



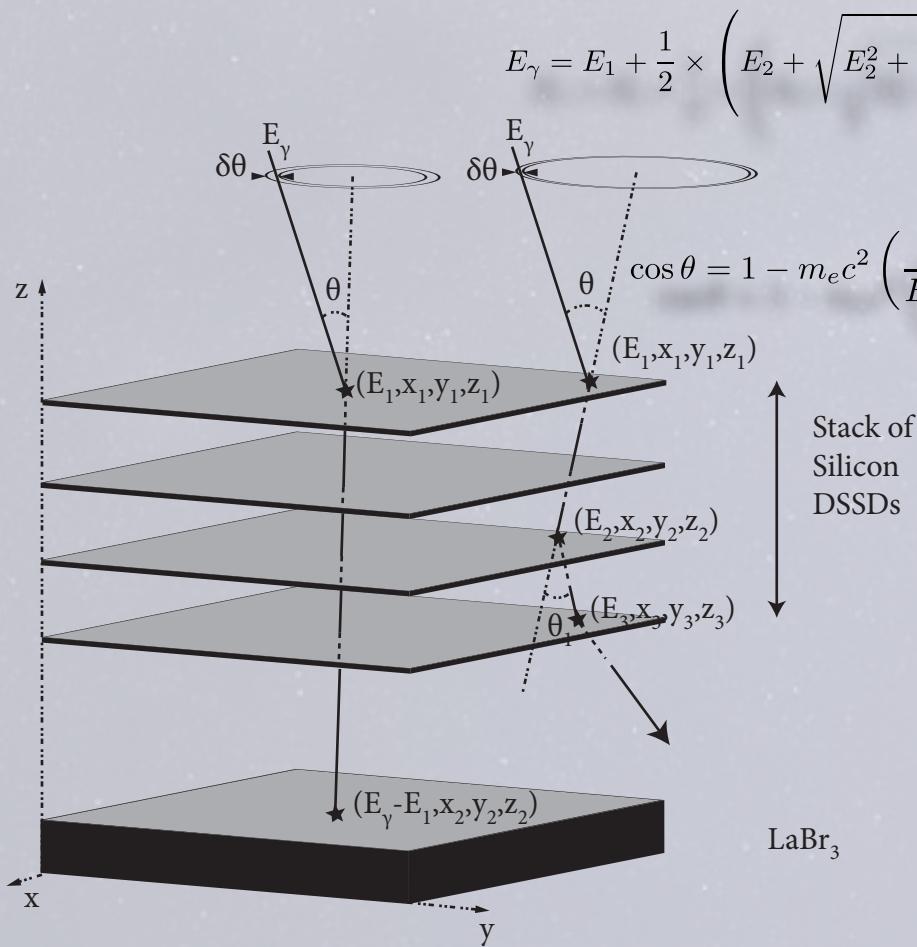
Instrumental issues for detection at nucleosynthesis energy range

R&D on a Compton telescope equipped
with Si-DSSDs

Philippe Laurent & Youri Dolgorouky

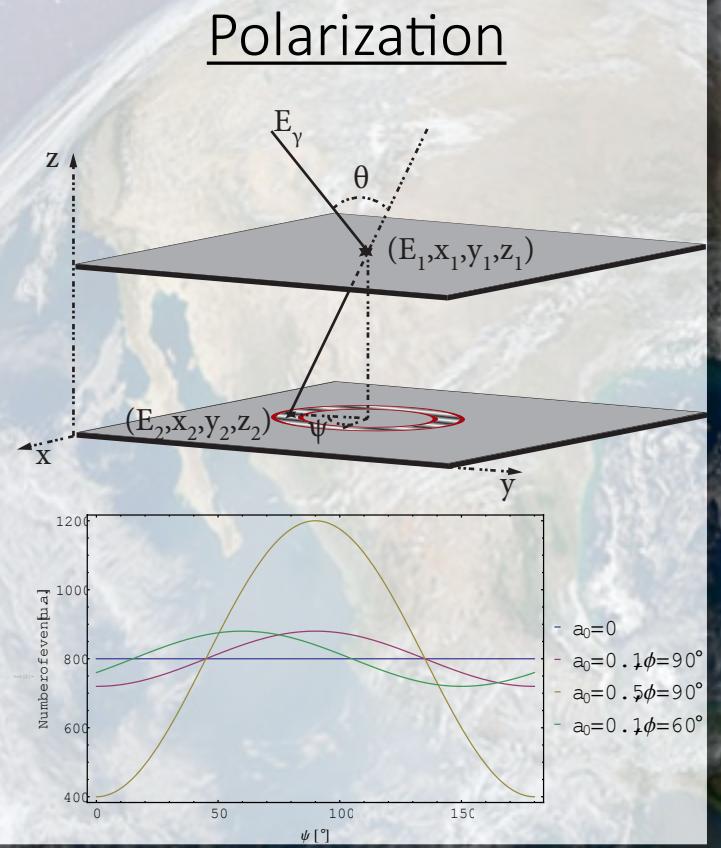
Compton Telescope

Energy and direction reconstruction of the cosmological sources



$$E_\gamma = E_1 + \frac{1}{2} \times \left(E_2 + \sqrt{E_2^2 + \frac{m_e c^2 E_2}{1 - \cos \theta_1}} \right)$$

$$\cos \theta = 1 - m_e c^2 \left(\frac{1}{E_\gamma - E_1} - \frac{1}{E_\gamma} \right)$$



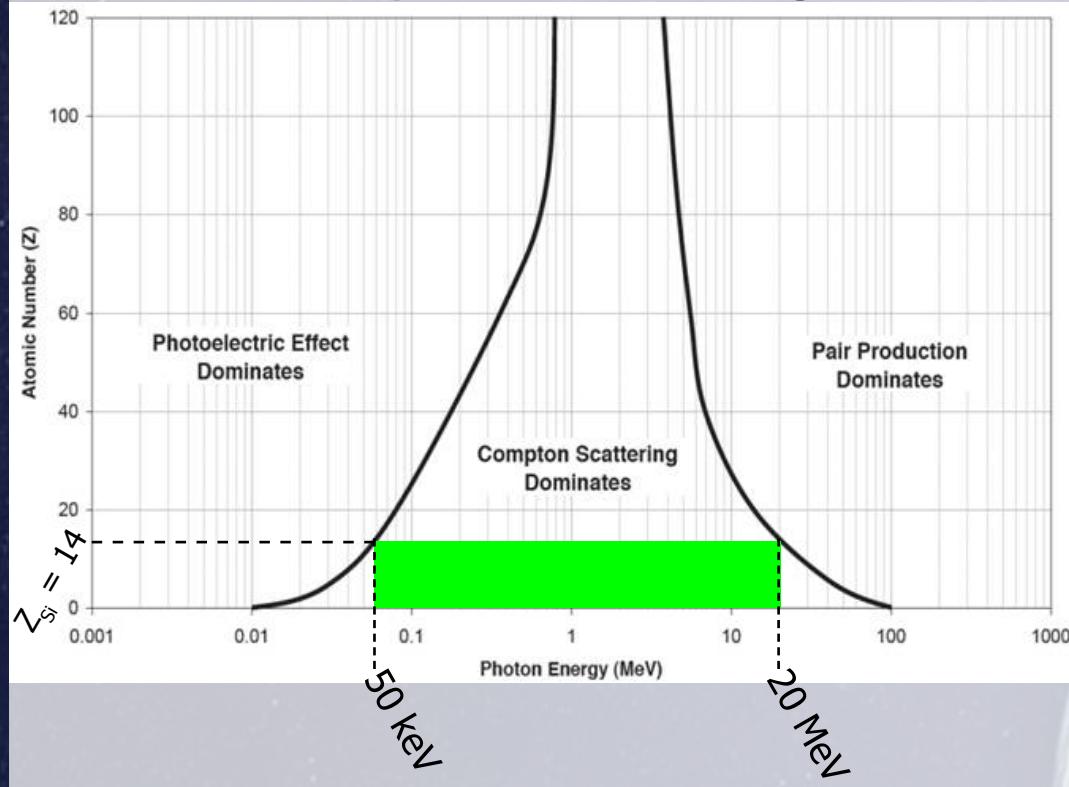
Main constrains

- Satellite :
 - **Mass budget**
 - **Electrical power budget**
 - **Technological Readiness Level**
- Environment :
 - **Radioactive background** at 1 MeV
- Compton telescope technic :
 - **Predominance of the Compton scattering** around the 1 MeV energy range



Silicon Double Sided Strip Detectors (Si-DSSD)

Compton Scattering



Stable element : low radioactivity

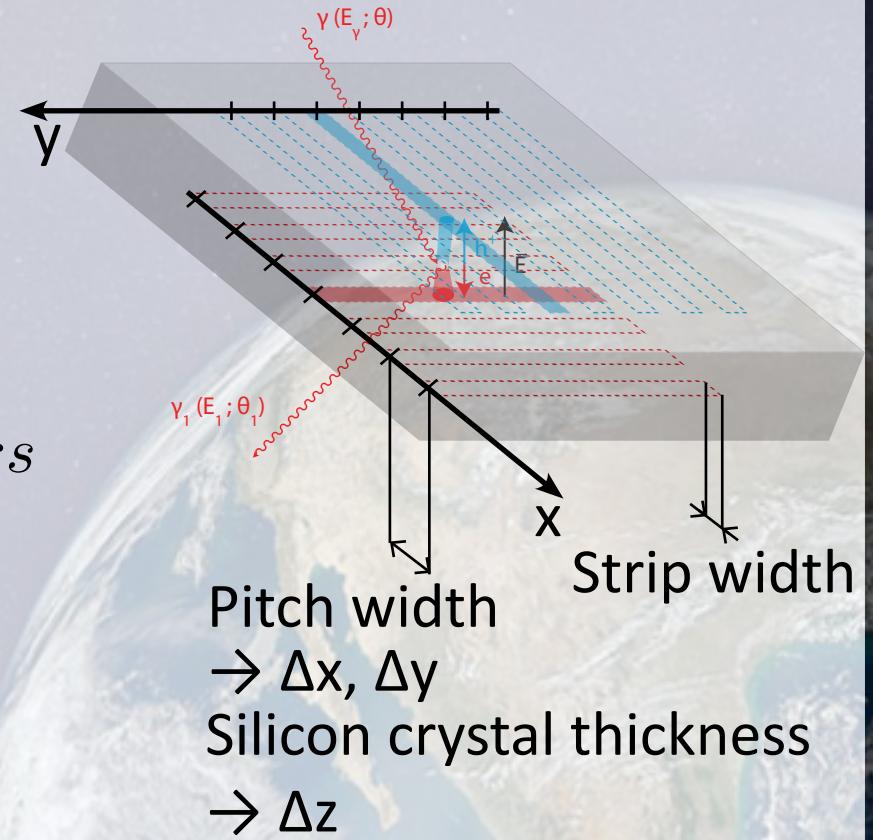
Light element : low mass

Silicon Double Sided Strip Detectors (Si-DSSD)

DSSD Principle

- Semiconductor material
 - PN Junction
 - Reverse Bias Voltage

$$E_{e^-} \propto \text{Number of hole/electron pairs}$$



Compton scattering position and energy measurement

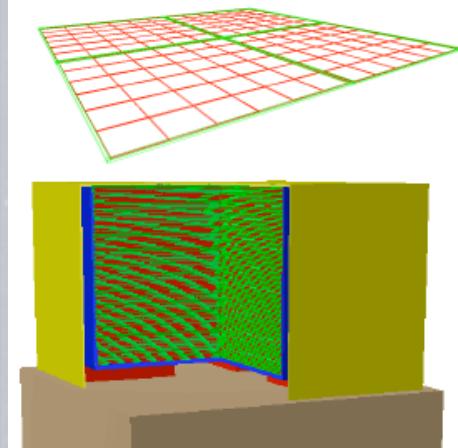


Energy and angular resolution

- Energy resolution } Compton telescope \Rightarrow DSSD
- Angular resolution } performances Energy resolution

MEGALIB simulation

CAPSiT configuration :



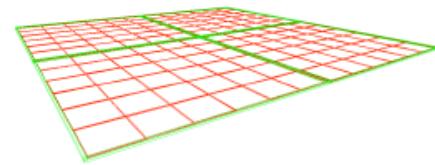
80 layers of Si-DSSD
 12×12 DSSDs
DSSD :
10 cm \times 10 cm
2 mm thickness
Pitch width : 1.5 mm



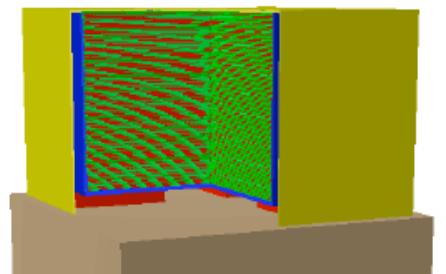
Energy and angular resolution

MEGALIB simulation

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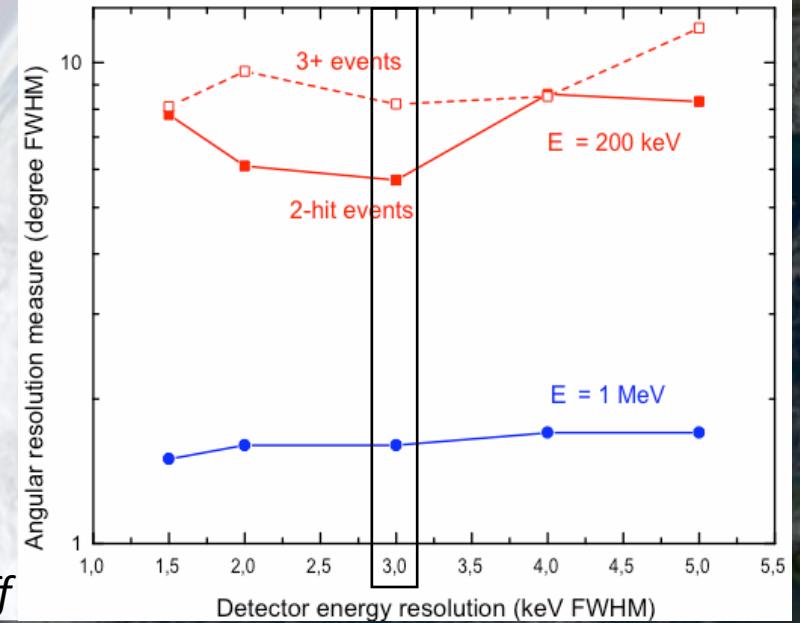
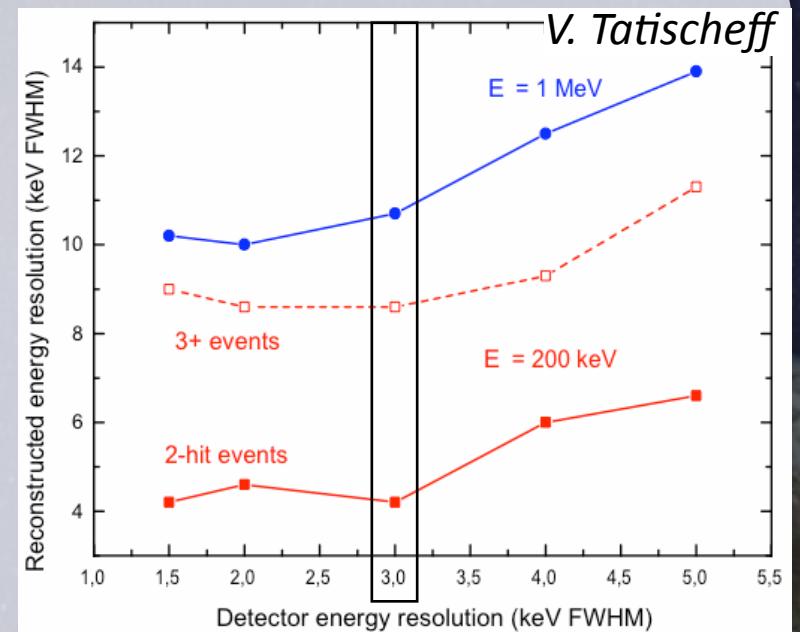
80 layers of Si-DSSD
 12×12 DSSDs



DSSD :
10 cm \times 10 cm
2 mm thickness
Pitch width : 1.5 mm

Energy resolution : 3keV / DSSD

V. Tatischeff

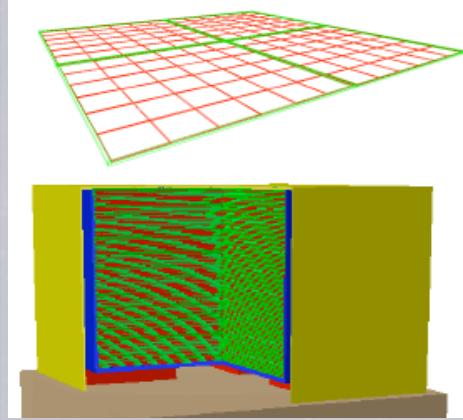




Energy and angular resolution

MEGALIB simulation

CAPSiTTe configuration :



80 layers of Si-DSSD

12×12 DSSDs

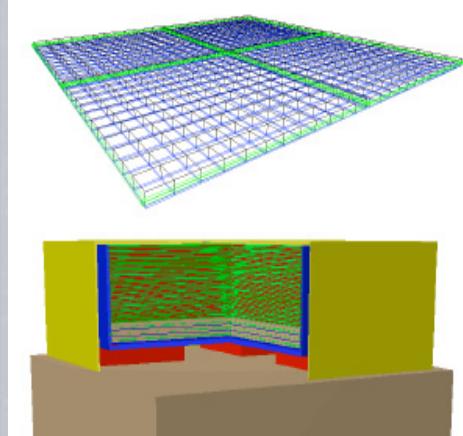
DSSD :

$10 \text{ cm} \times 10 \text{ cm}$

2 mm thickness

Pitch width : 1.5 mm

CoCoTe configuration :



30 layers of Si-DSSD

12×12 DSSDs

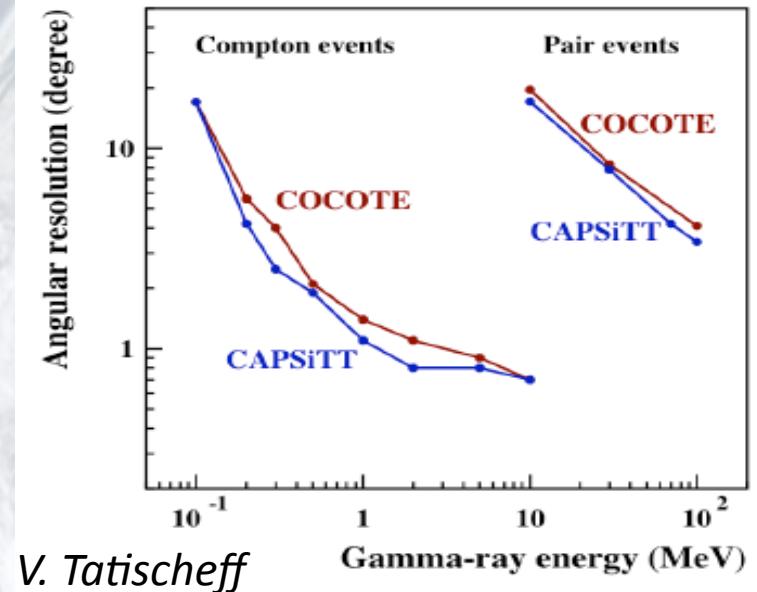
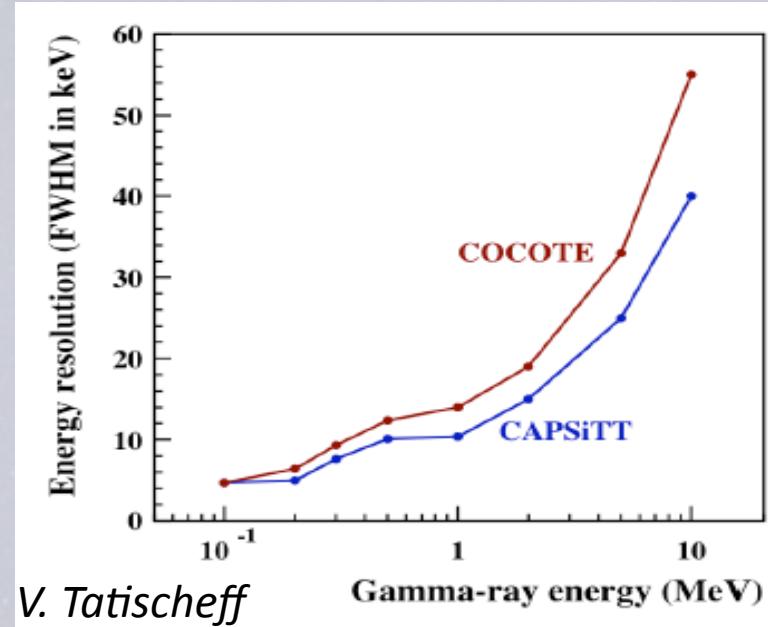
LaBr_3 calorimeter

DSSD :

$10 \text{ cm} \times 10 \text{ cm}$

1.5 mm thickness

Pitch width : 1.5 mm



Energy resolution and Equivalent Noise Charge

$$\Delta E^2 = 2.35^2 \varepsilon_w^2 \cdot ENC_{tot}^2 + 2.35^2 F \varepsilon_w \cdot E_\gamma + \Delta E_{perme}^2$$

$$ENC_{tot}^2 = \frac{\alpha_s}{\tau_s} \cdot (C_{DSSD} + C_p + C_e)^2 + \alpha_f \cdot (C_{DSSD} + C_p + C_e)^2 + \alpha_p \tau_s \cdot I_{Leak.}$$

DSSD capacitance

DSSD Leakage current

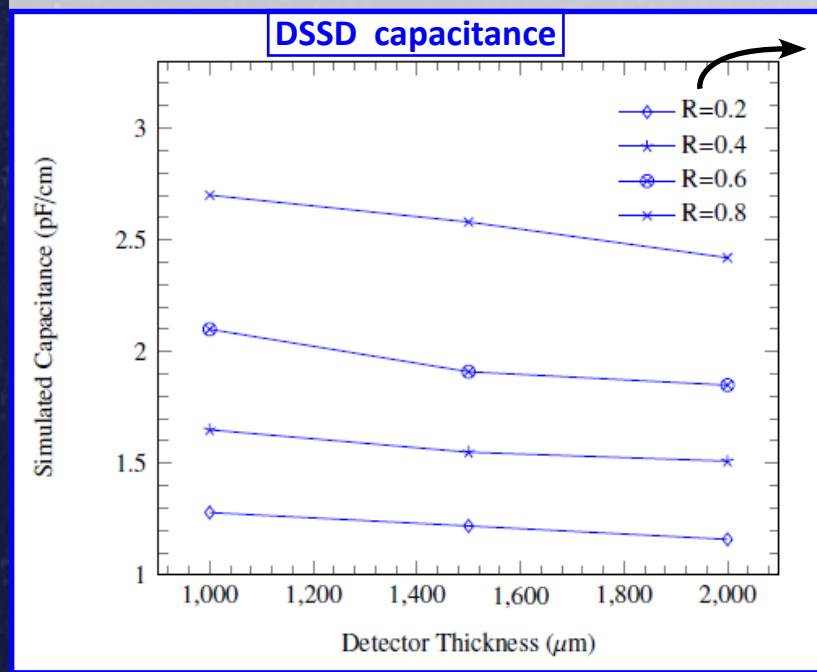
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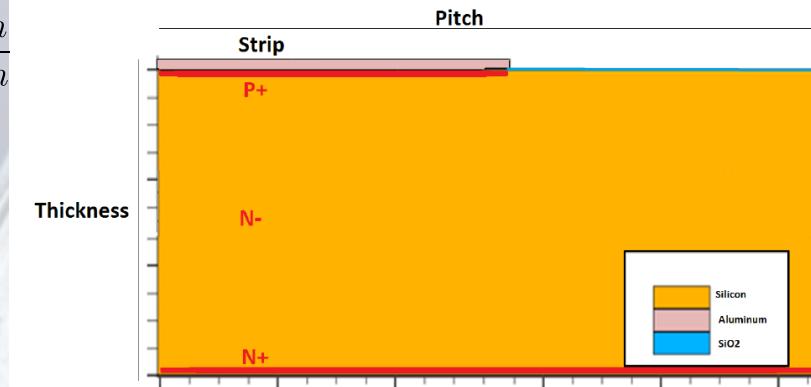
$$ENC_{tot}^2 = \frac{\alpha_s}{\tau_s} \cdot (C_{DSSD} + C_p + C_e)^2 + \alpha_f \cdot (C_{DSSD} + C_p + C_e)^2 + \alpha_p \tau_s \cdot I_{Leak.}$$

DSSD capacitance

TCAD Simulation



$$\text{Ratio} = \frac{\text{strip width}}{\text{pitch width}}$$



M. Khalil et al. (2013)

Energy resolution and Equivalent Noise Charge

$$\Delta E^2 = 2.35^2 \varepsilon_w^2 \cdot ENC_{tot}^2 + 2.35^2 F \varepsilon_w \cdot E_\gamma + \Delta E_{perme}^2$$

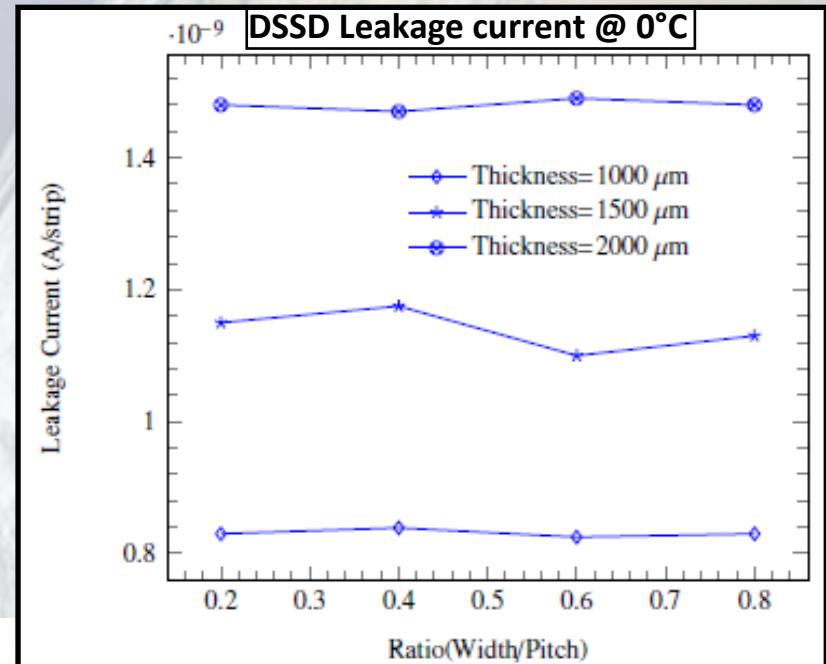
$$ENC_{tot}^2 = \frac{\alpha_s}{\tau_s} \cdot (C_{DSSD} + C_p + C_e)^2 + \alpha_f \cdot (C_{DSSD} + C_p + C_e)^2 + \alpha_p \tau_s \cdot I_{Leak.}$$

DSSD Leakage current

TCAD Simulation

Good approximation :

$$\frac{I_{Leak}(T_2)}{I_{Leak}(T_1)} = 2^{\frac{T_2 - T_1}{7}}$$



M. Khalil et al. (2013)

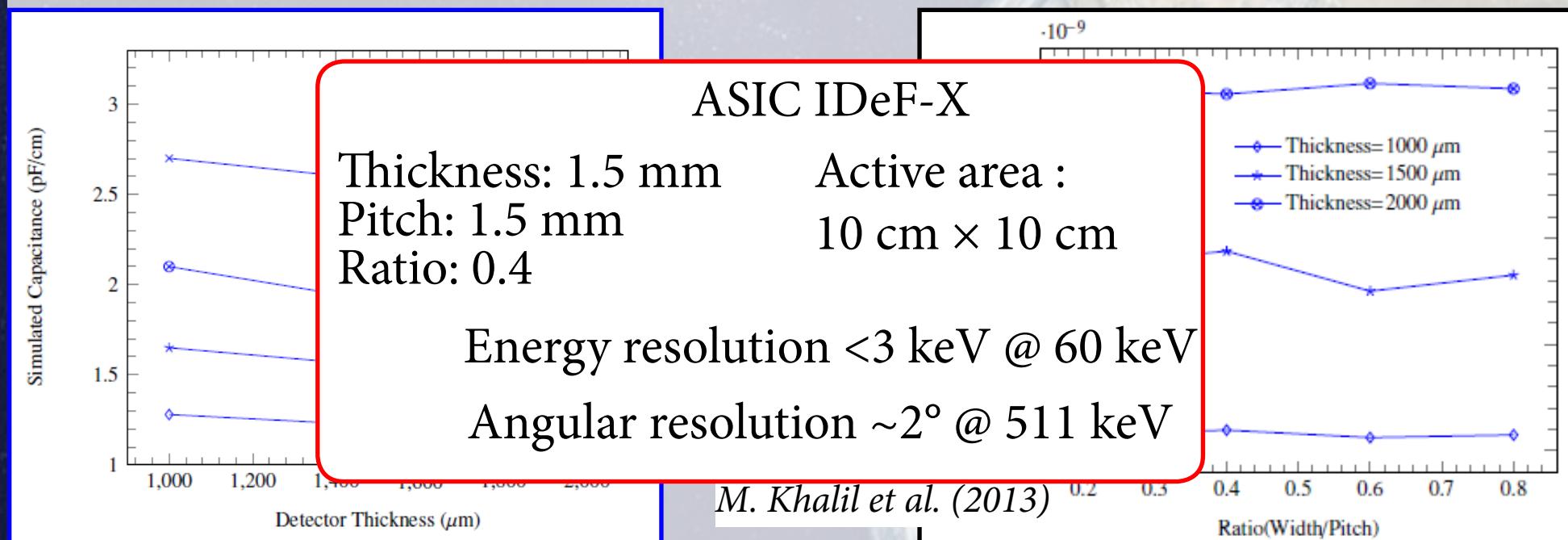
Energy resolution and equivalent noise charge

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↑
DSSD capacitance
↑
DSSD Leakage current

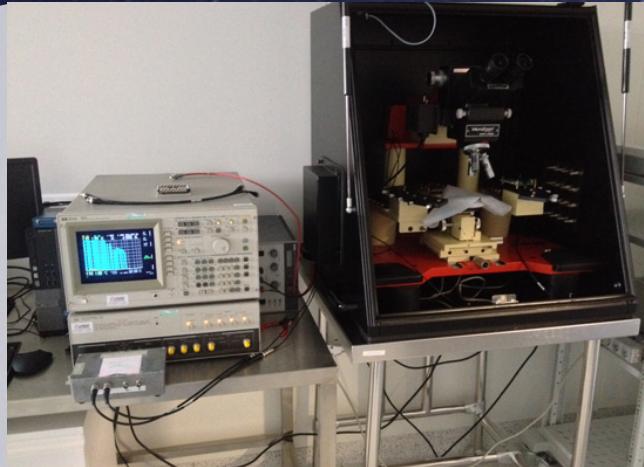
TCAD Simulation





Instrumental issues for detection at nucleosynthesis energy range

R&D on a Compton telescope with Si-DSSDs

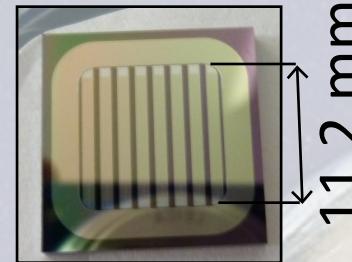


Electrical characterizations

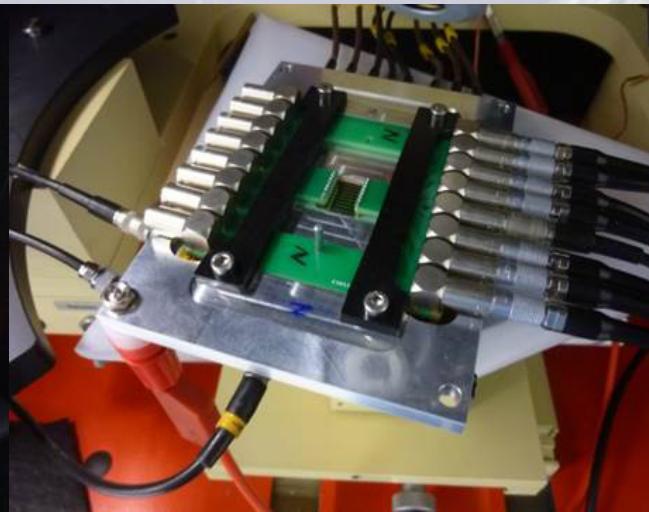
Experimental set-up at APC Laboratory

Baby DSSD (SINTEF)

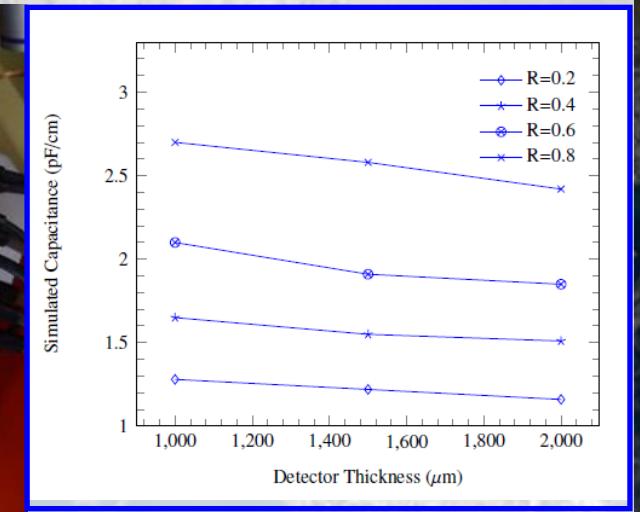
Thickness : 2 mm



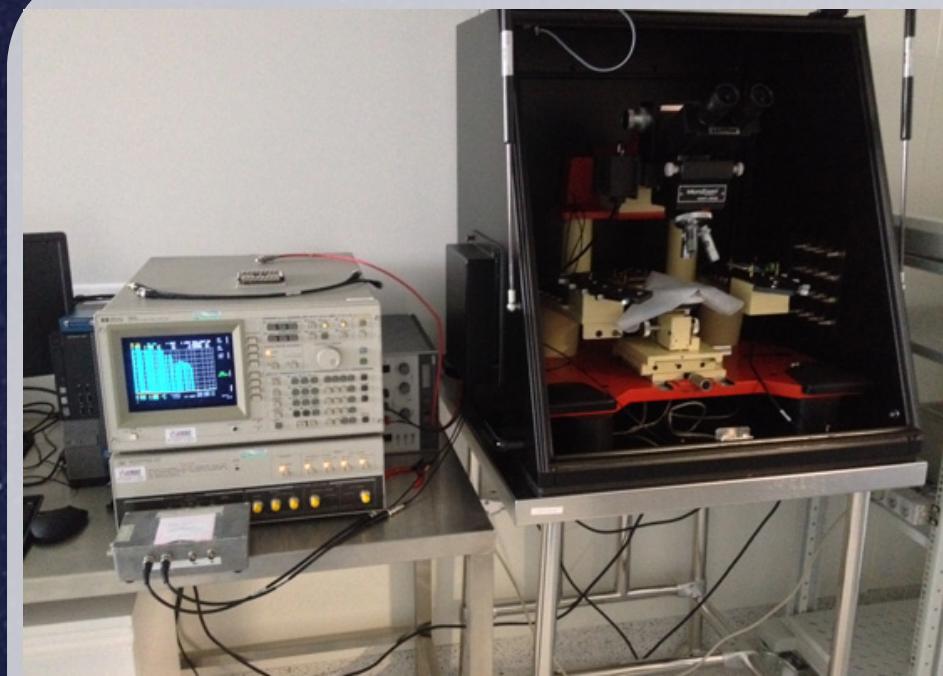
C vs V curve



$C_{DSSD} = 3.1 \text{ pF/cm}$ @ 550 V and 1MHz
Ratio = 0.93



Simulation

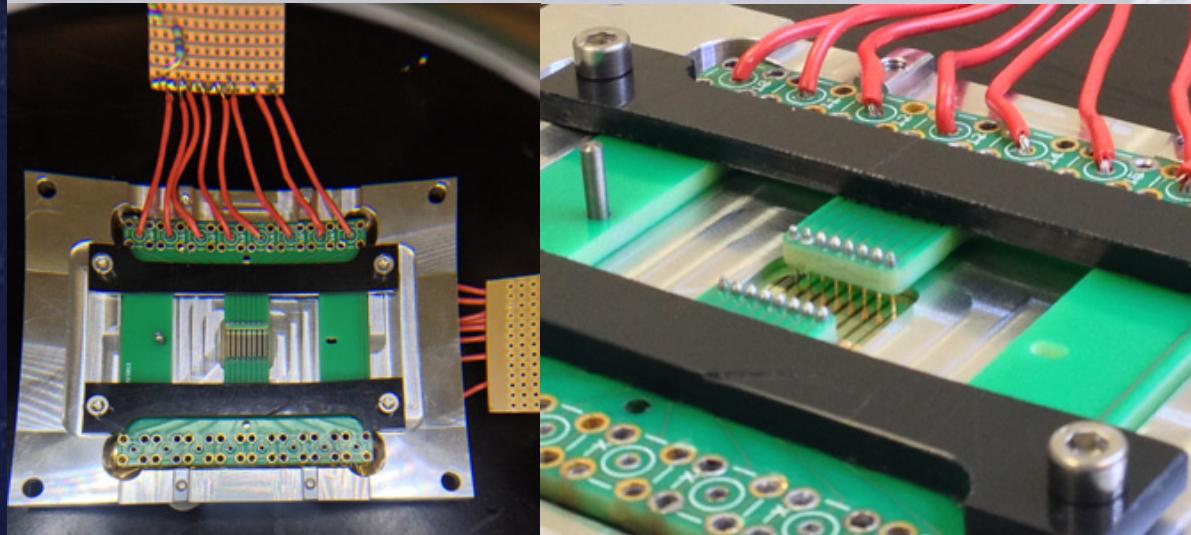
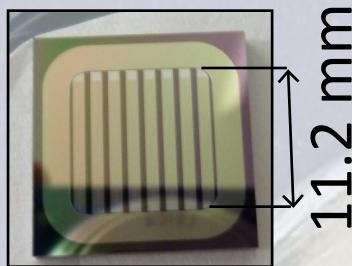


Electrical characterizations

Experimental set-up at APC Laboratory

Baby DSSD (SINTEF)

Thickness : 2 mm

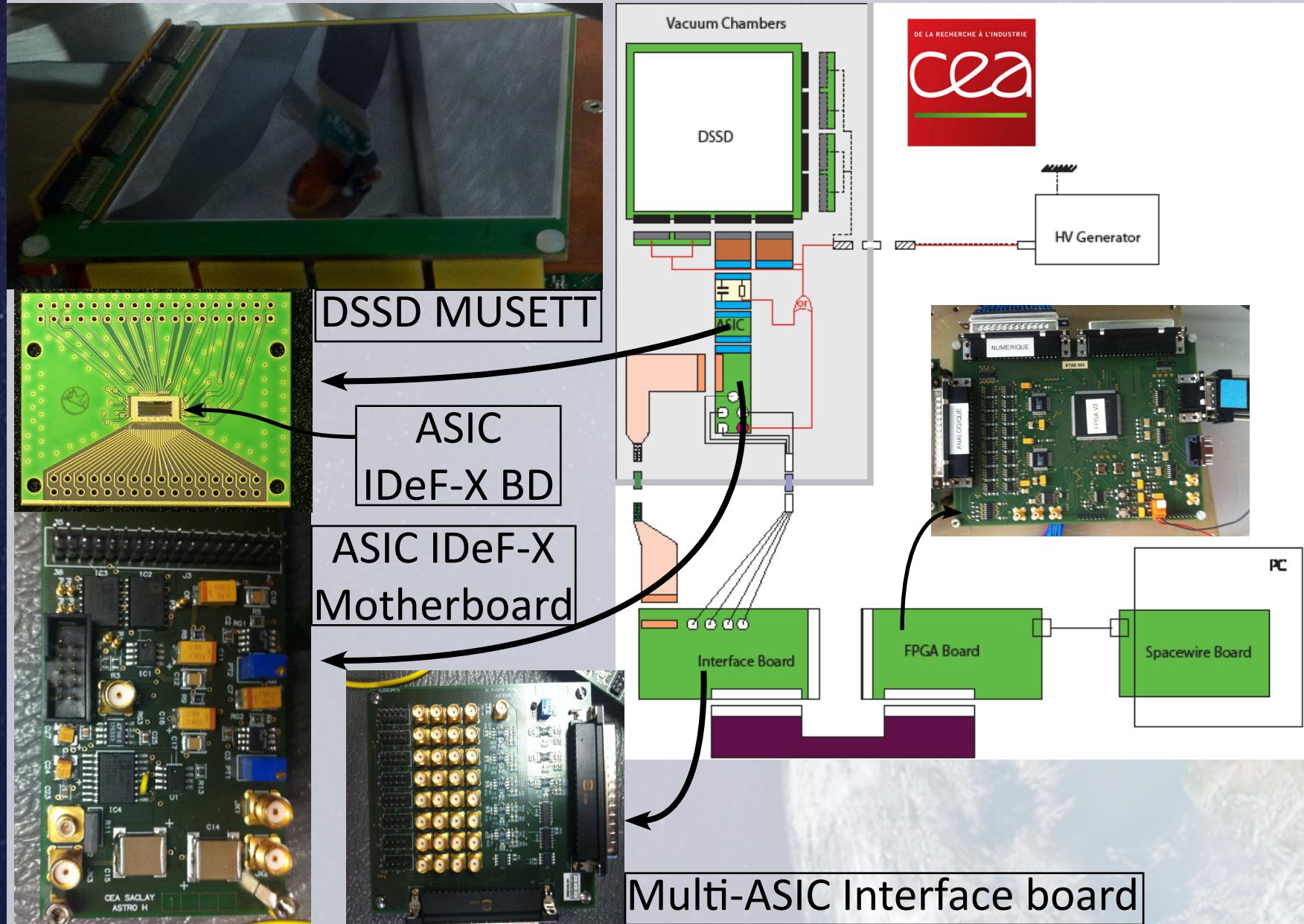


Test :
Forward or reverse bias ?
Reverse bias :
 $< 10 \text{ nA} @ 20^\circ\text{C and } 200 \text{ V}$

I vs V curve

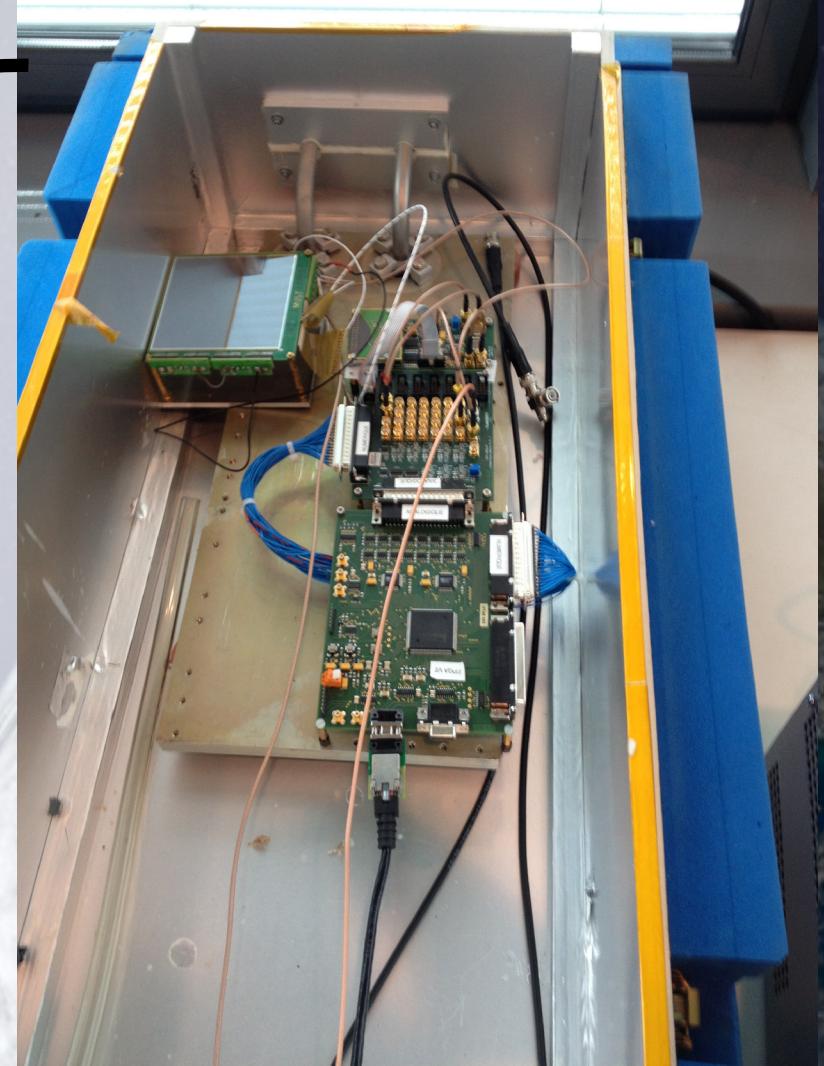


Spectrometric test bench





Spectrometric test bench



Installation at APC Laboratory



Spectrometric test bench On going activities

Equip the box :

- Nitrogen flushing → ↓ dew point
- Cool down device → -20°C → $I_{\text{Leak}}/64$

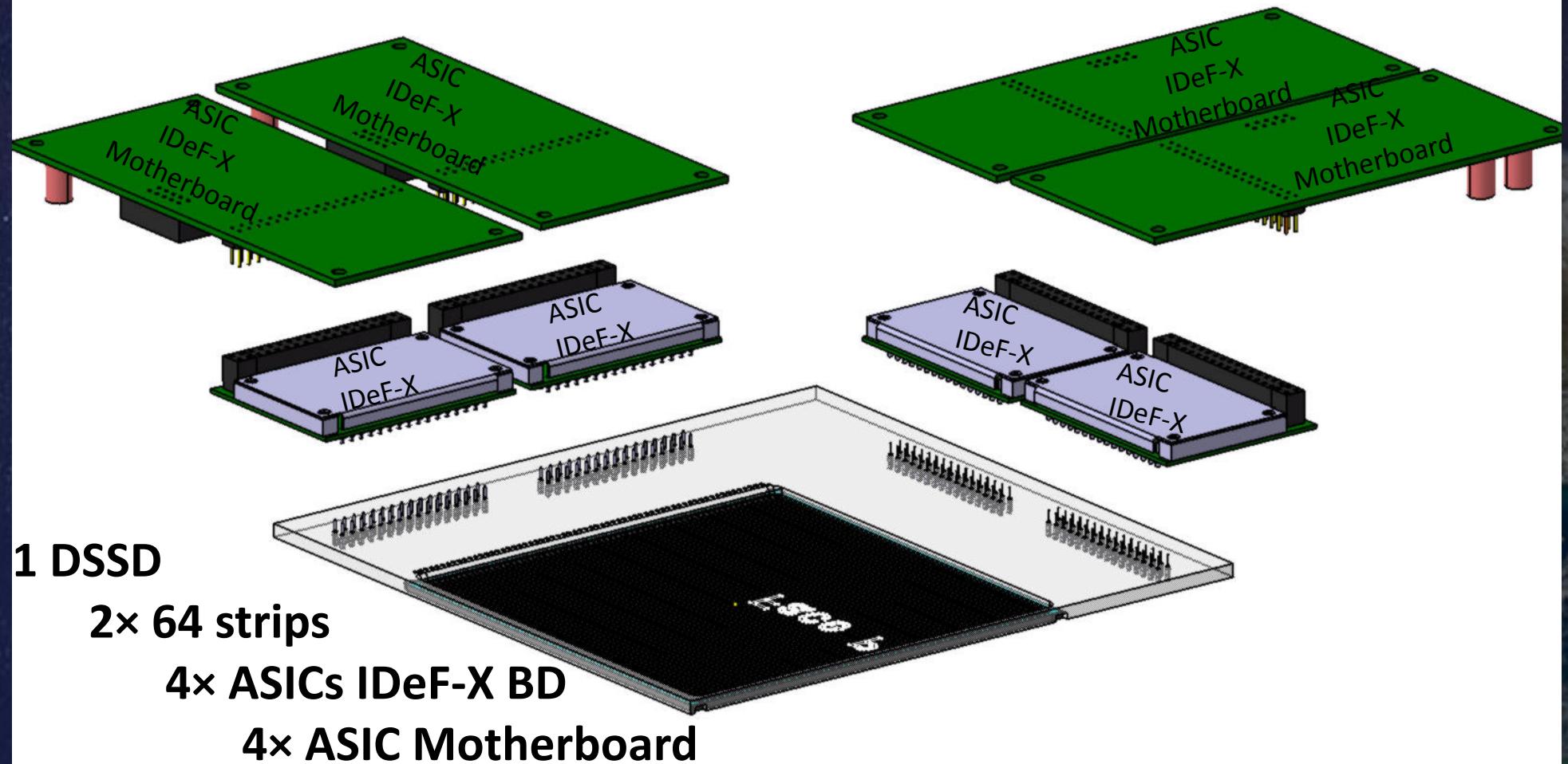
Electromagnetic compatibility

Flashing issues

**Test with Musett and the 60 keV ray of ^{241}Am source
with 2 ASICs IDeF-X BD**

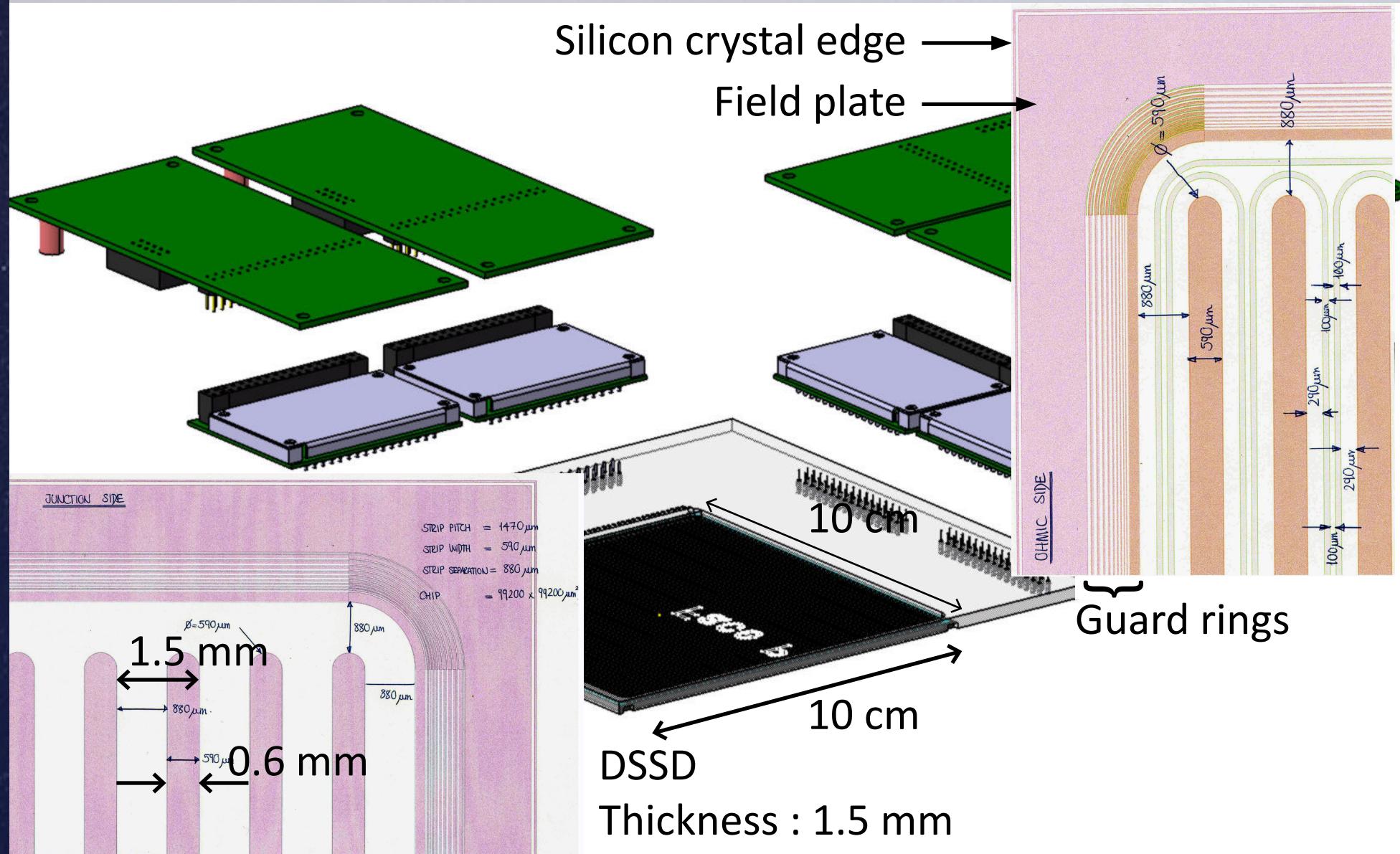


DSSDs & PCB - design



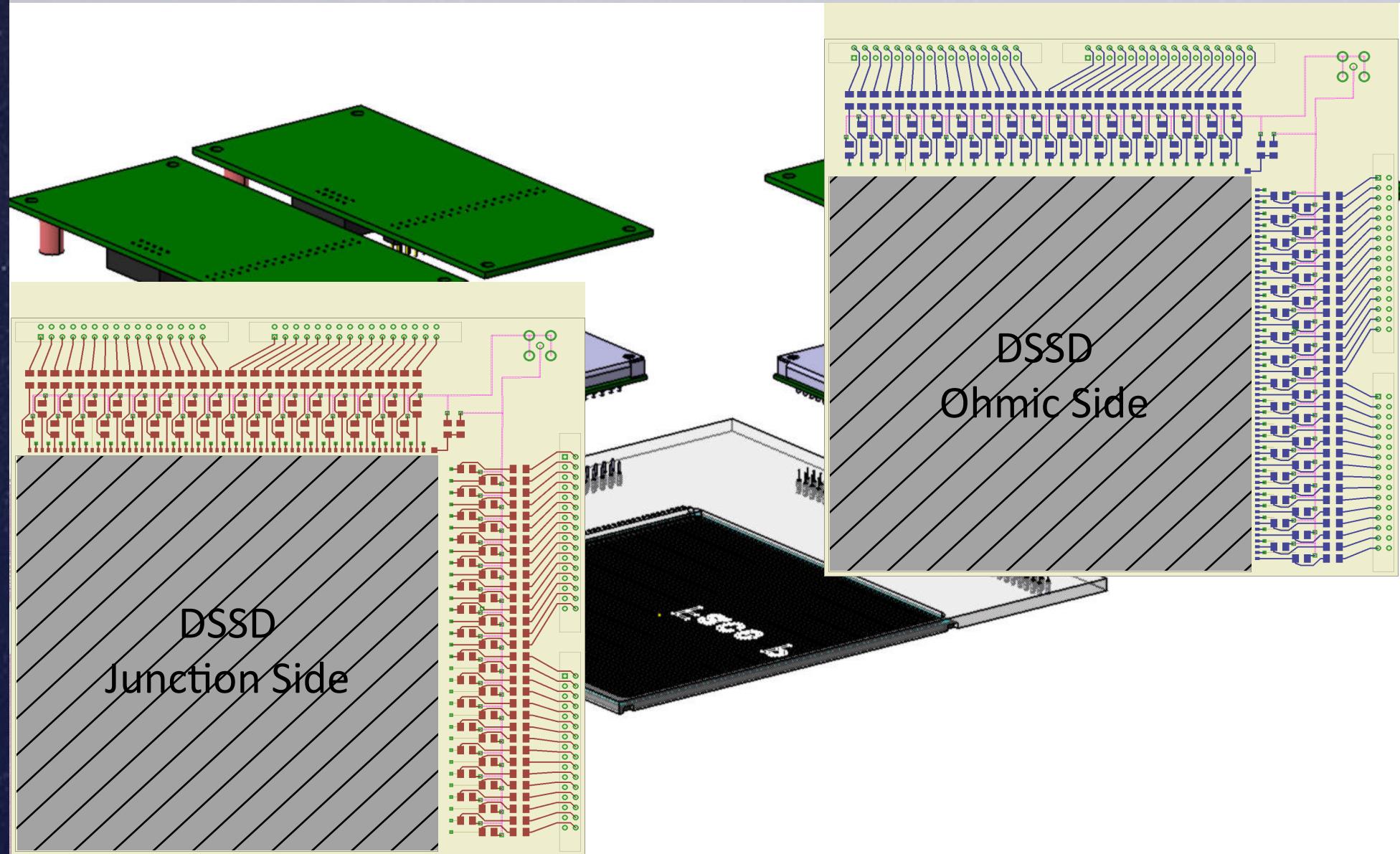


DSSDs & PCB - design

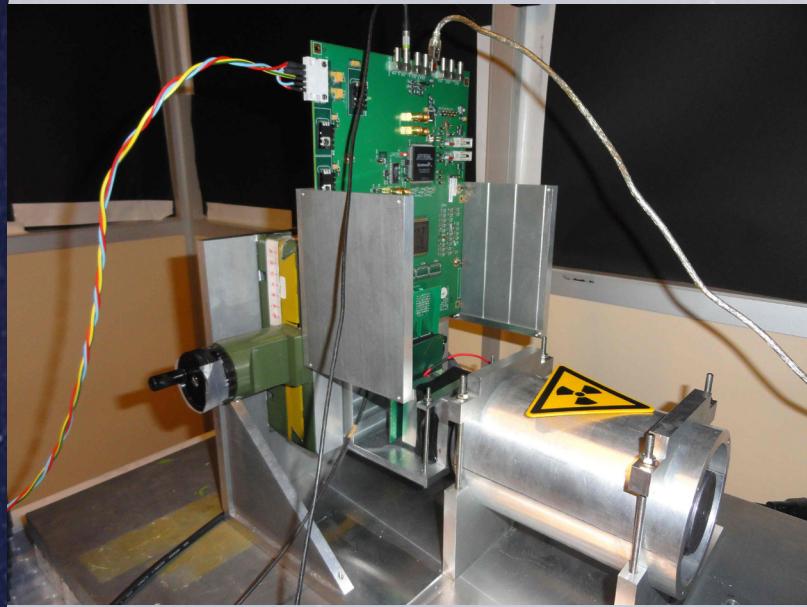




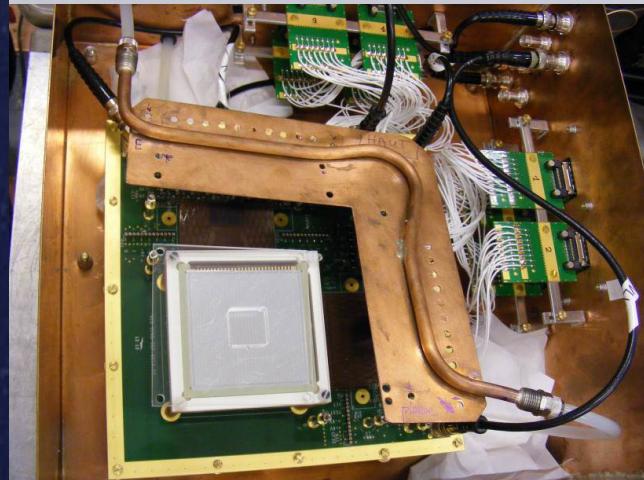
DSSD & PCB - design



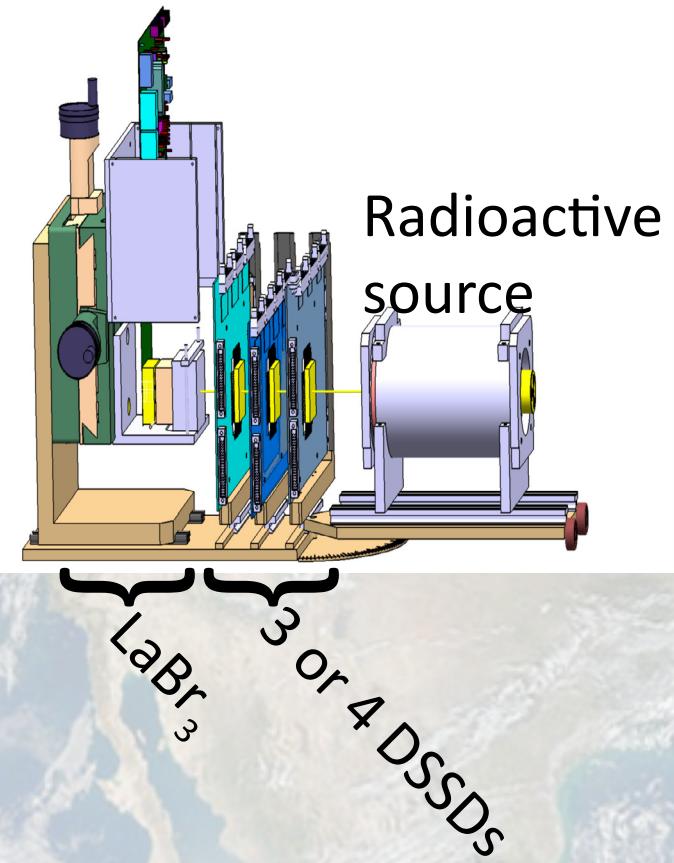
CoCoTe : Compact Compton Telescope



Existing experimental set-up at CSNSM for LaBr_3 tests



Experimental setup developed at IPNO



3 or 4 DSSDs :
■ 2 from IPNO
■ 2 from APC Lab.



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R&D on a Compton telescope with Si-DSSDs



THANK YOU