



Origin of Dune Orientation on Titan

New Insights from experimental & field measurements

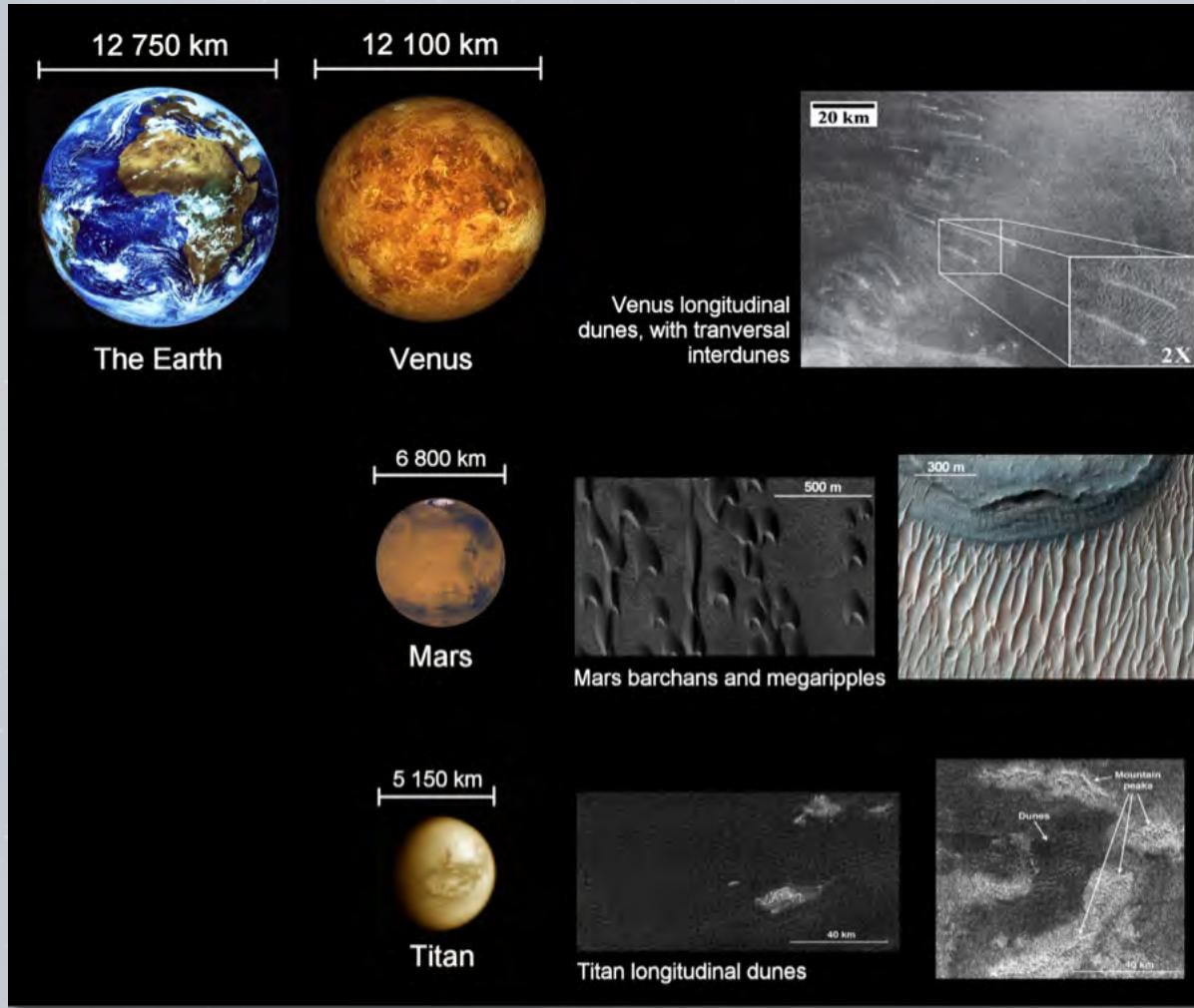
A. Lucas, S. Rodriguez, C. Narteau, B. Charnay,
A. Garcia, S. Courrech du Pont, A.G. Hayes





Sand Seas, a common landscape in the Solar System

We need sediment and atmosphere



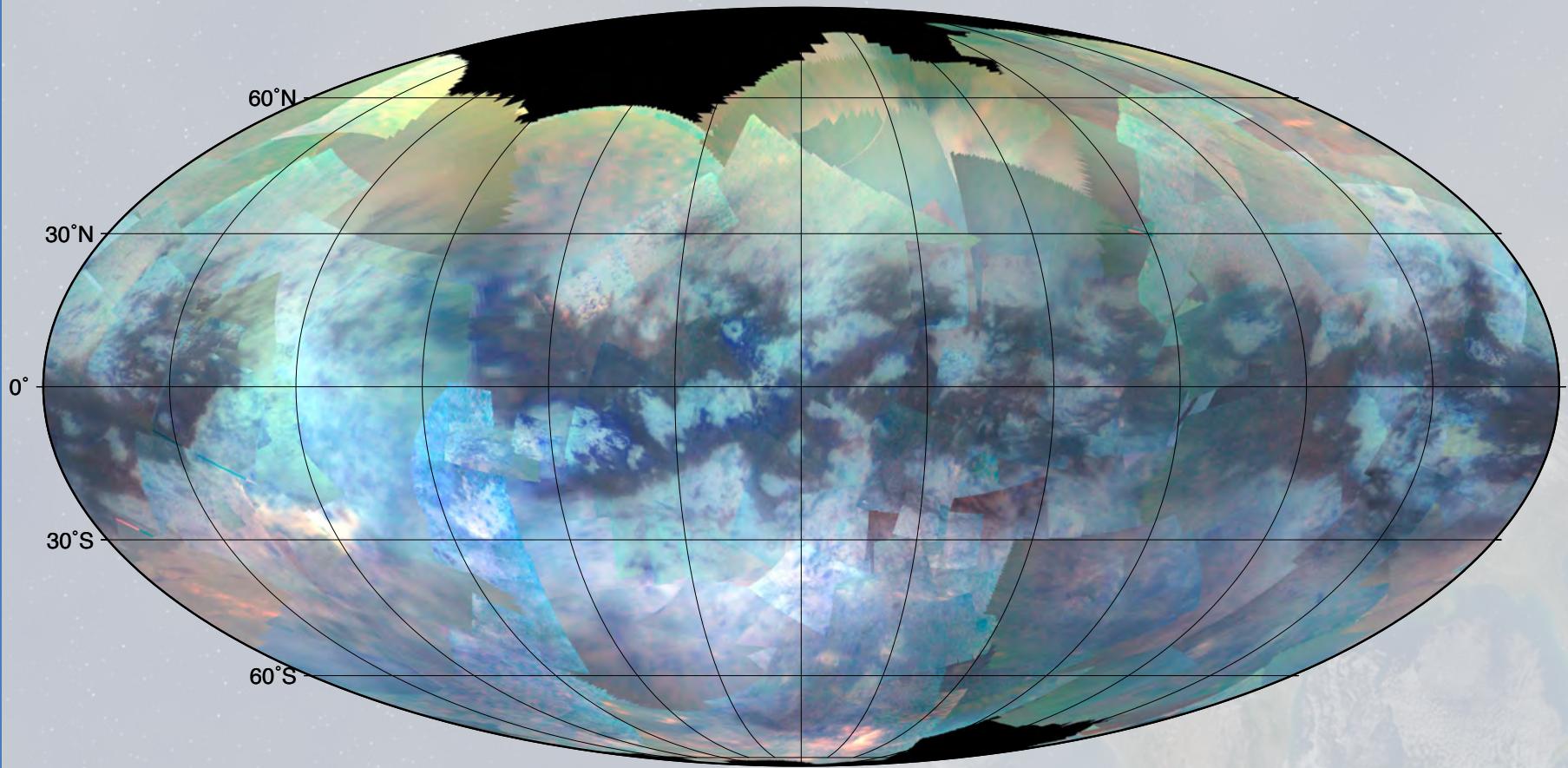


Origin of Dune Orientation on Titan

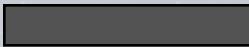




Titan seen by VIMS/Cassini

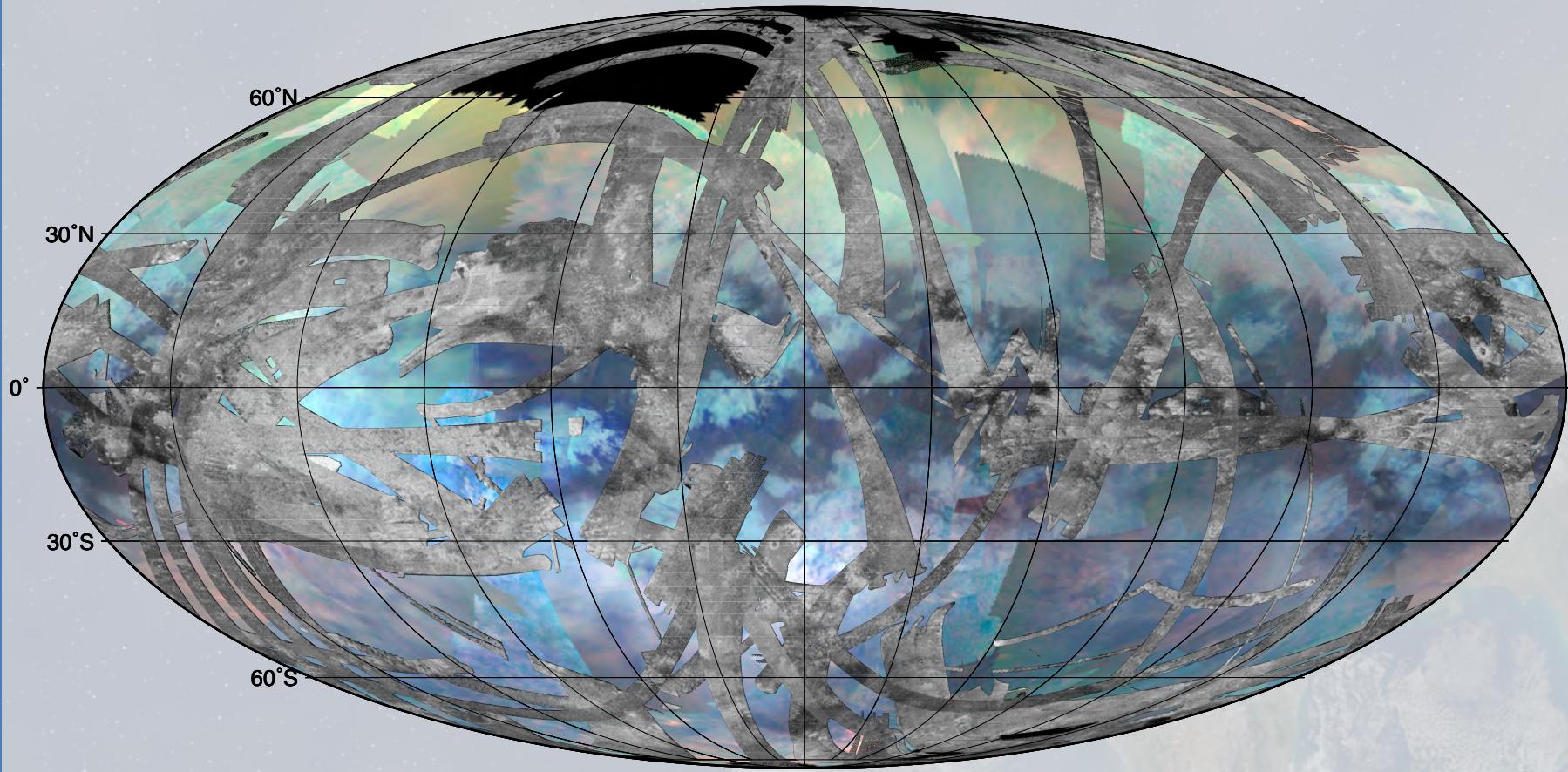


2 500 km (eq.)



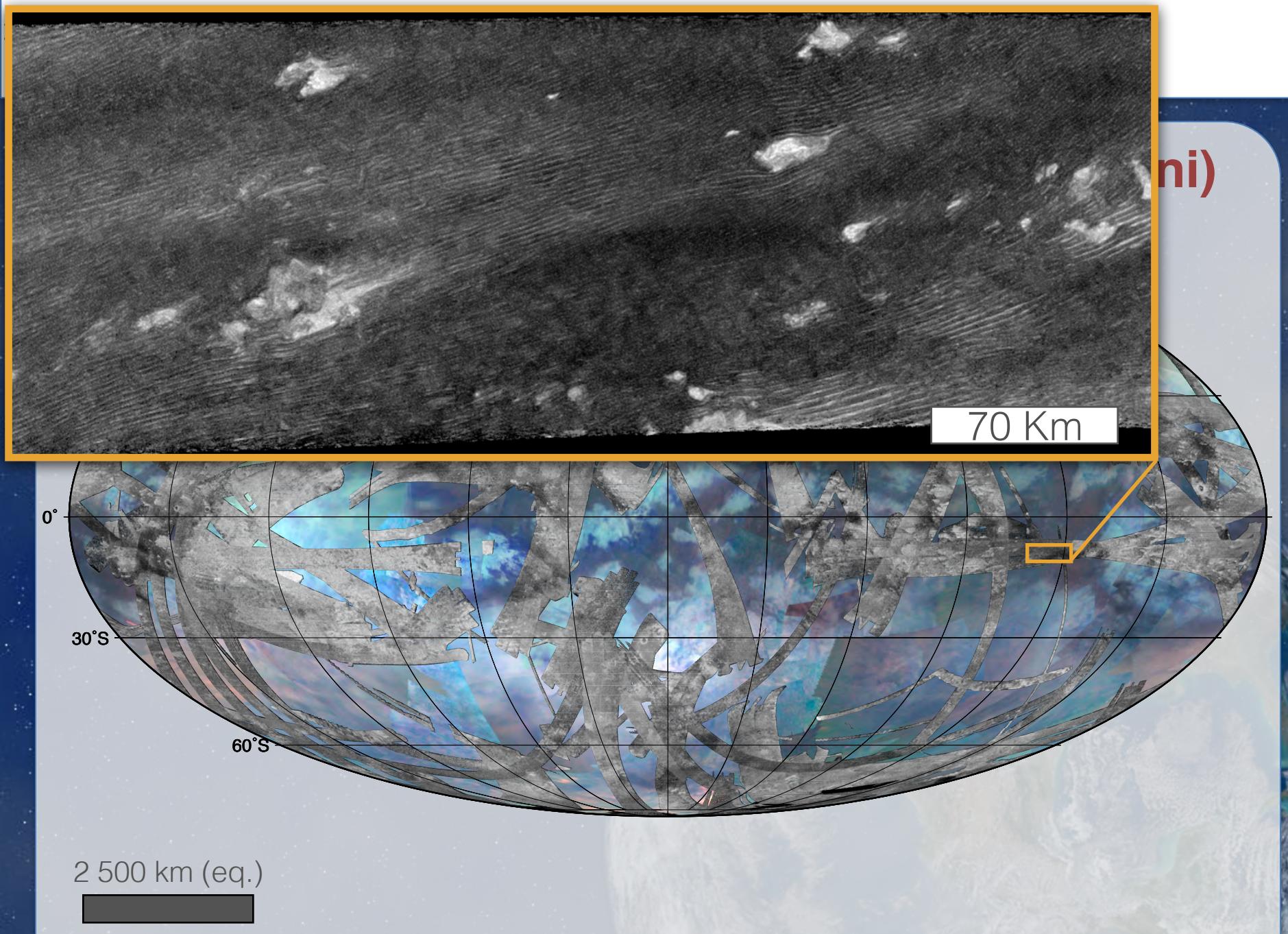


Titan seen by RADAR in SAR Mode (Cassini)

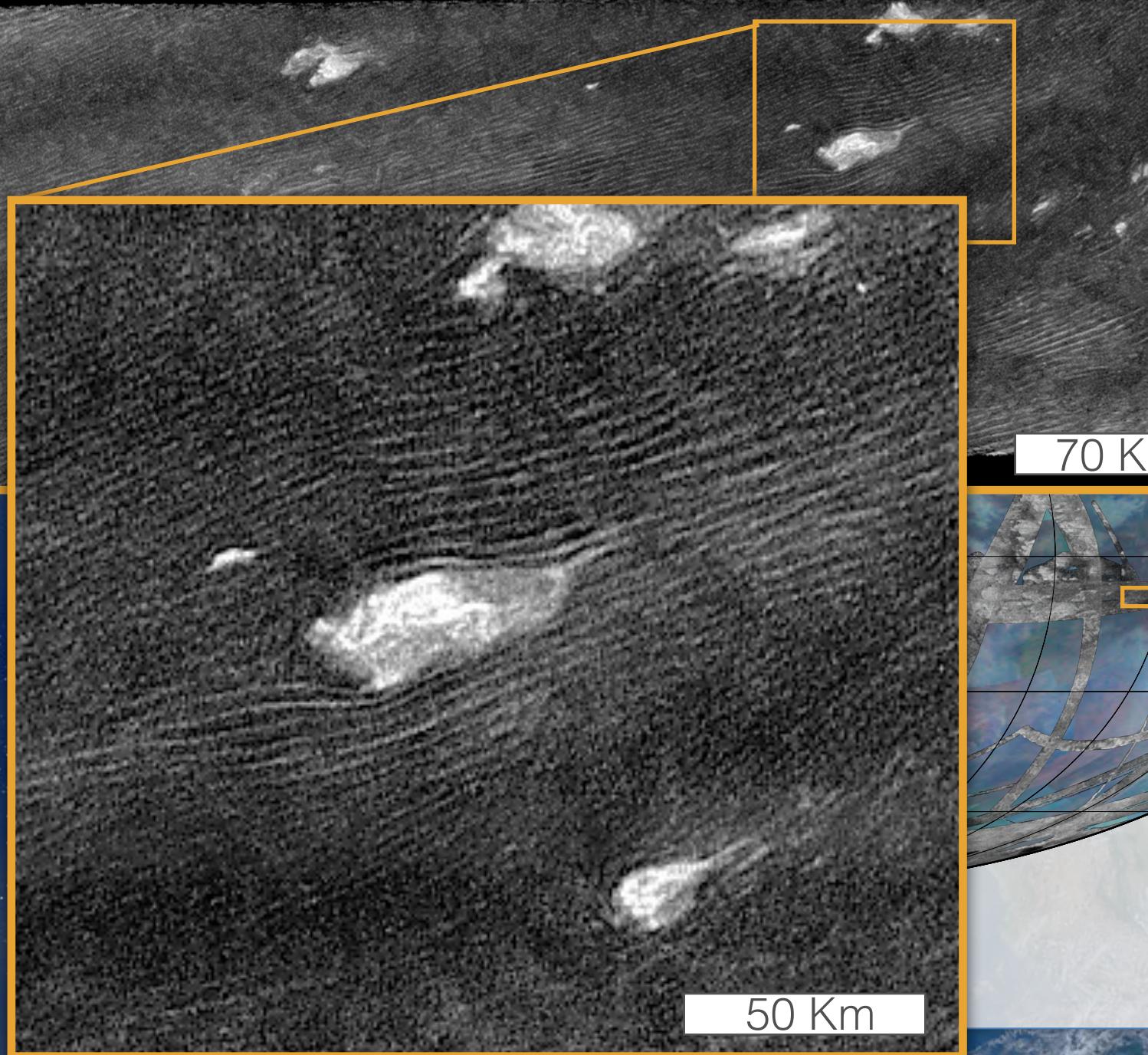


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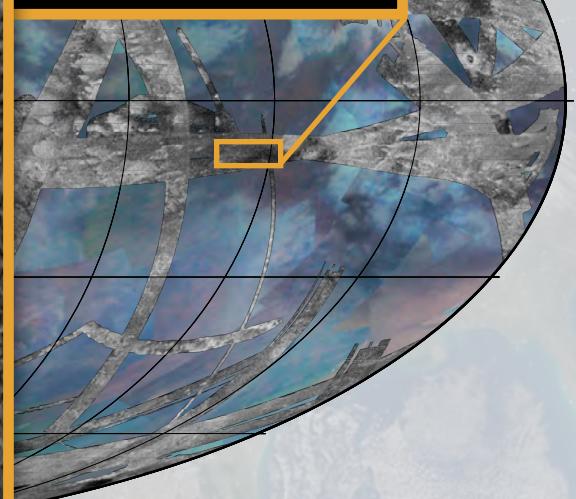
ni)



ni)



70 Km

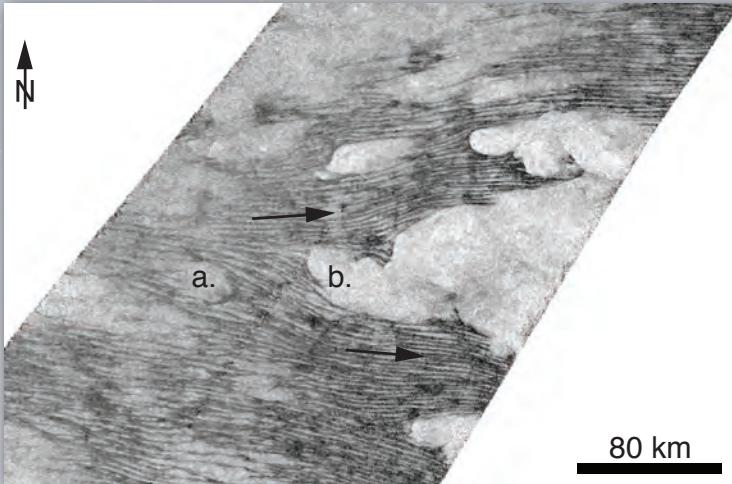


May 16th 2014

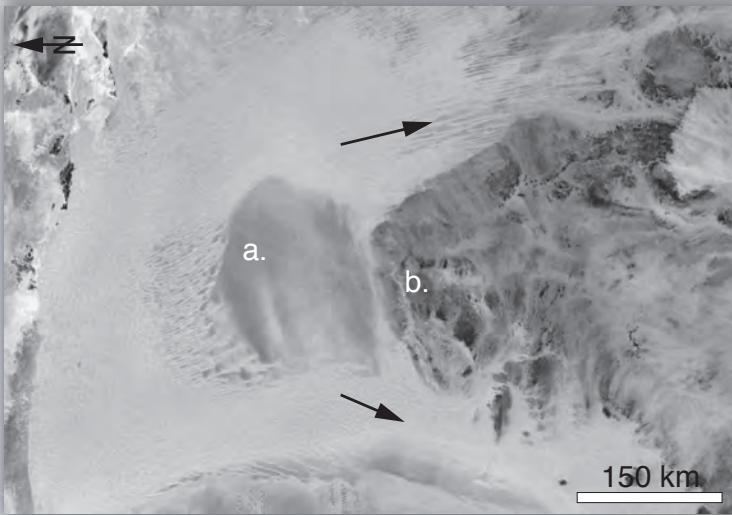


Eastward dune propagation

Titan



Egypt





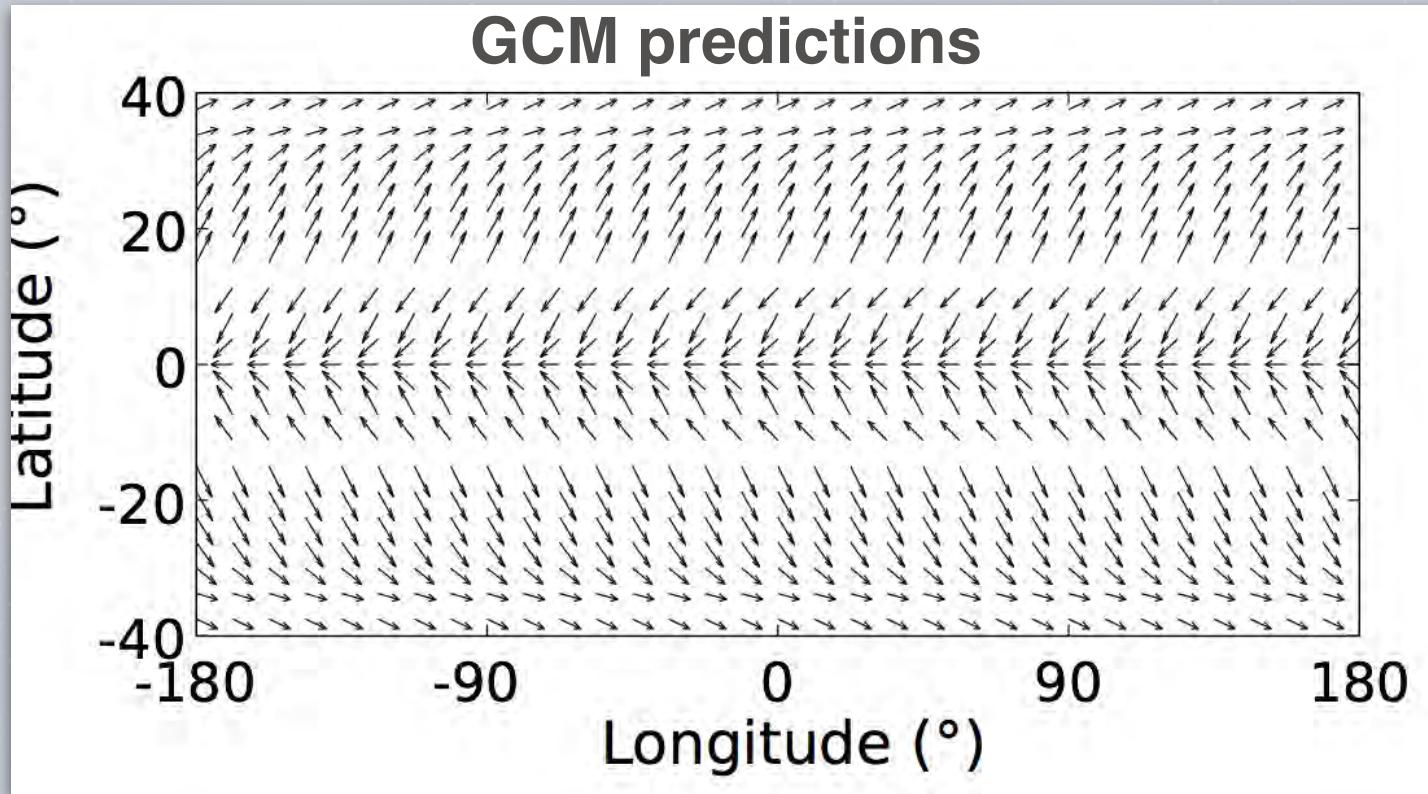
Titan's Dunes properties

Global direction	Type	Length	Wavelength	Height	Location
Eastward	Linear	100-1000' km	~3 km	~ 100' m	equatorial band (+/- 30°)

Lorenz et al., 2006; 2008; Radebaugh et al. 2008; 2010; 2013; Rubin & Hesp, 2009; Neish et al. 2010; Le Gall et al., 2011; 2012; Arnold et al. 2013; Rodriguez et al. 2013



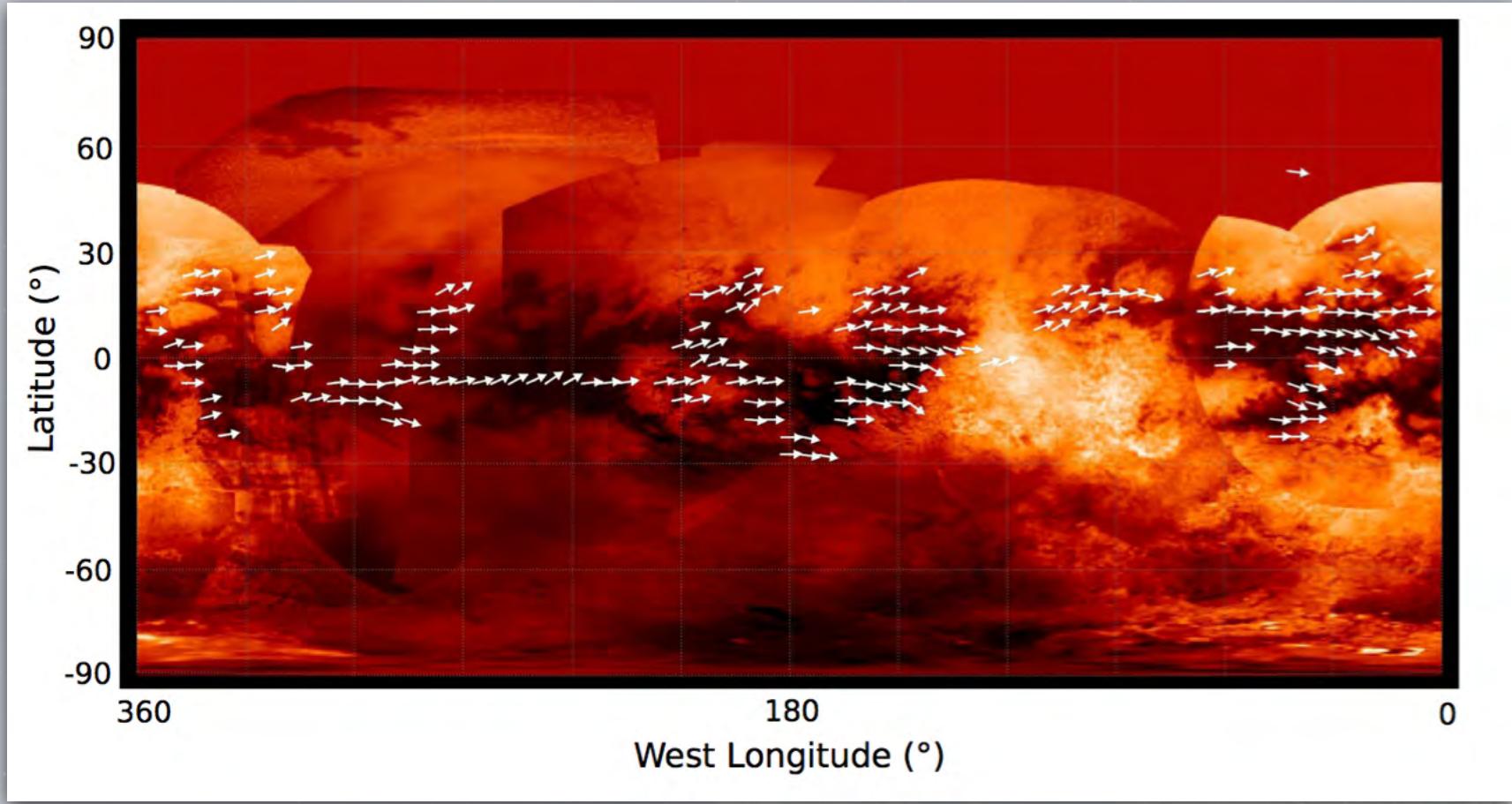
Global Circulation Model of Titan



*Charnay et al. AGU 2013, &
Charnay et al (incl. Lucas) in review for Nat. Geosc.*



Dunes conundrum



Propagation direction after Lorenz *et al.* (2006) and following works



Fondamental questions

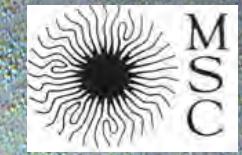
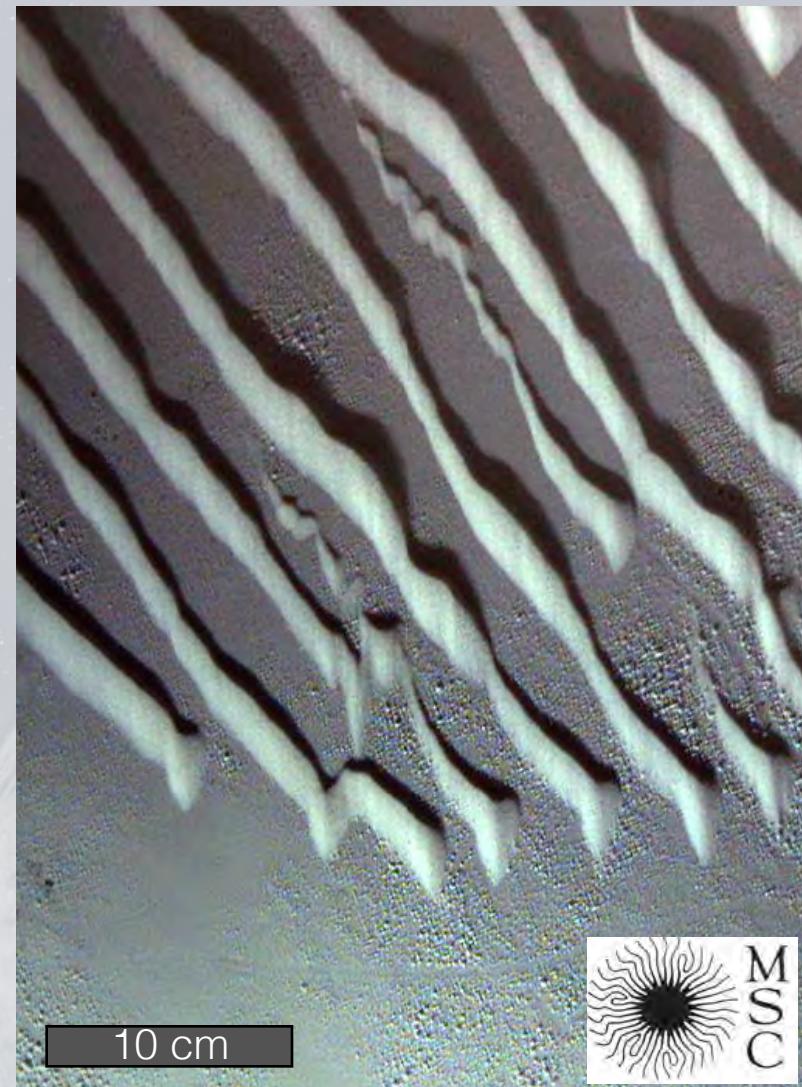
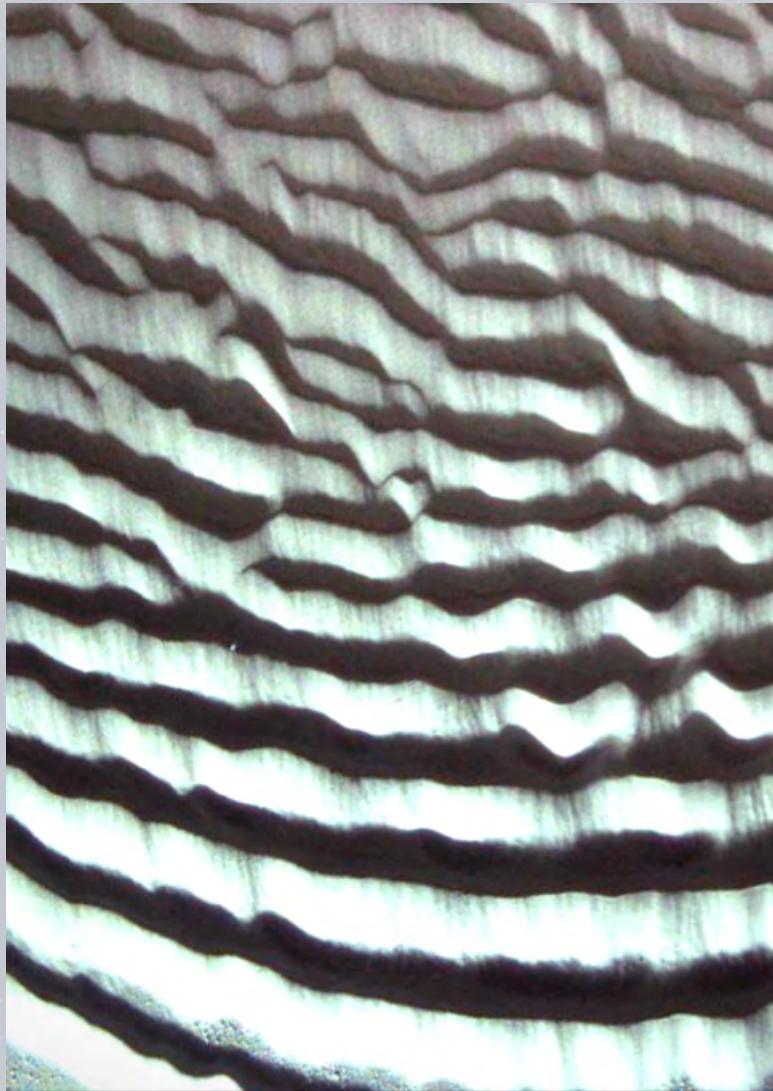
- How to explain their geographic confinement?
- How to explain their linear shape and orientation?
- How to explain their direction of propagation?
- Are they still active?



To answer these questions, we take advantage of new experimental, numerical and analytical results

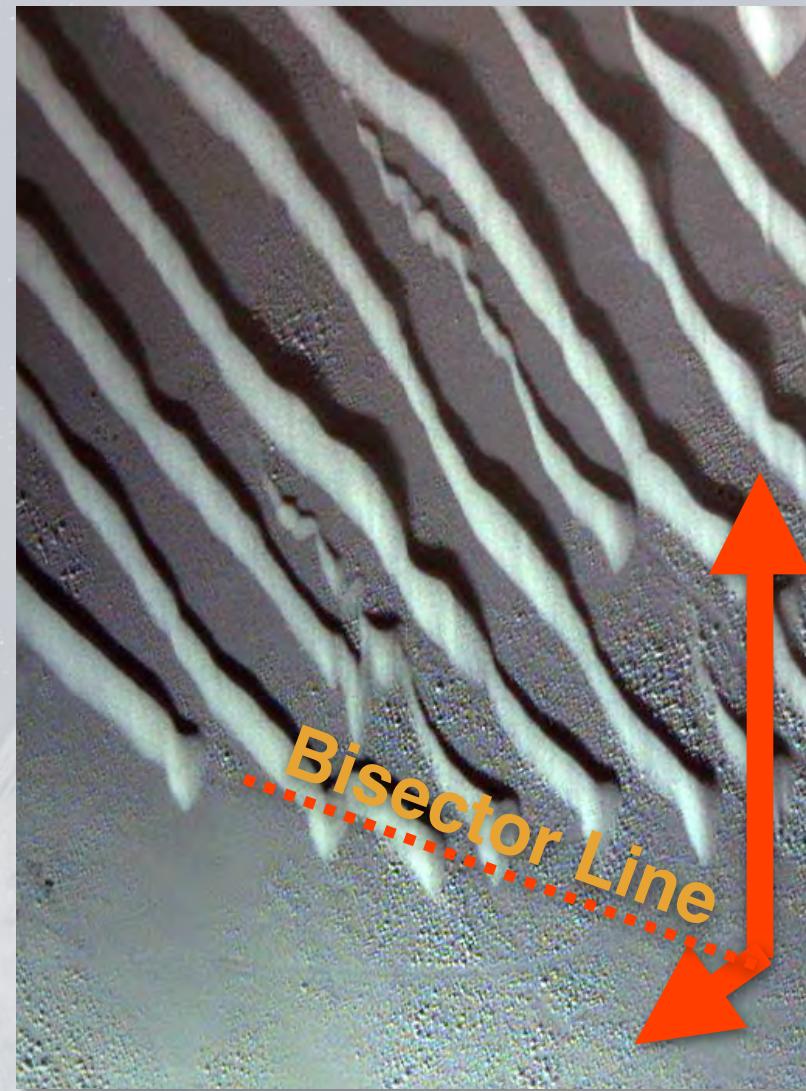
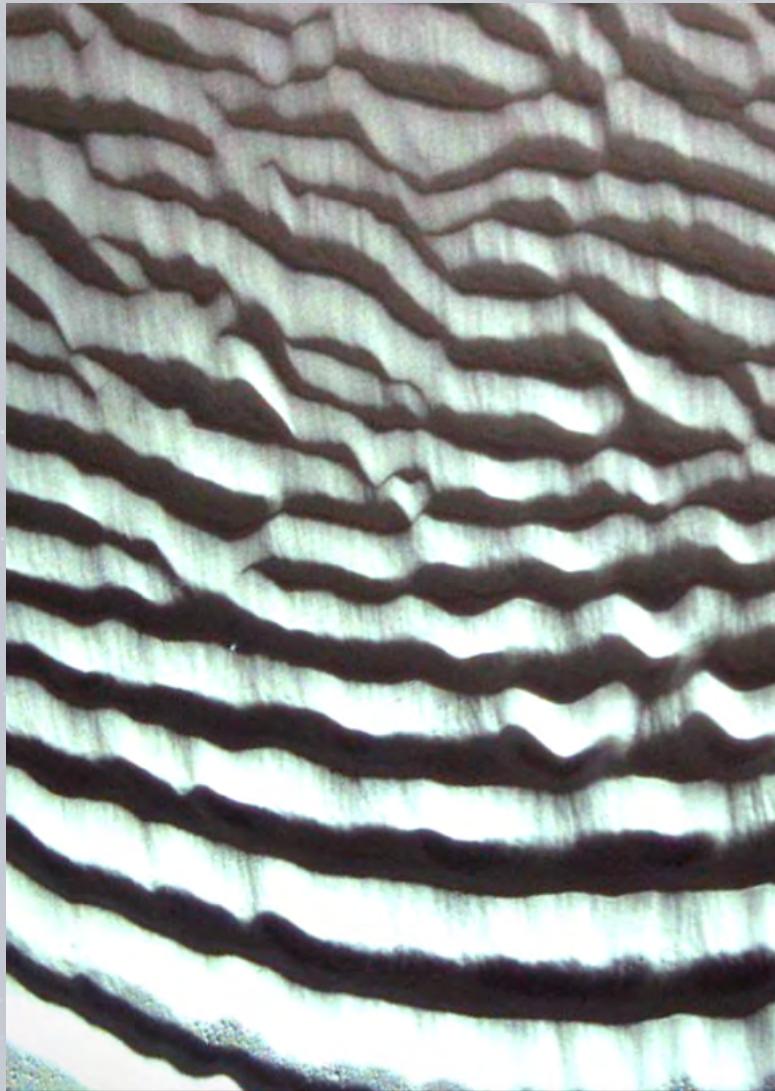


Origin of Dune Orientation on Titan





Origin of Dune Orientation on Titan



Two modes of dune orientation in the lab.

Controlled by the Sediment Supply



Two modes of dune orientation in the lab.



Courrech du Pont et al., (2014)

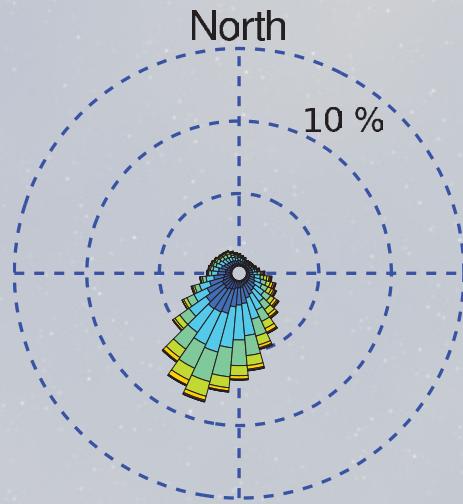
Origin of Dune Orientation on Titan

A Method to Derive two Dune Orientations from Wind Data

Wind Data



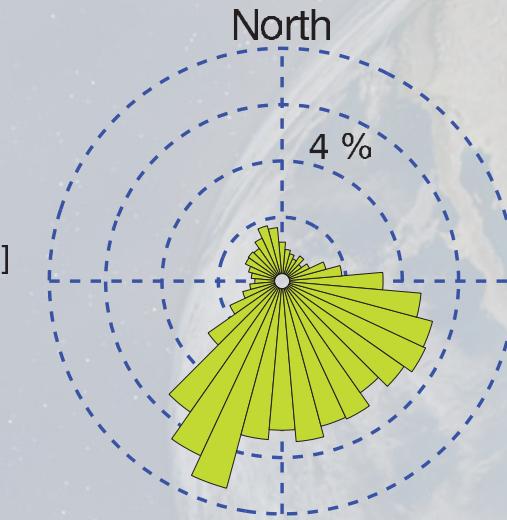
Wind Regime



Sediment transport law

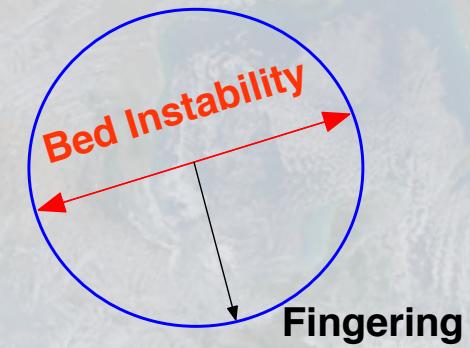
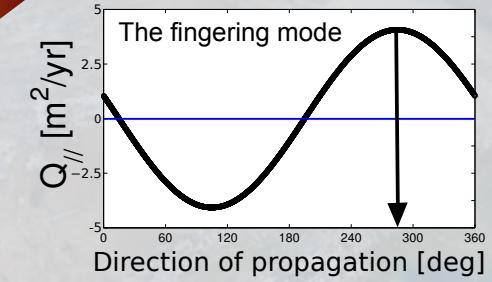
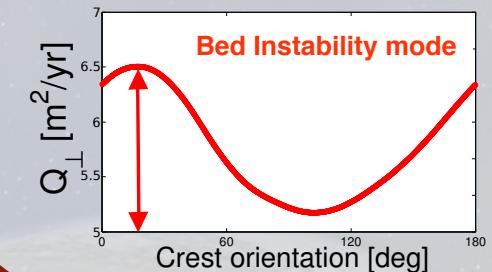
$$Q = f(u, u_{th}^*, d, \rho_s, \rho_f, g)$$

Sediment flux

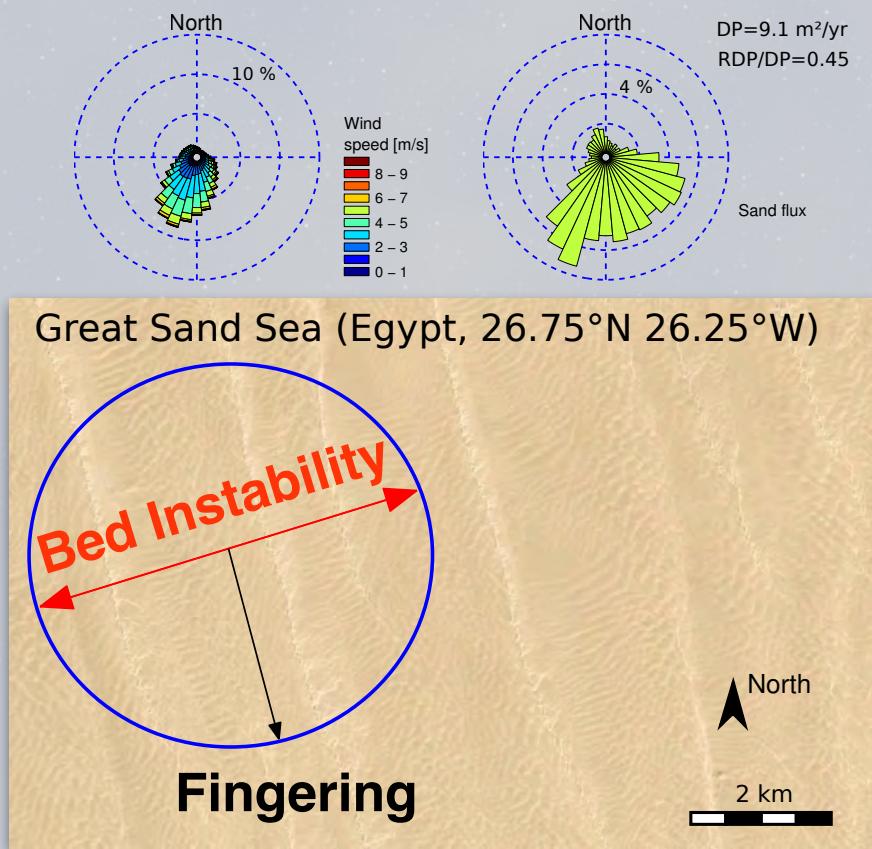


Dune orientation

Courrech du Pont et al., (2014)



A Method to Derive two Dune Orientations from Wind Data



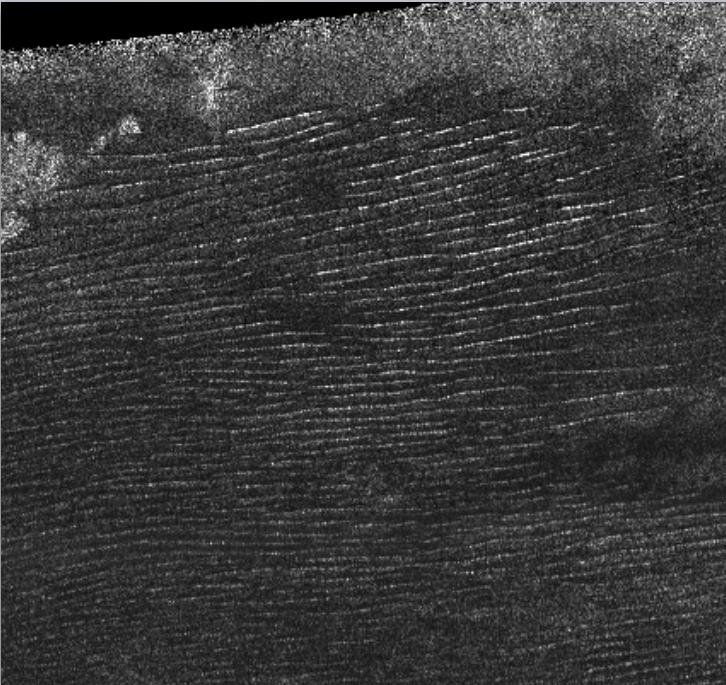
We have a method to derive dune alignment from **wind data** accounting for two **dune growth mechanisms** and their corresponding **crest orientations**

Courrech du Pont & Narteau, in review for Geology



Back to Titan

We need an exhaustive and global map orientation
but SAR images suffers from noise (speckle)

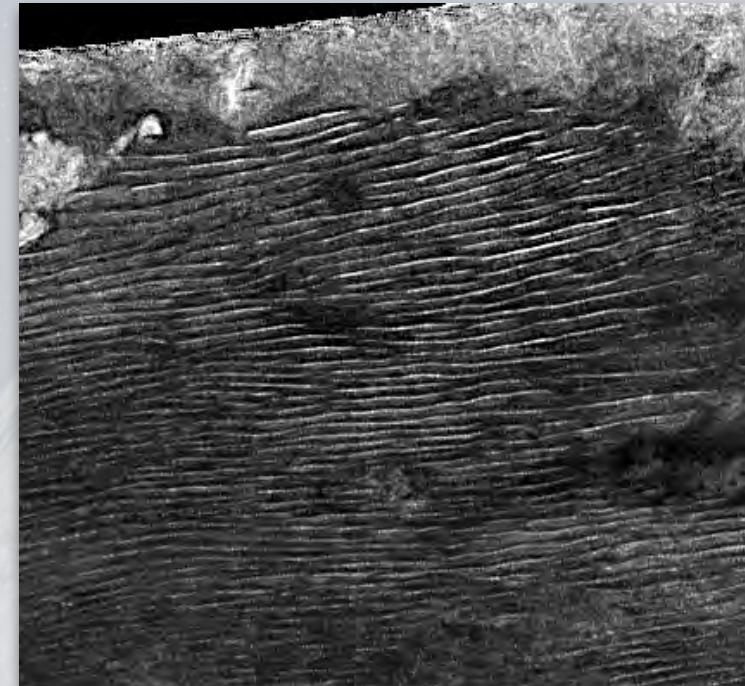
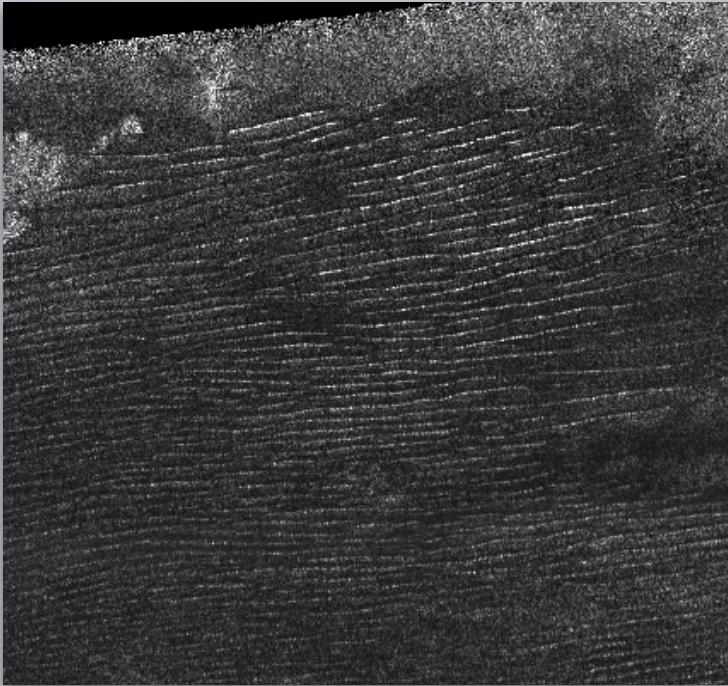


Lucas et al. (*in revision for JGR*)
and presented @AGU 2011 and LPSC 2012



Back to Titan

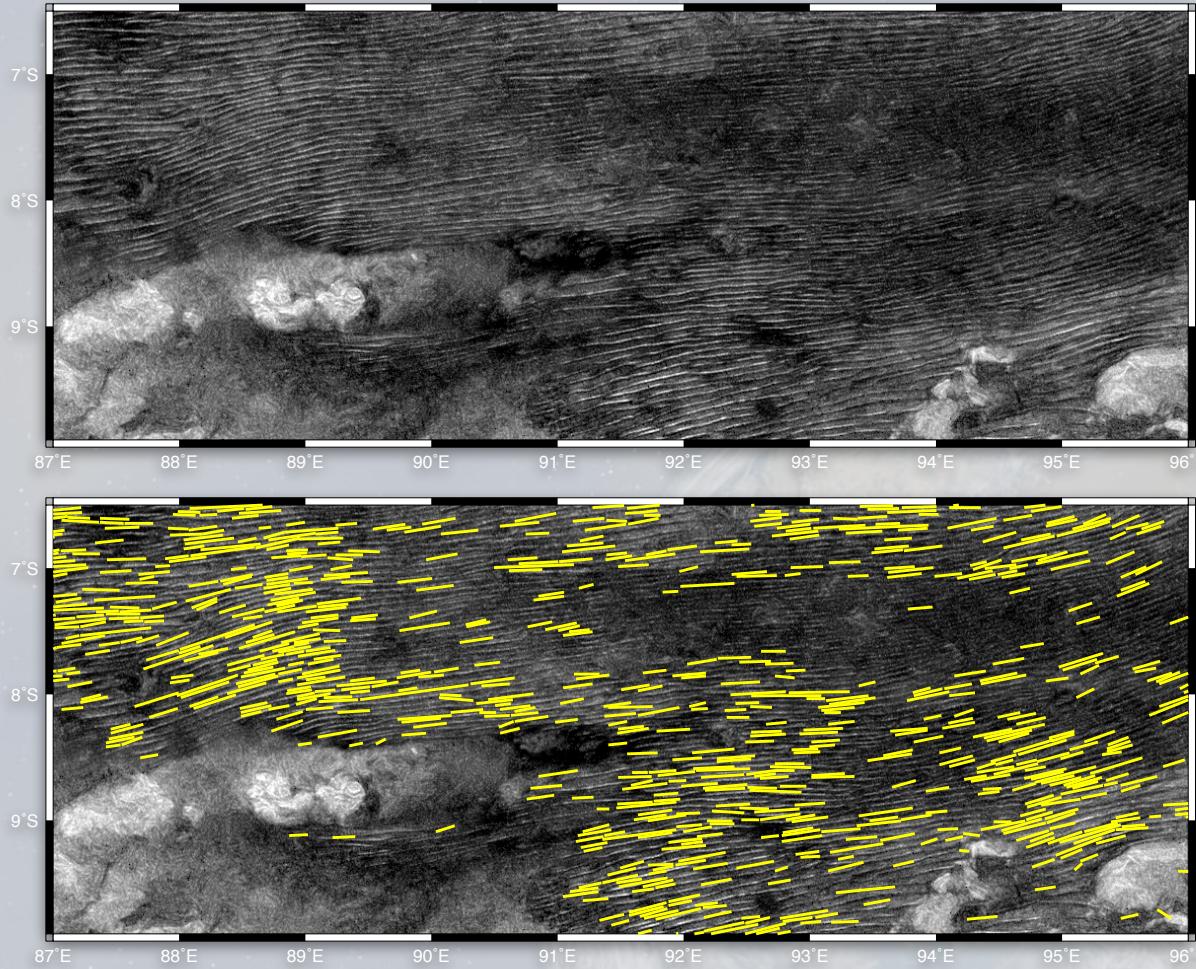
We need an exhaustive and global map orientation
De-noising the SAR data



Lucas et al. (*in revision for JGR*)
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Origin of Dune Orientation on Titan

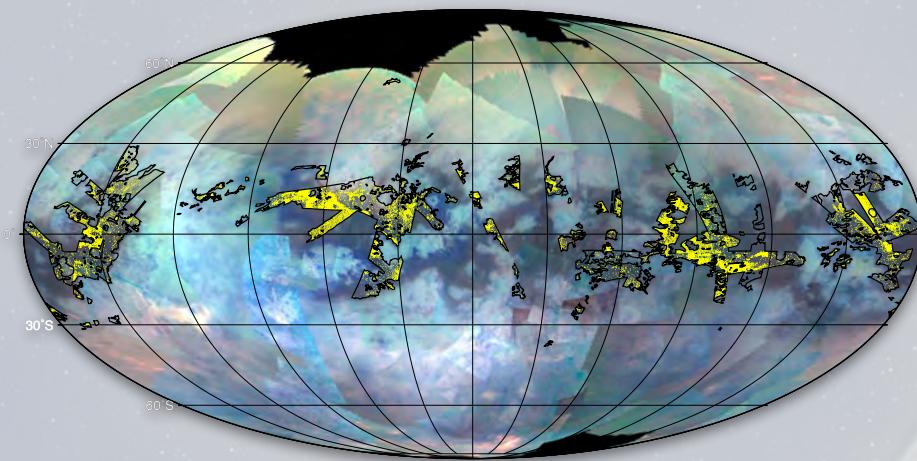
We need an exhaustive and global map orientation

LSD: Linear Segment Detection

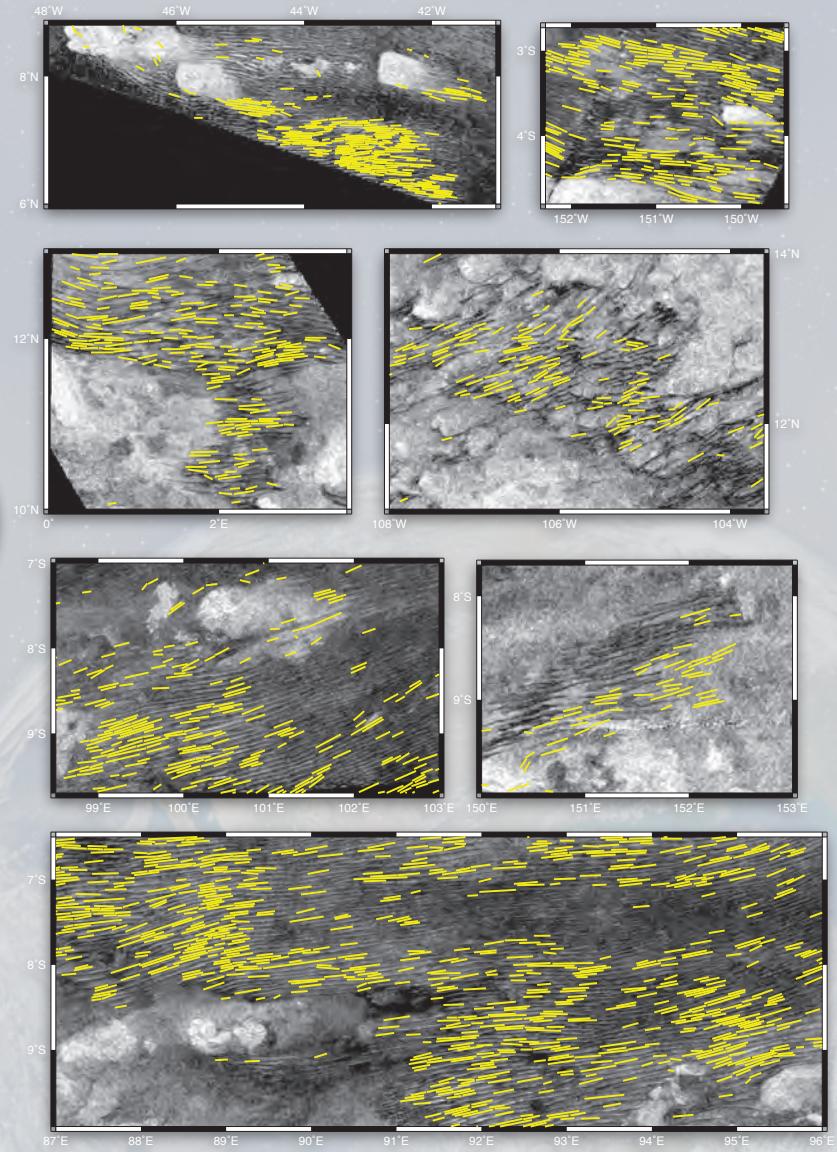


Origin of Dune Orientation on Titan

Global extraction



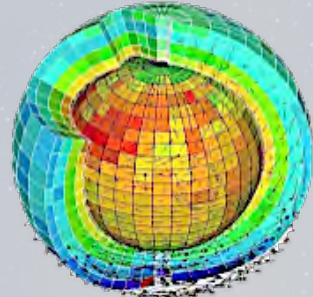
We can do some statistics on that



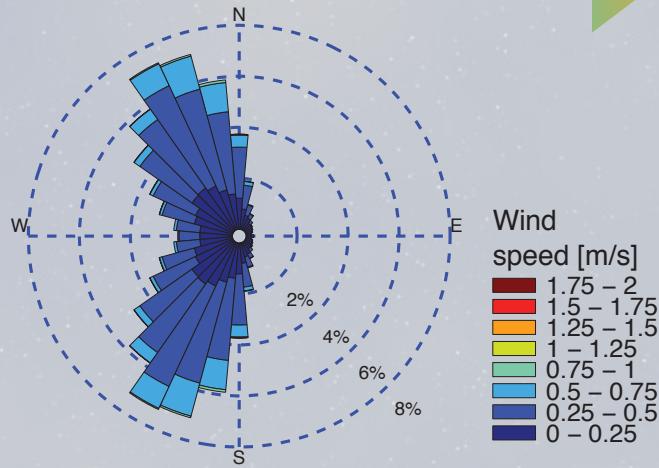
Prediction of Dune Orientation on Titan

Titan's GCM (IPSL-LMD)

Charnay et al., (2013)



Wind Regime

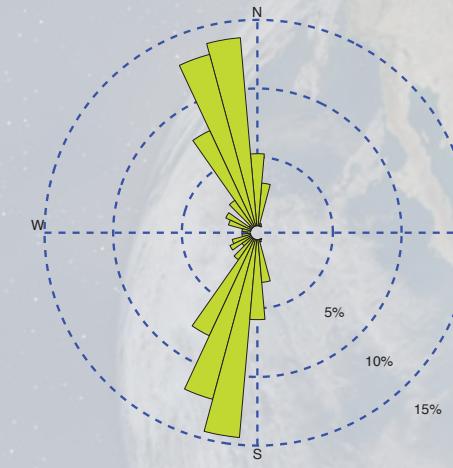


Sediment transport law

Lorenz et al., (1995)

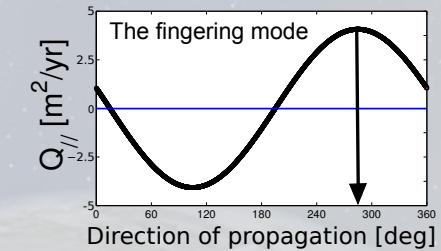
$$Q = f(u, u_{th}^*, d, \rho_s, \rho_f, g)$$

Sediment fluxes

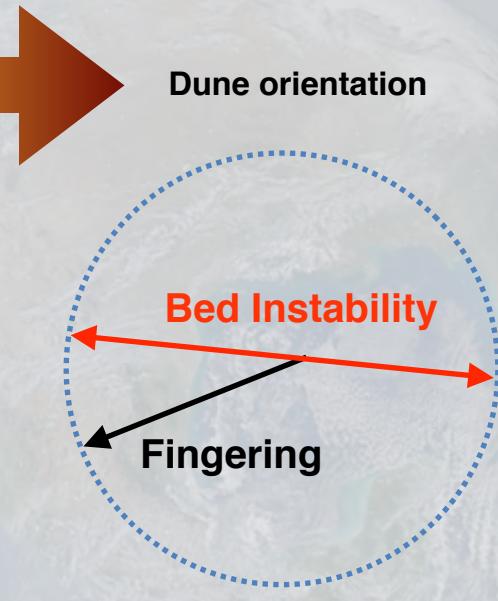


Dune orientation

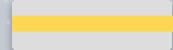
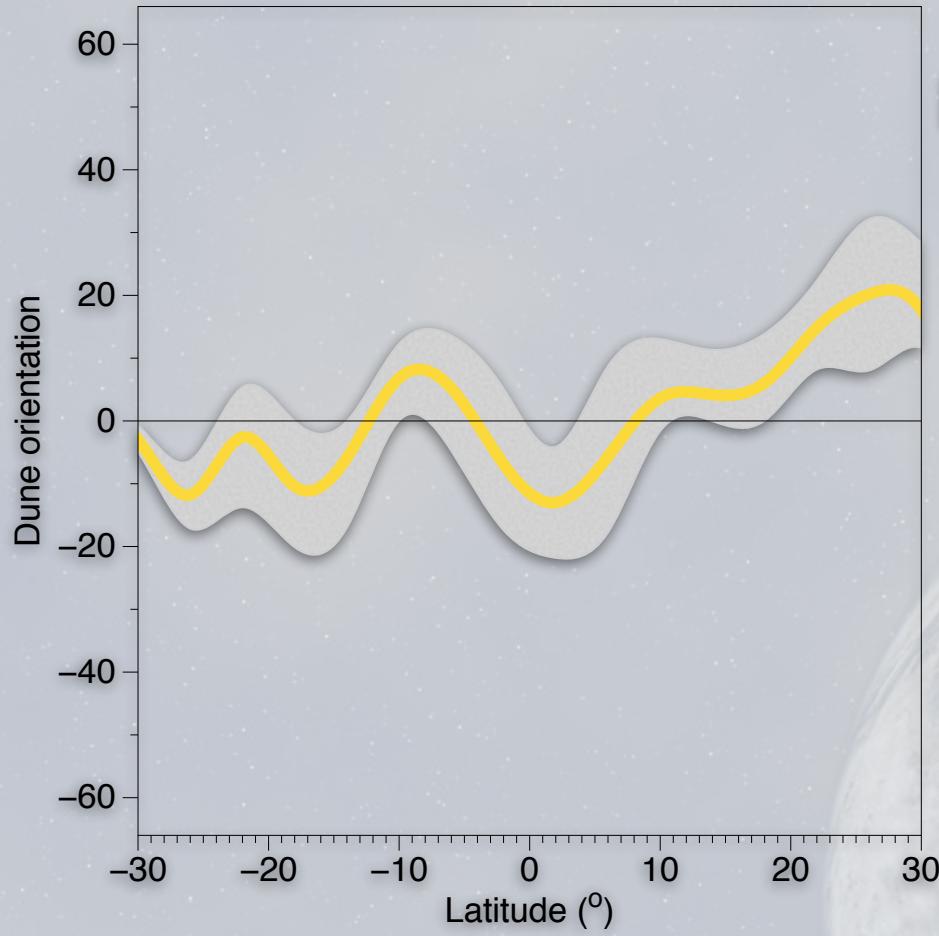
Courrech du Pont et al., (2014)



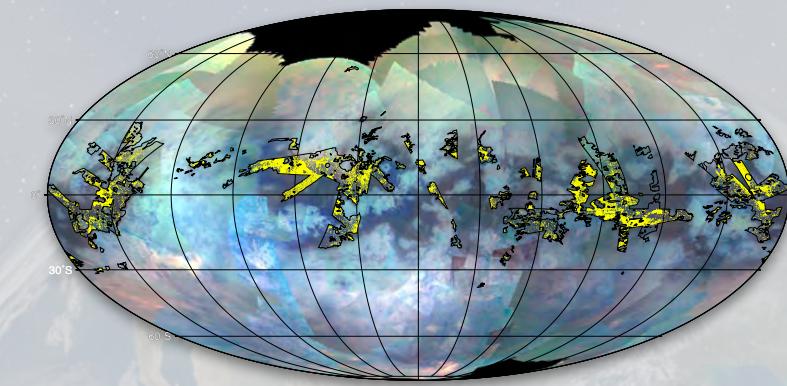
Dune orientation



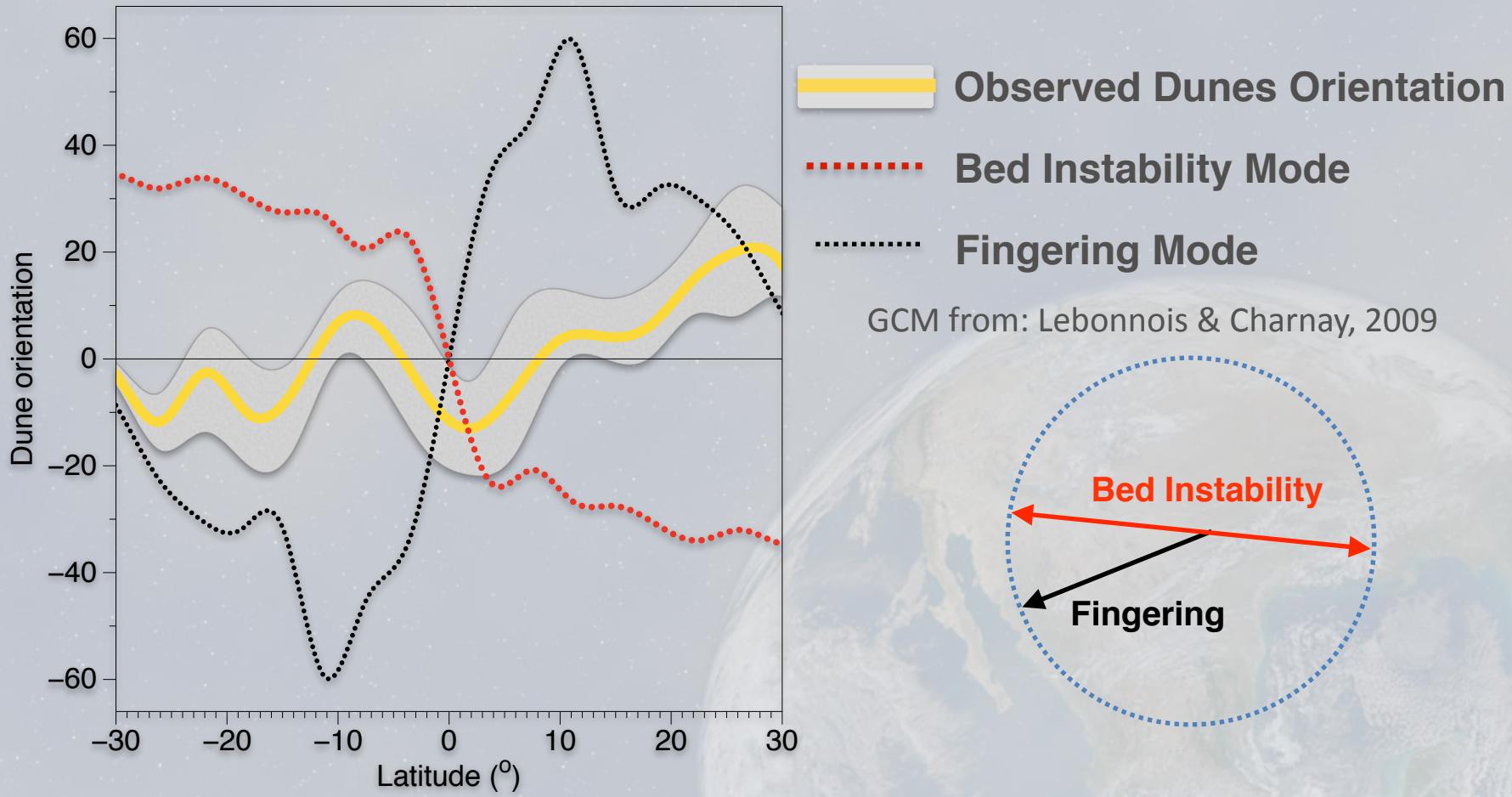
Observed and Predicted Dune Orientation



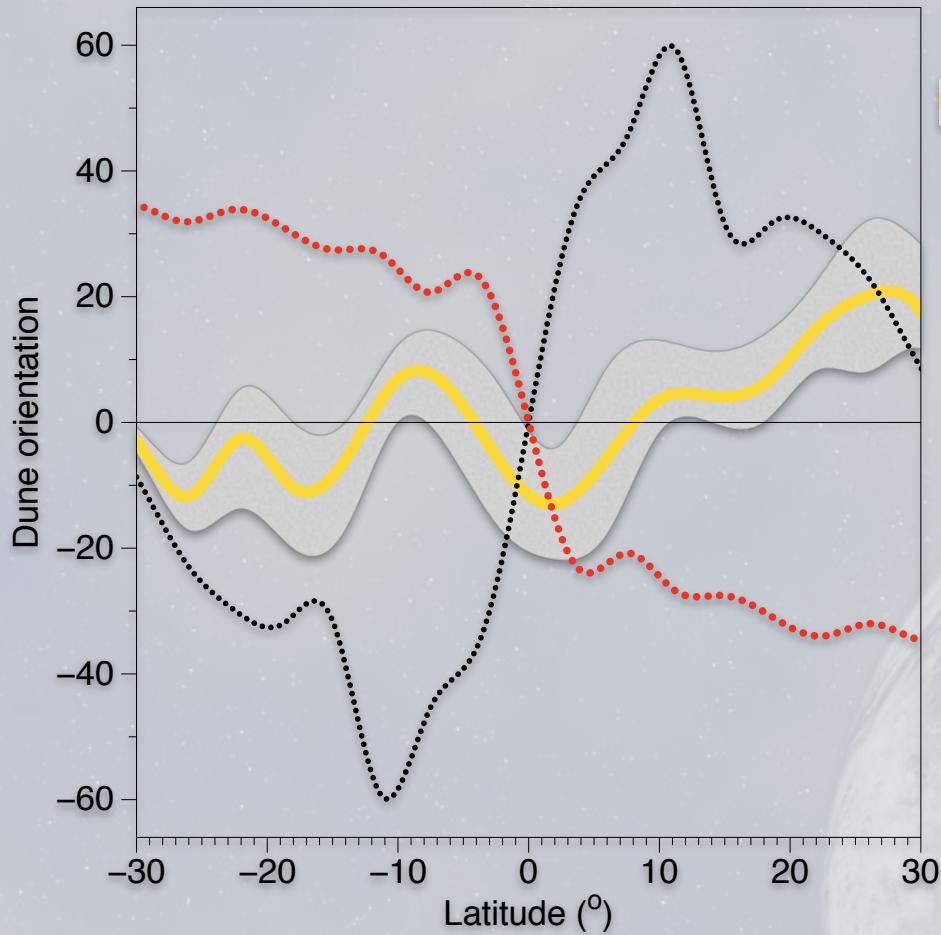
Observed Dunes Orientation



Observed and Predicted Dune Orientation



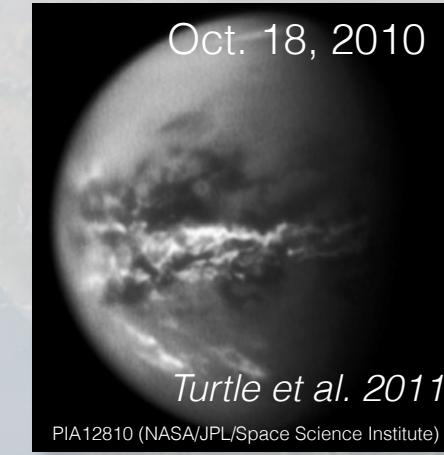
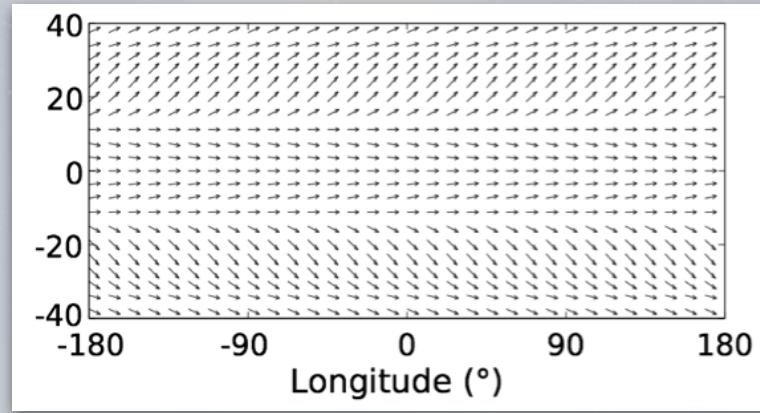
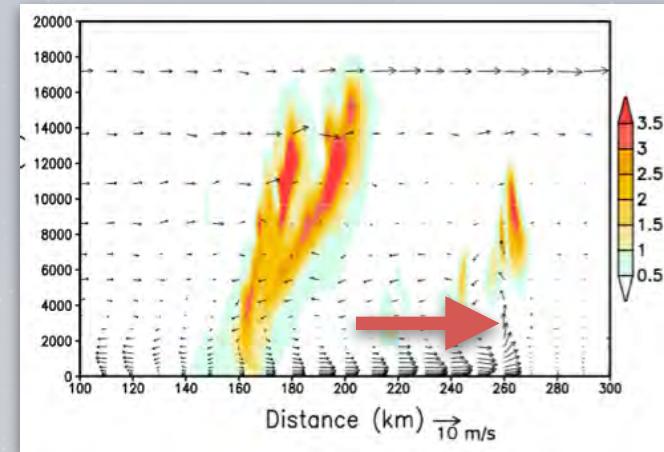
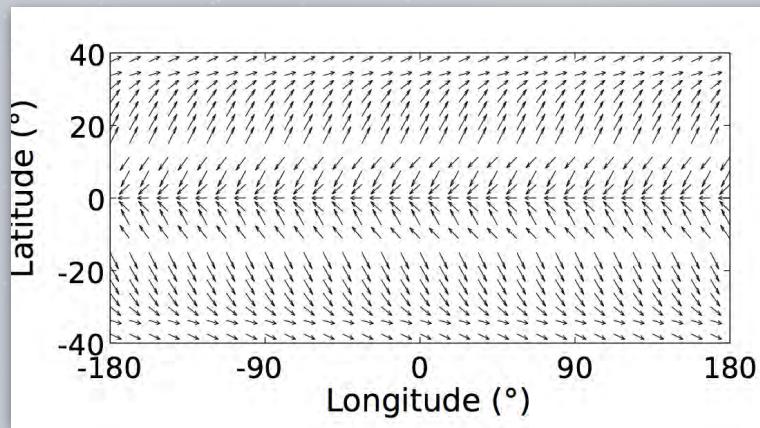
Observed and Predicted Dune Orientation



GCM from: Lebonnois & Charnay, 2009



Equinoctial storms

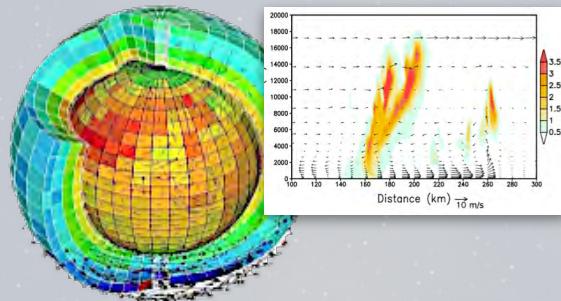


Charnay et al. AGU 2013, &
Charnay et al. in review for Nat. Geosc.

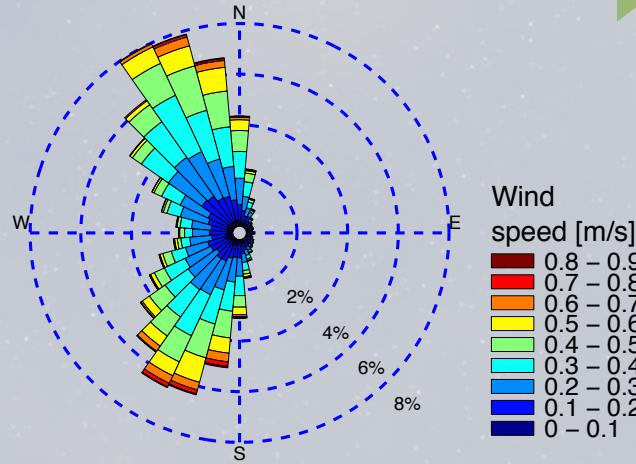
Prediction of Dune Orientation on Titan

Titan's GCM + Storms

Charnay et al., (2013)



Wind Regime



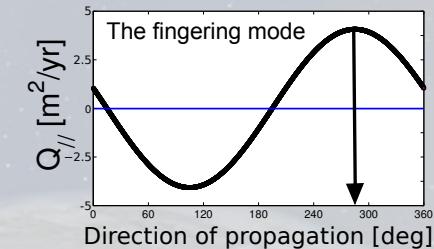
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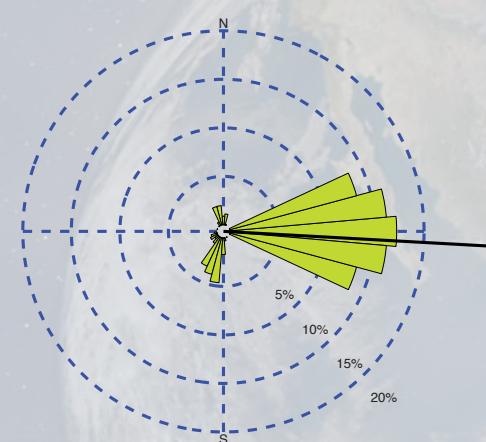
$$Q = f(u, u_{th}^*, d, \rho_s, \rho_f, g)$$

Dune orientation

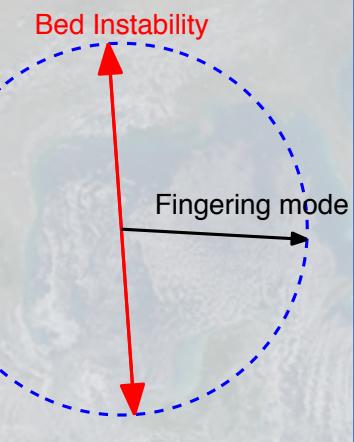
Courrech du Pont et al., (2014)



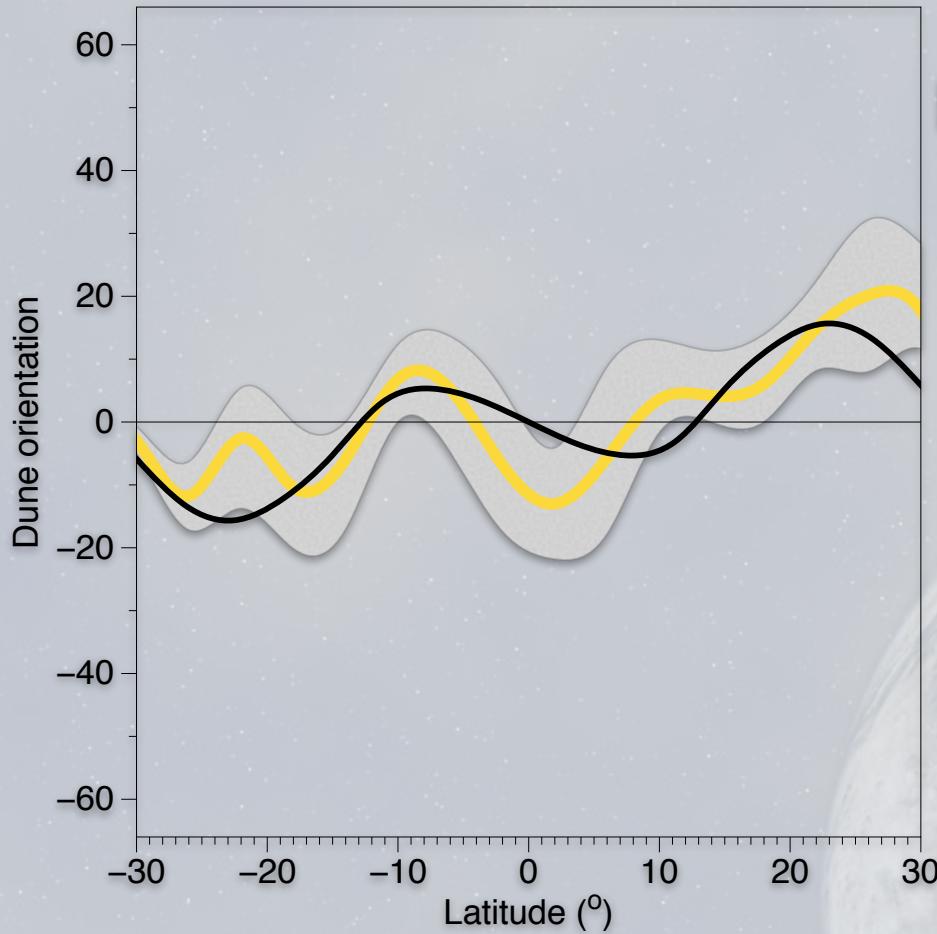
Sediment fluxes



Dune orientation



Observed and Predicted Dune Orientation



Observed Dunes Orientation

Bed Instability Mode

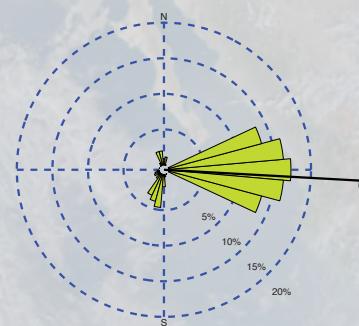
Fingering Mode

Equinoctial Storms:

GCM from: Lebonnois & Charnay, 2009
Meso-scale in Charnay et al. under review

Fingering Mode with Storms

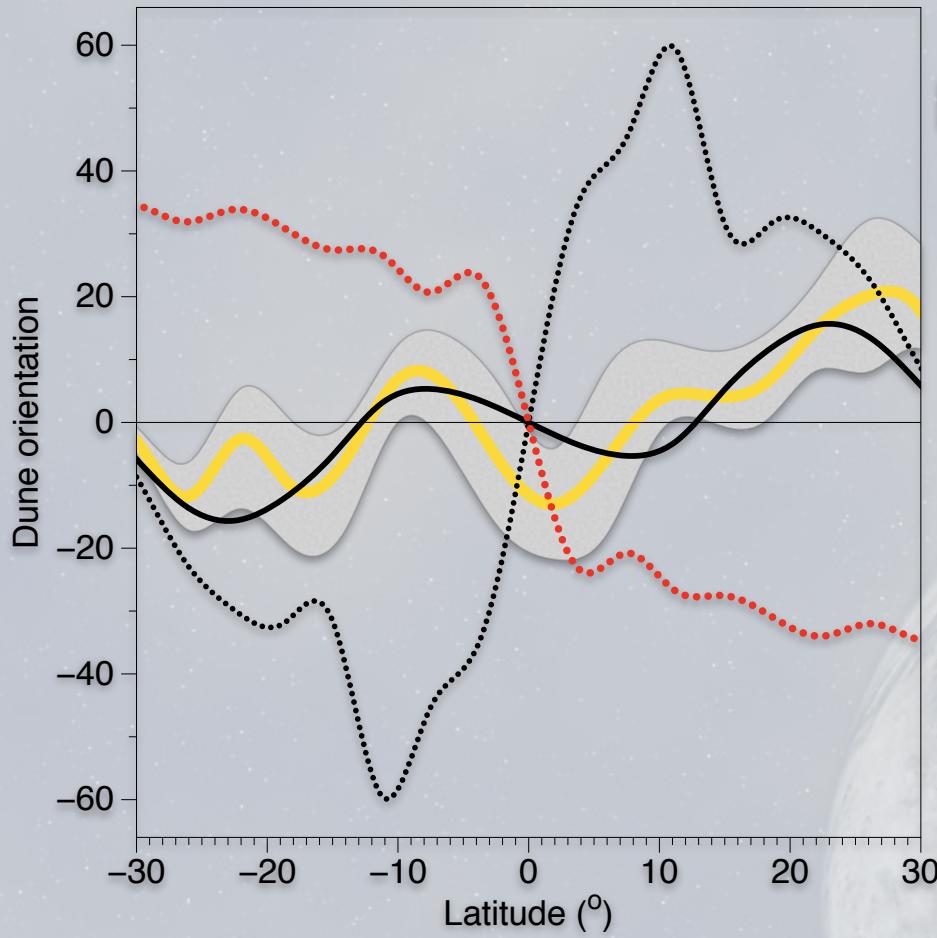
Sediment fluxes



Dune orientation



Observed and Predicted Dune Orientation



Observed Dunes Orientation

Bed Instability Mode

Fingering Mode

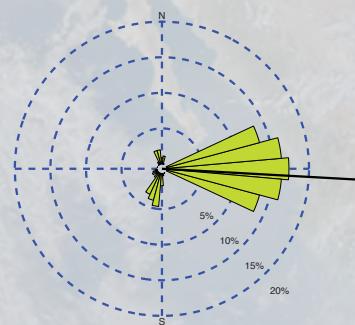
Equinoctial Storms:

GCM from: Lebonnois & Charnay, 2009

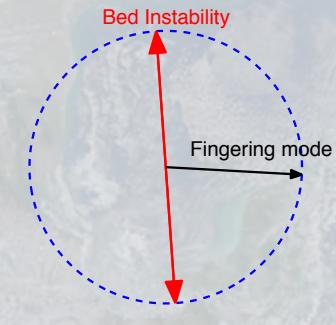
Meso-scale in Charnay et al. under review

Fingering Mode with Storms

Sediment fluxes



Dune orientation

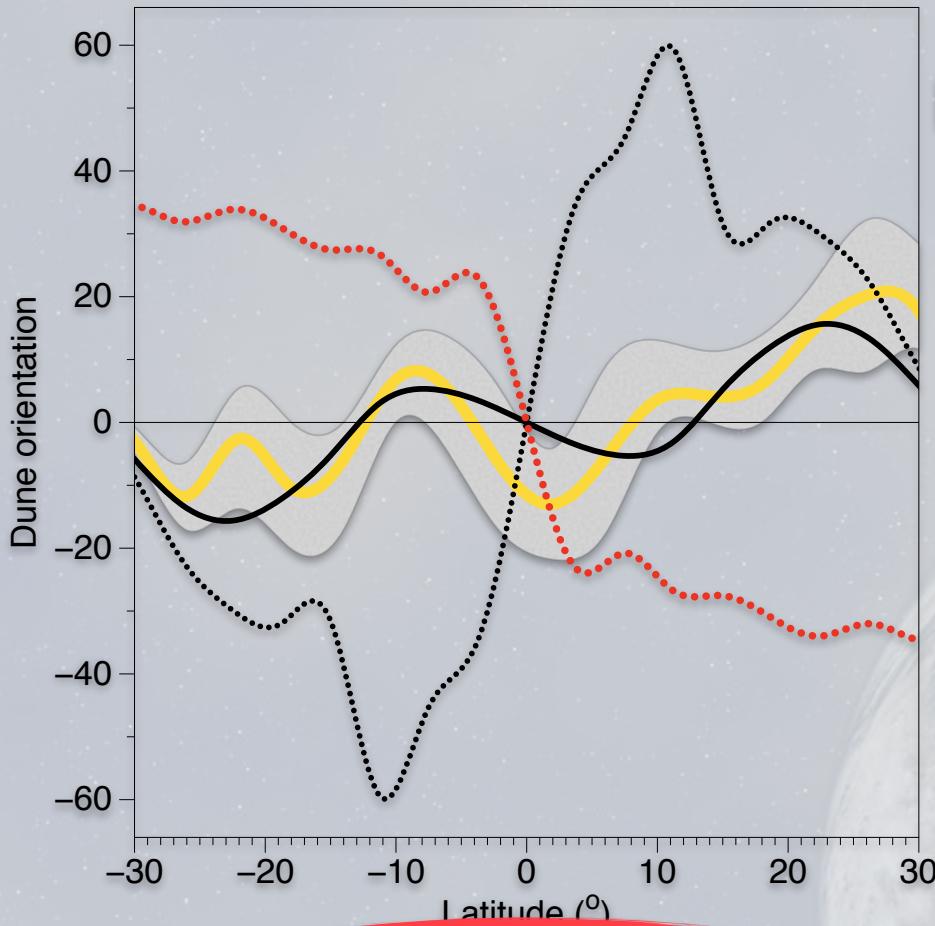


Origin of Dune Orientation on Titan



Bed instability Mode with Storms

Observed and Predicted Dune Orientation



Observed Dunes Orientation

Bed Instability Mode

Fingering Mode

Equinoctial Storms:

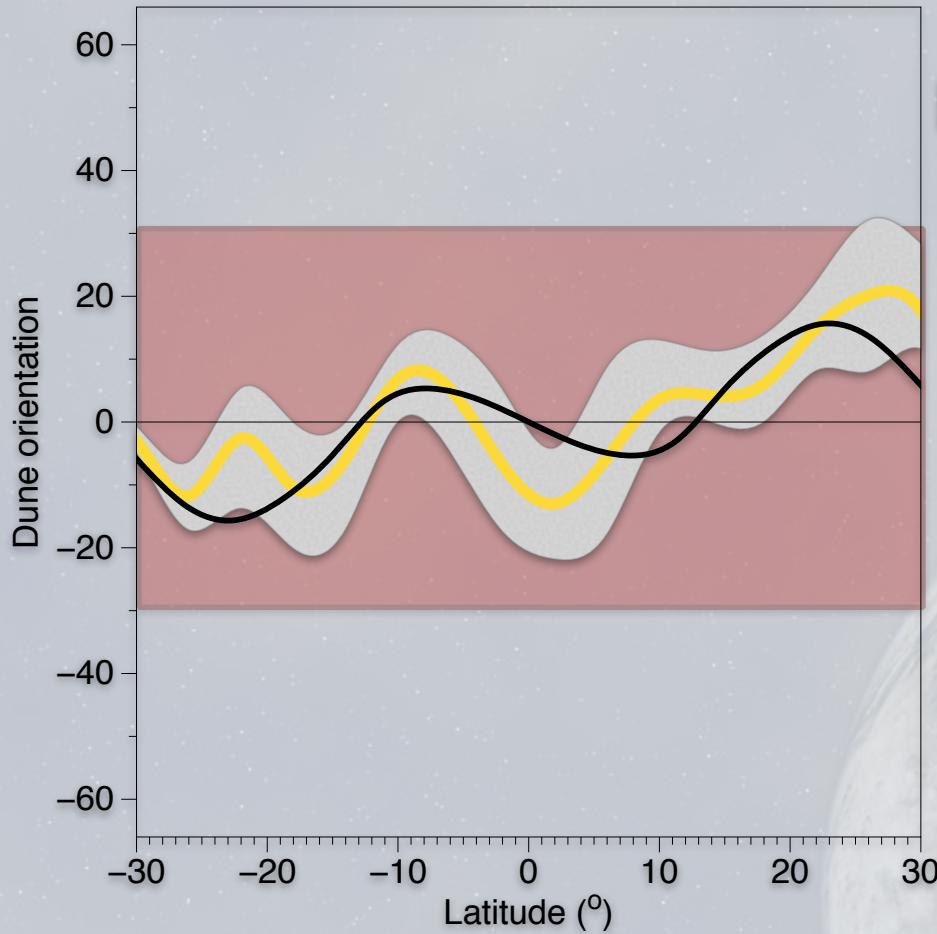
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Fingering Mode with Storms

Bed Instability Mode

Bed instability Mode with Storms

Observed and Predicted Dune Orientation



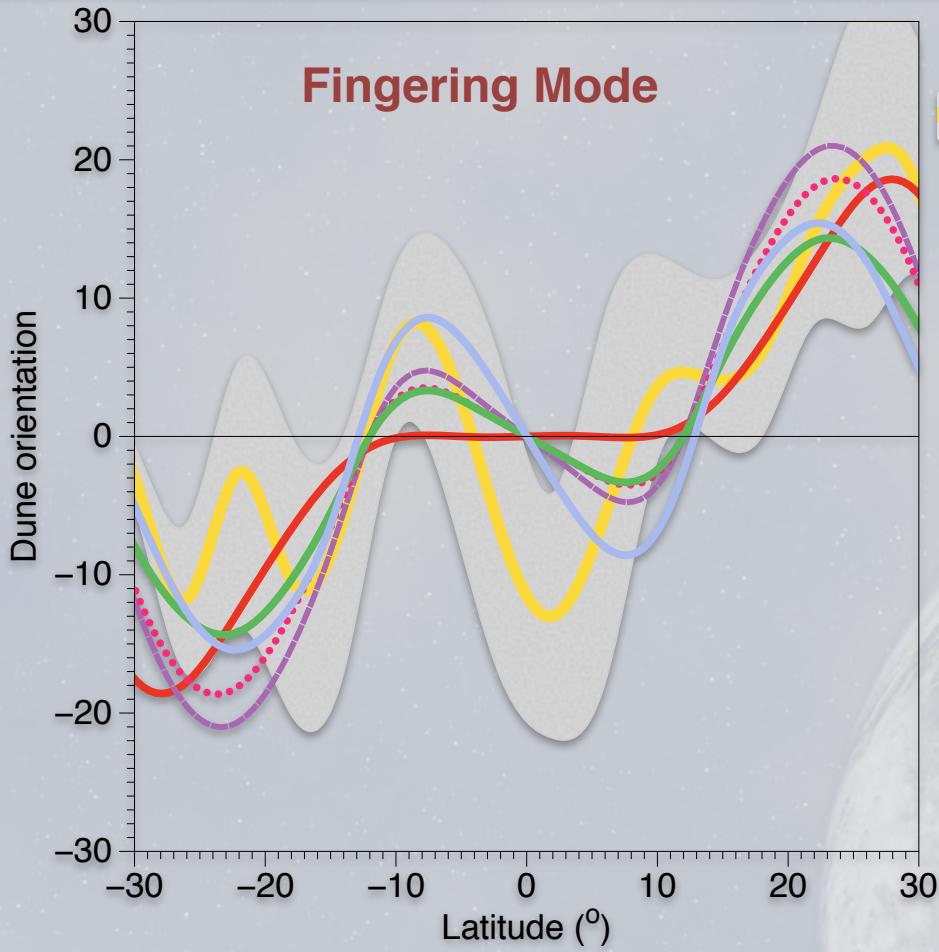
Observed Dunes Orientation

Equinoctial Storms:

GCM from: Lebonnois & Charnay, 2009
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Fingering Mode with Storms

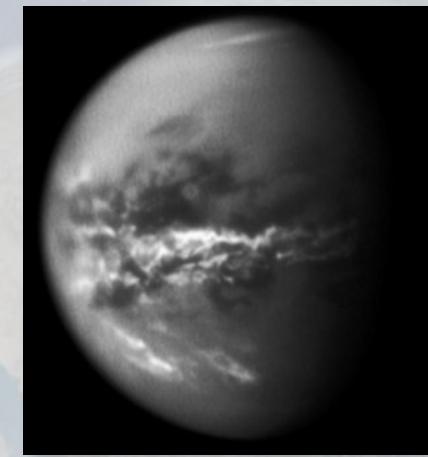
Storm properties effects on predictions



Observed Dunes Orientation

Equinoctial Storms:

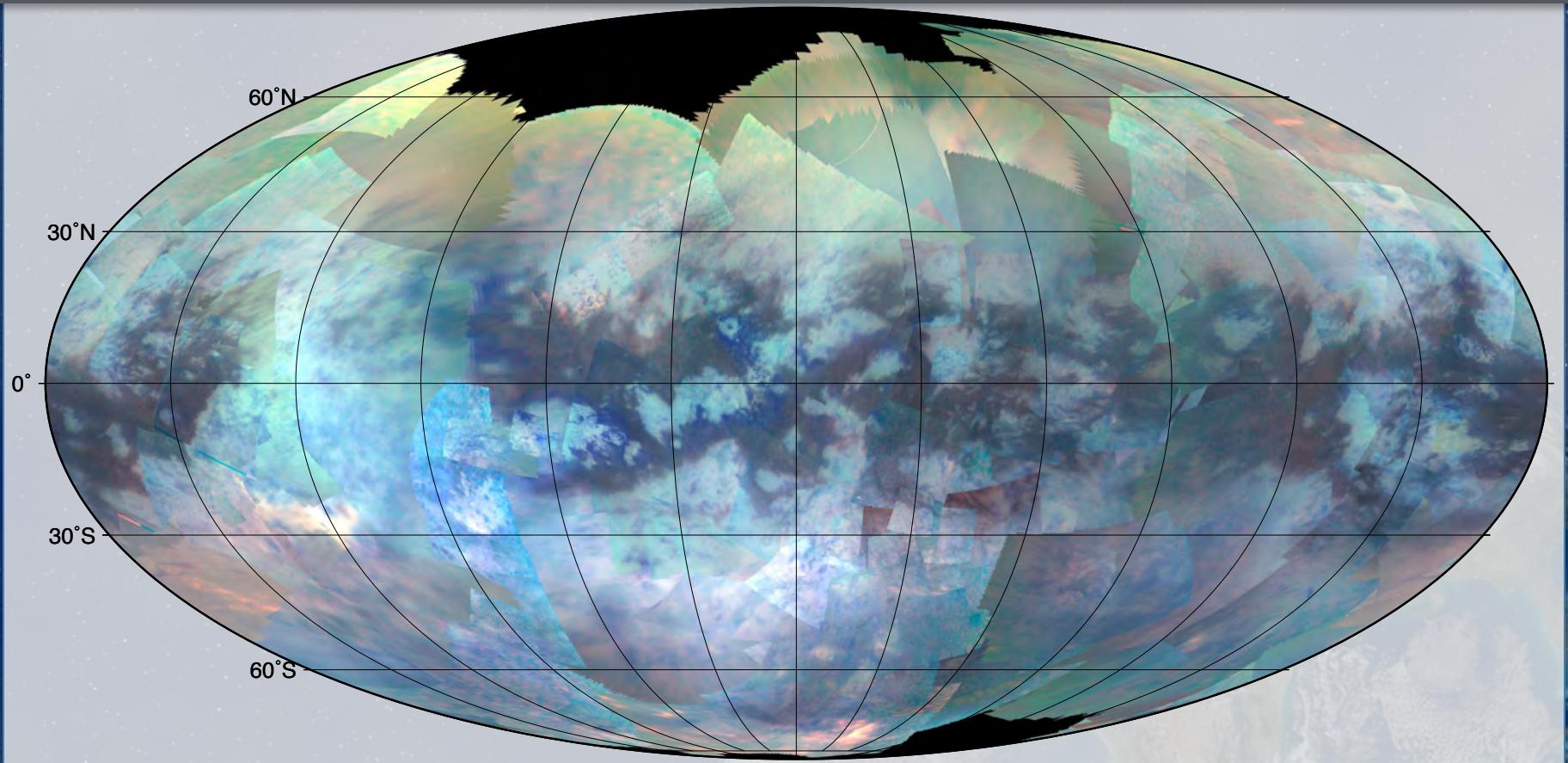
GCM from: Lebonnois & Charnay, 2009
Meso-scale in Charnay et al. under review



- Size of storms
- Frequency of storms
- Friction Velocity threshold



Can we explain the equatorial segregation?

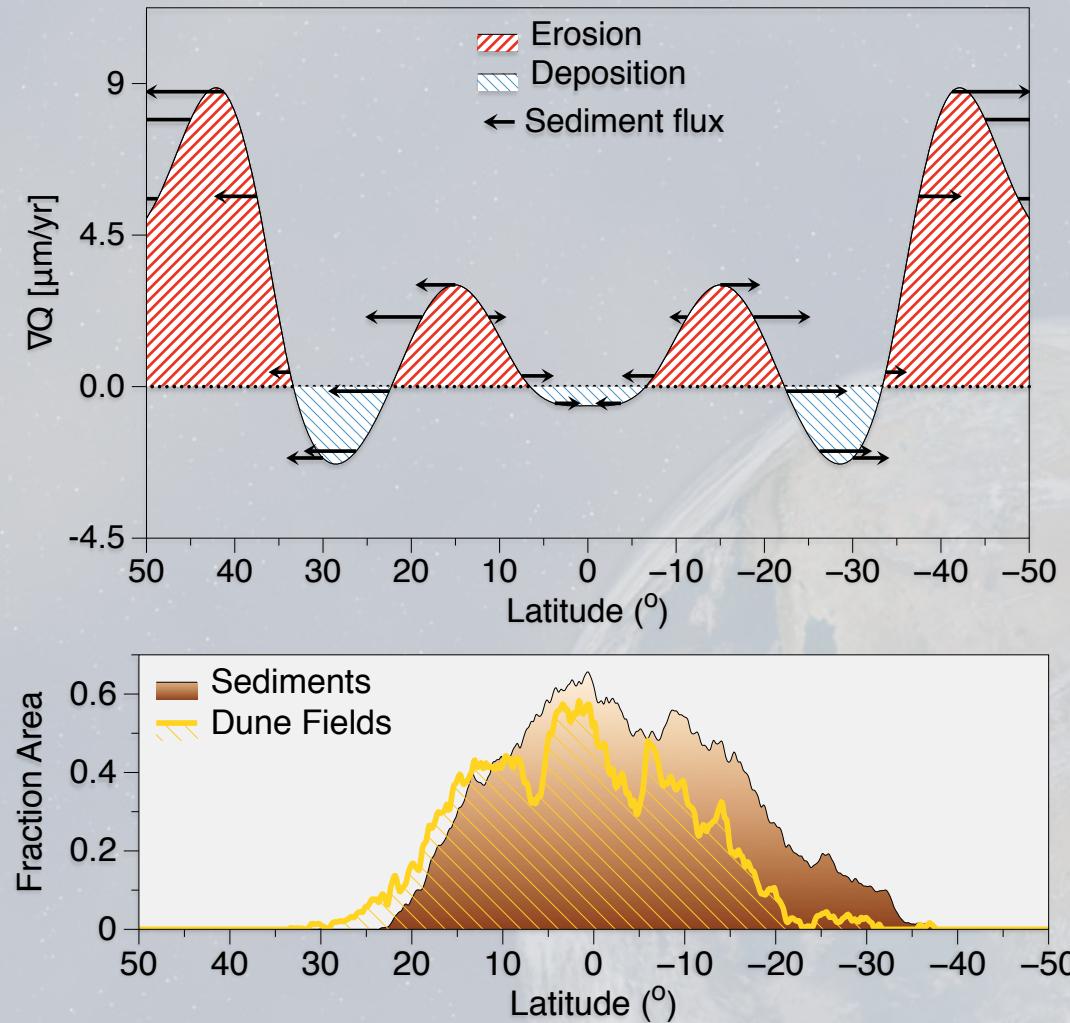


2 500 km (eq.)

Conservation of mass
using wind data :

$$\partial_x Q = -\partial_t h$$

Can we explain the equatorial segregation?





New Insights on Titan's dune orientation

How to explain their geographic confinement?

- ✓ Sand fluxes maintain the sediment in the equatorial belt

How to explain their shape and orientation?

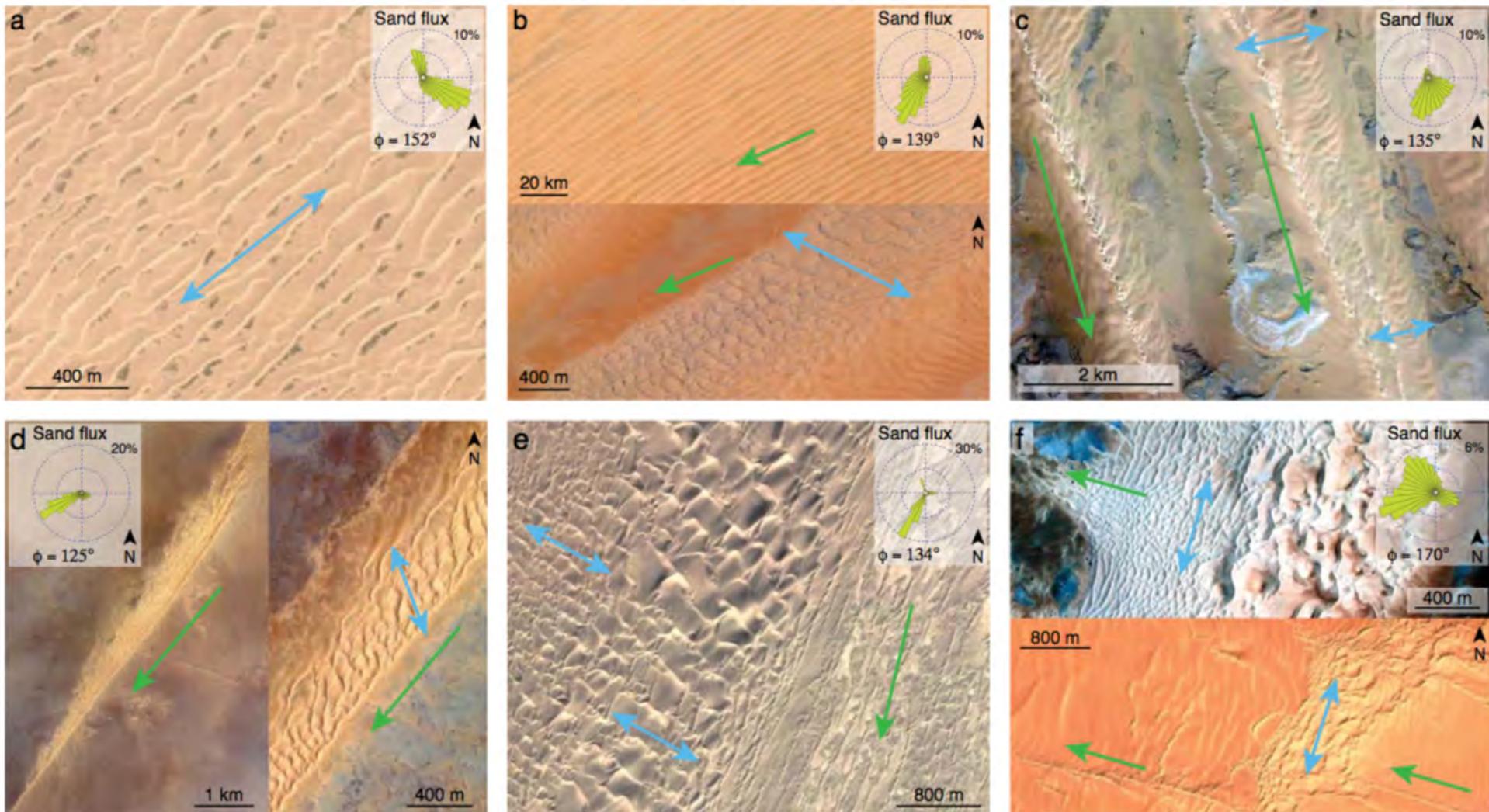
- ✓ Dunes form in the fingering mode.
- ✓ Controlled by the **current two-winds regime** (secondary incl.)
- ✓ Elongating (“fingering”) mode of growth

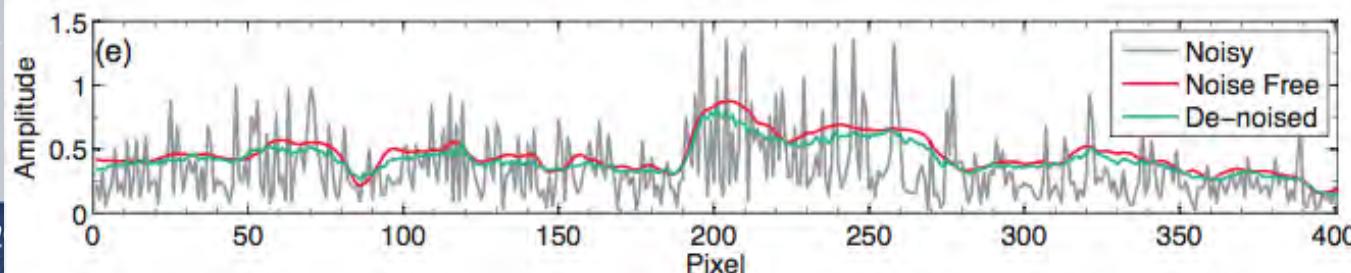
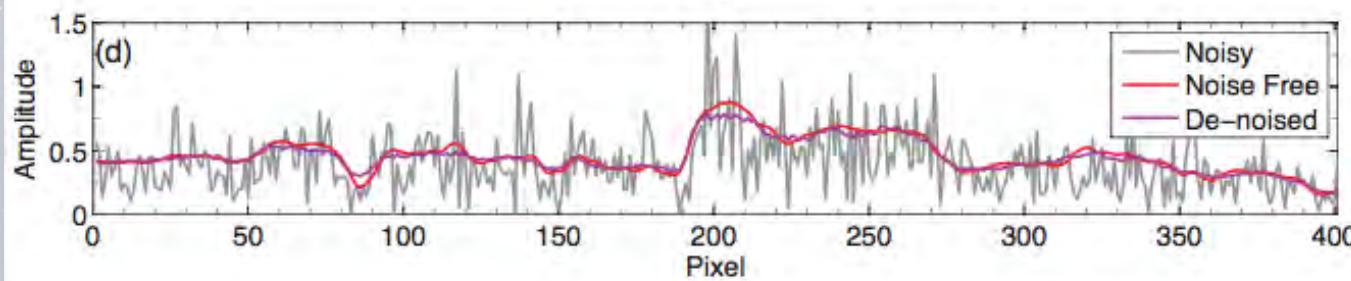
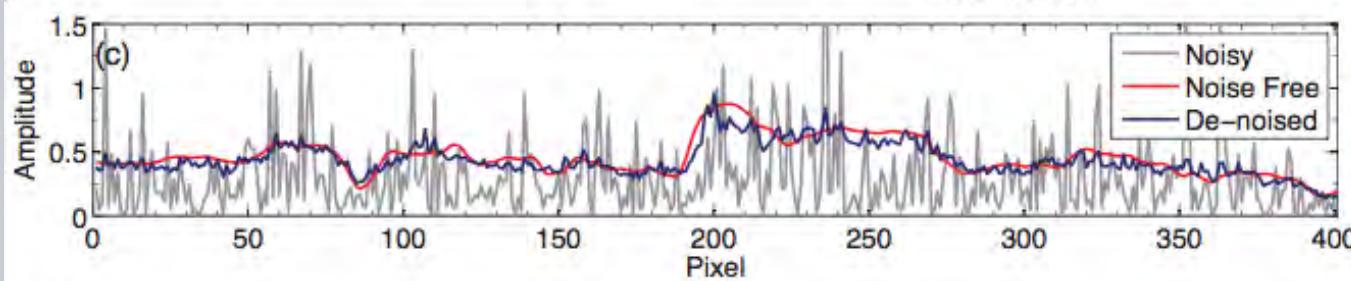
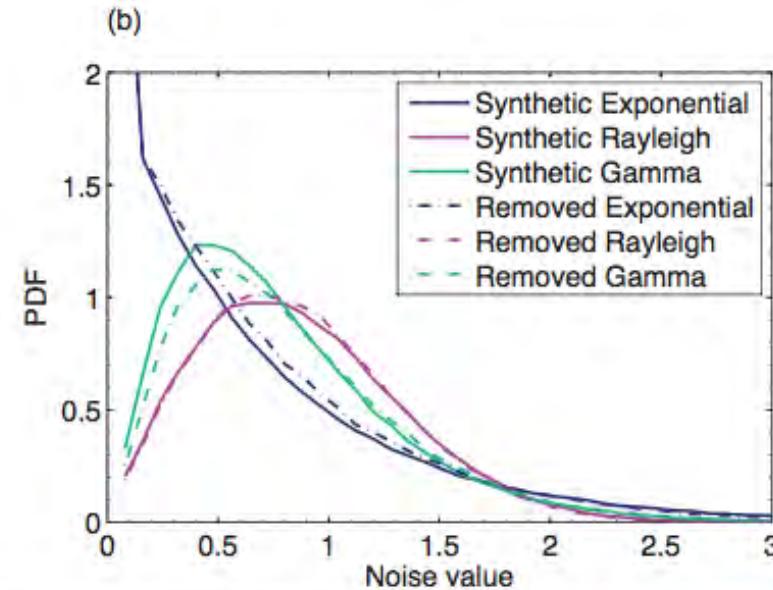
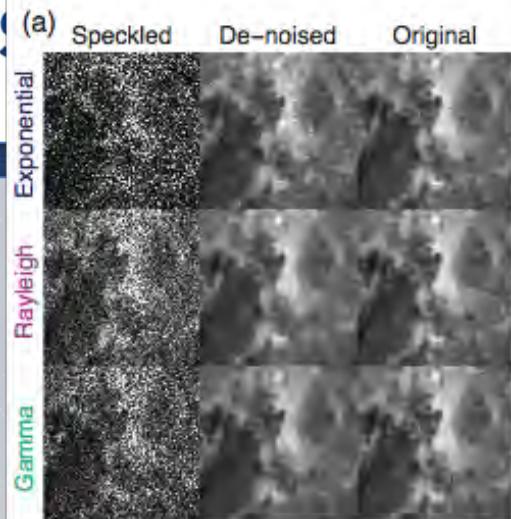
How to explain their direction of propagation?

- ✓ Dunes elongate like fingers on a non-erodible bedrock from a localized source (where?)



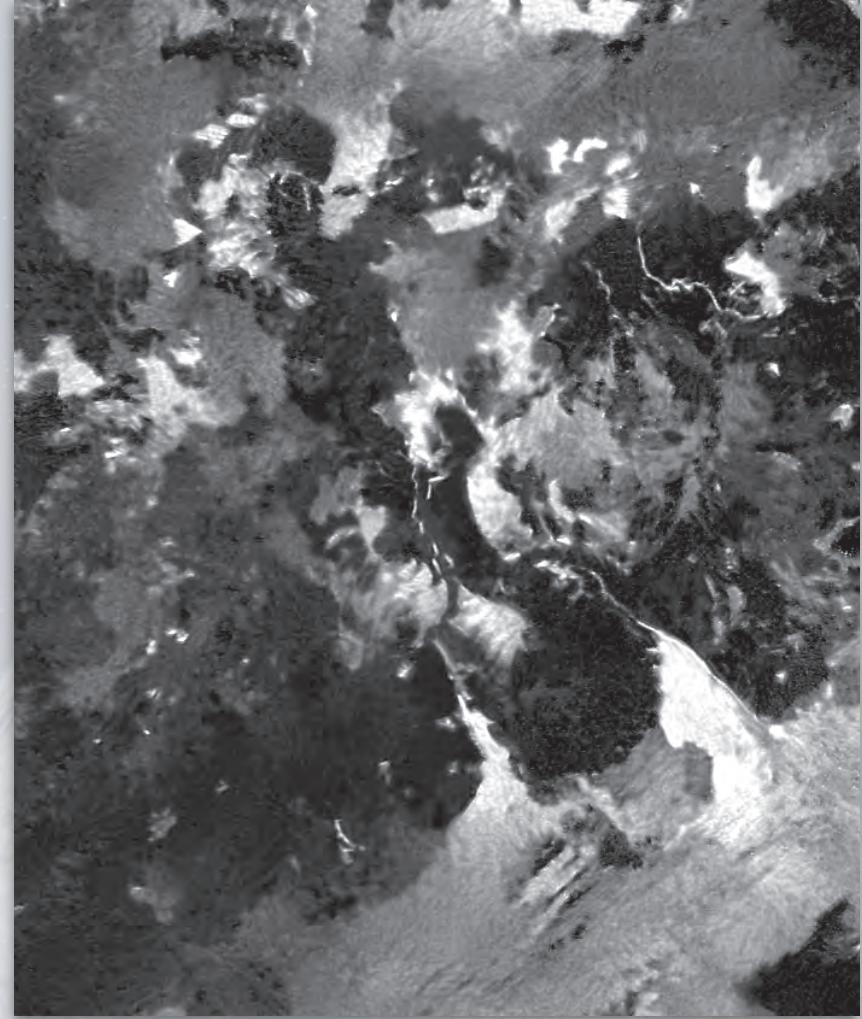
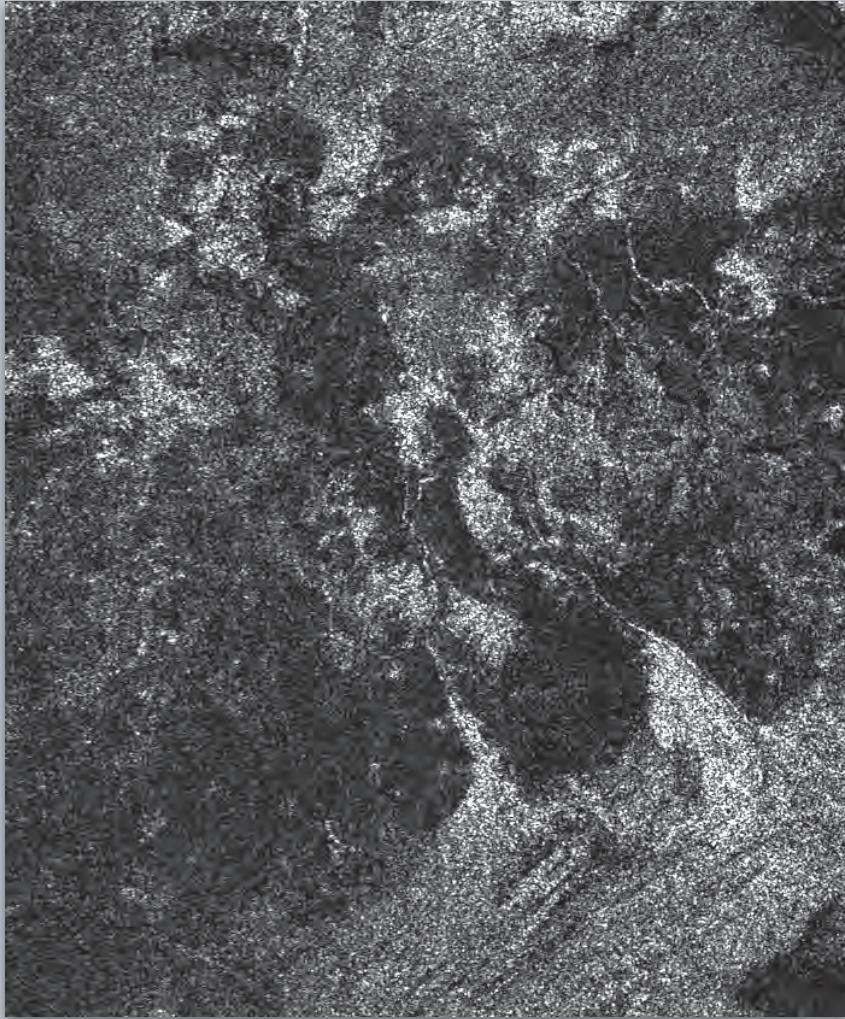
Origin of Dune Orientation on Titan





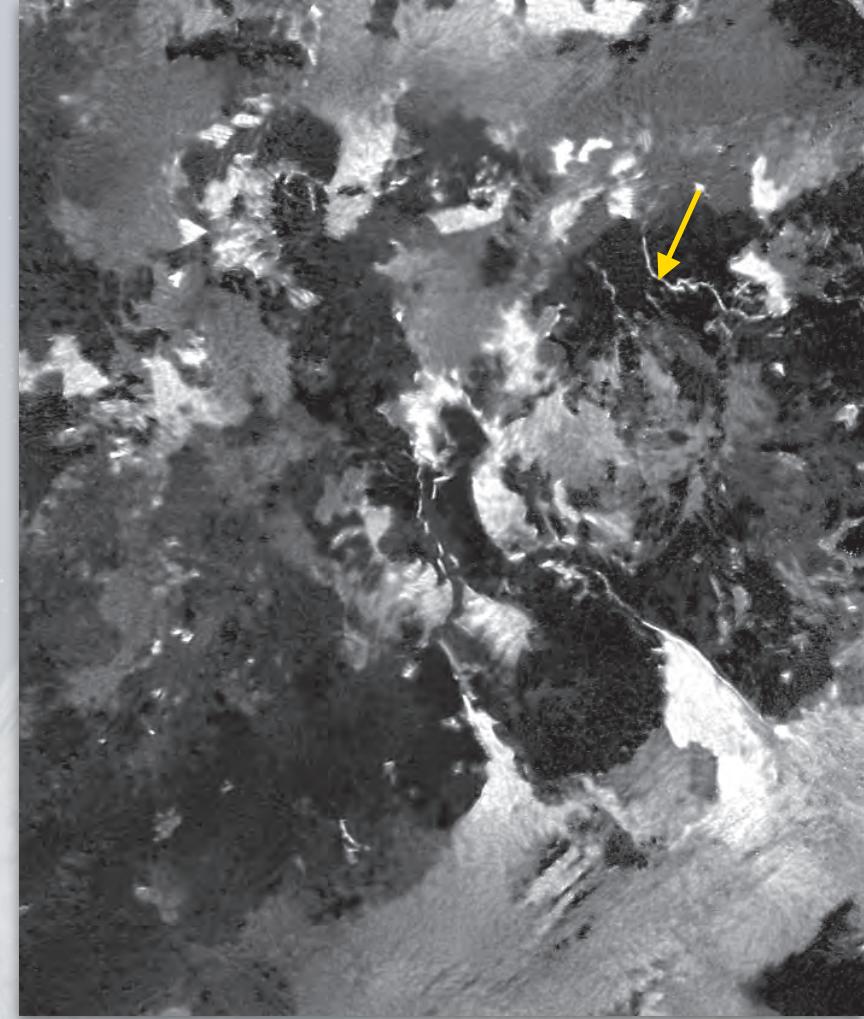
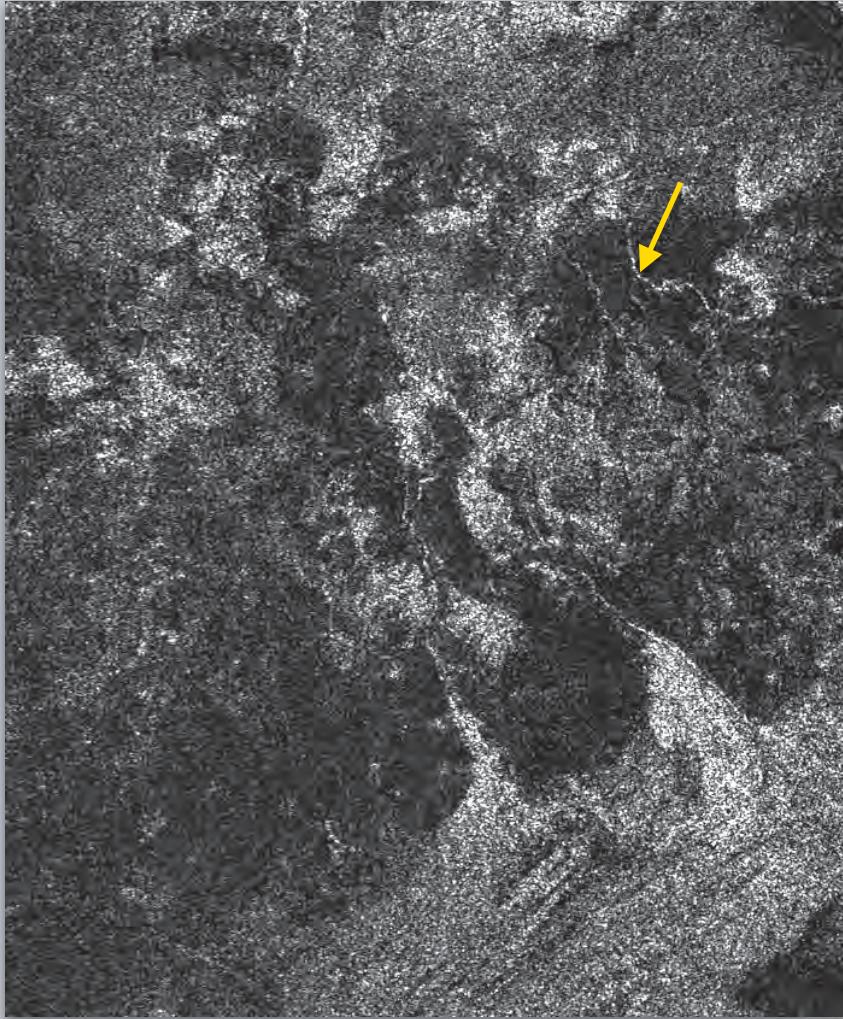


De-noising the SAR data

Lucas et al. (*in revision for JGR*)

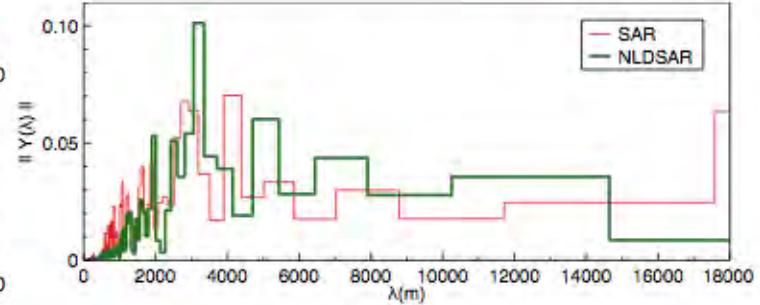
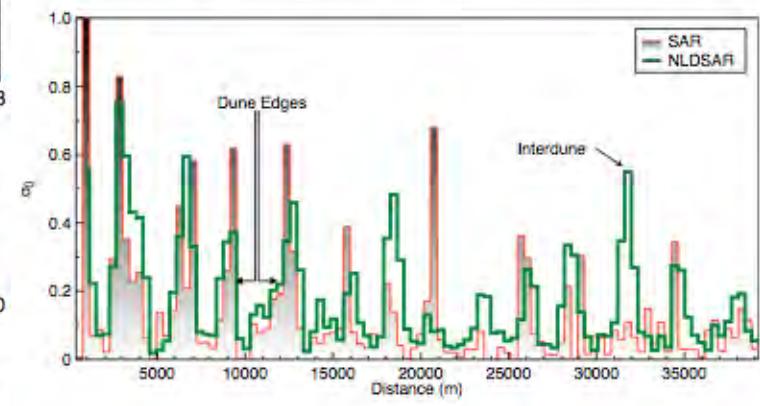
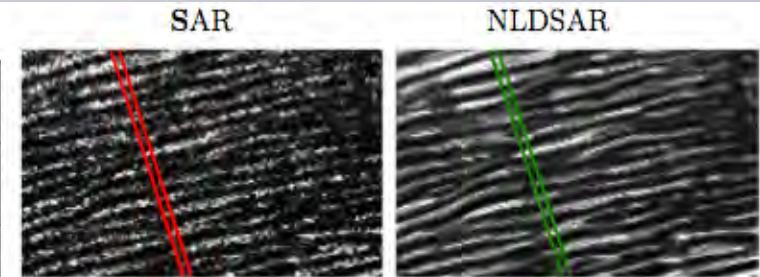
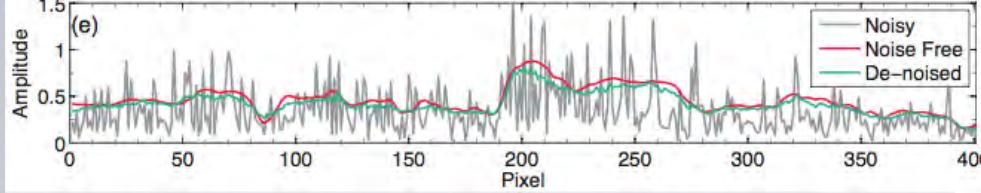
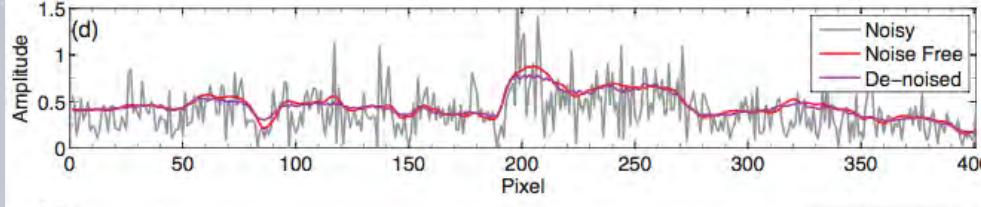
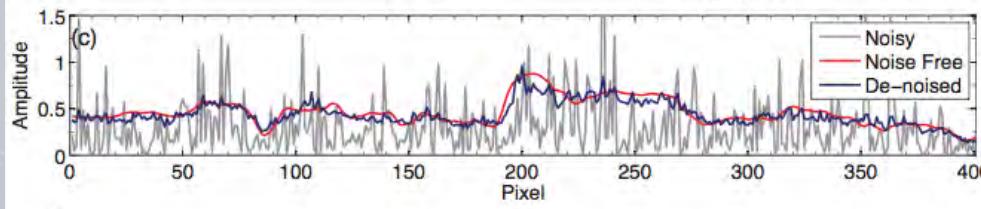
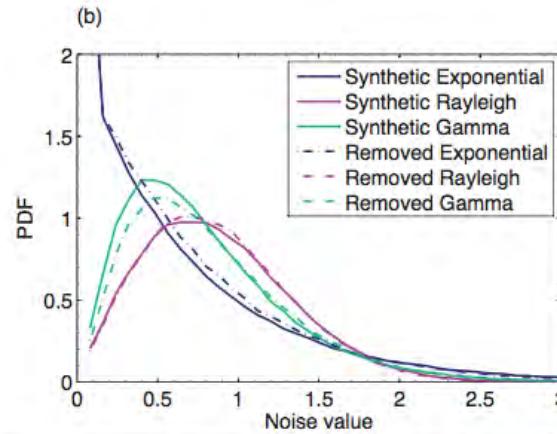
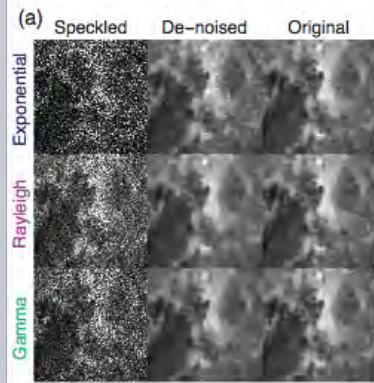
Origin of Dune Orientation on Titan

De-noising the SAR data



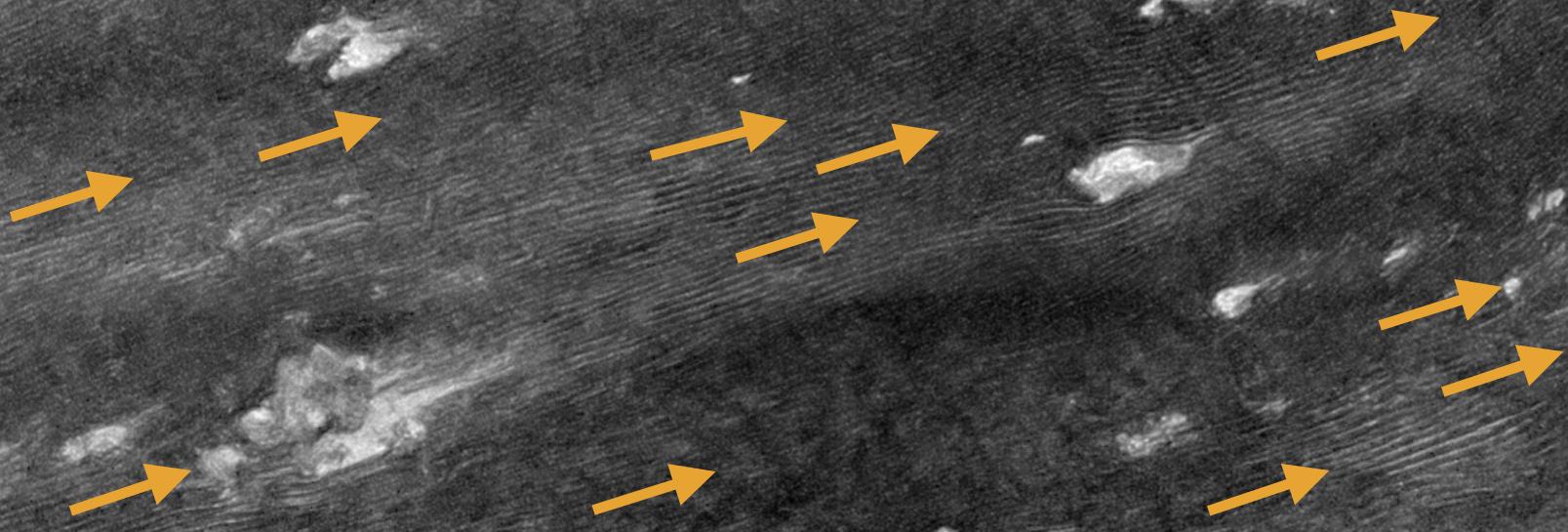
Lucas et al. (*in revision for JGR*)

Origin of Dune Orientation on Titan

Lucas et al. (*in revision for JGR*)

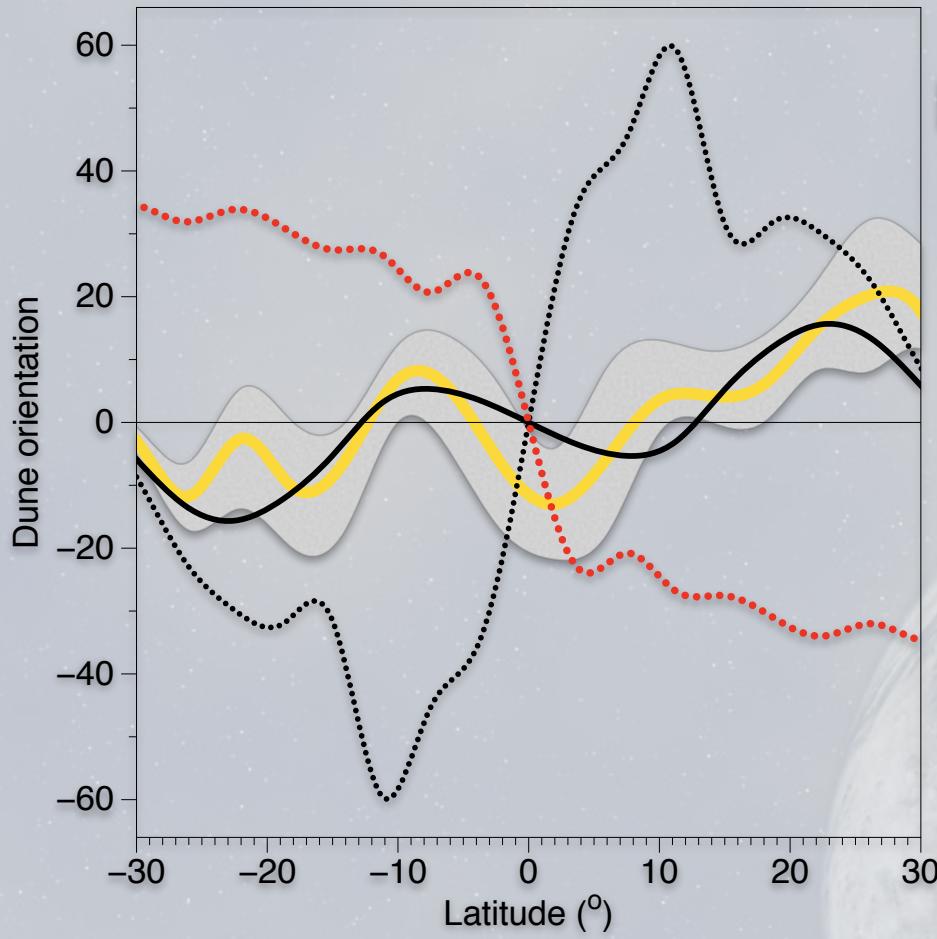


Dunes conundrum





Observed and Predicted Dune Orientation



Observed Dunes Orientation

Bed Instability Mode

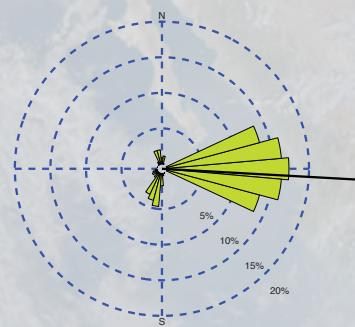
Fingering Mode

Equinoctial Storms:

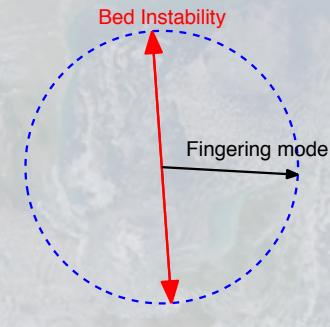
GCM from: Lebonnois & Charnay, 2009
Meso-scale in Charnay et al. under review

Fingering Mode with Storms

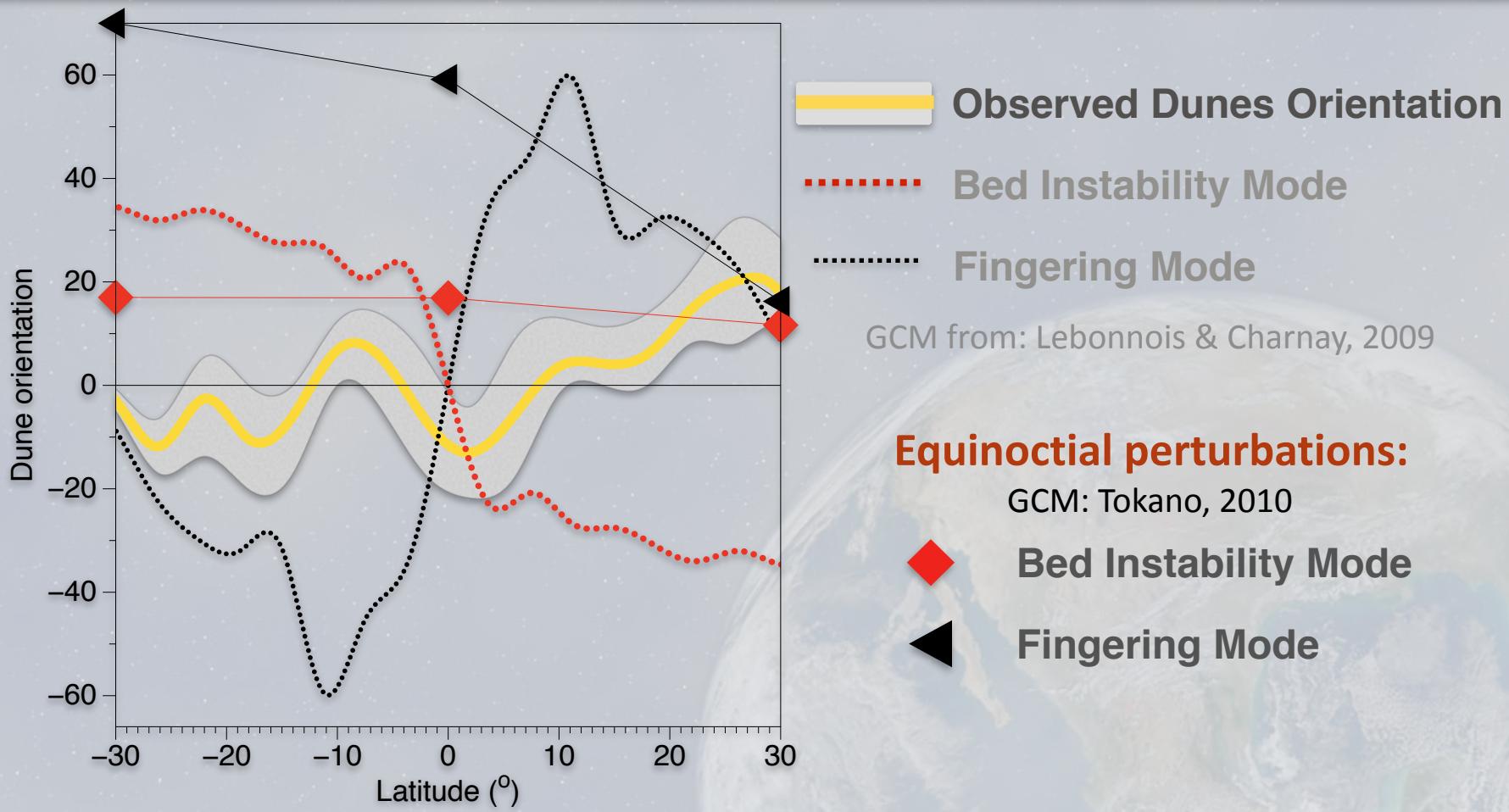
Sediment fluxes



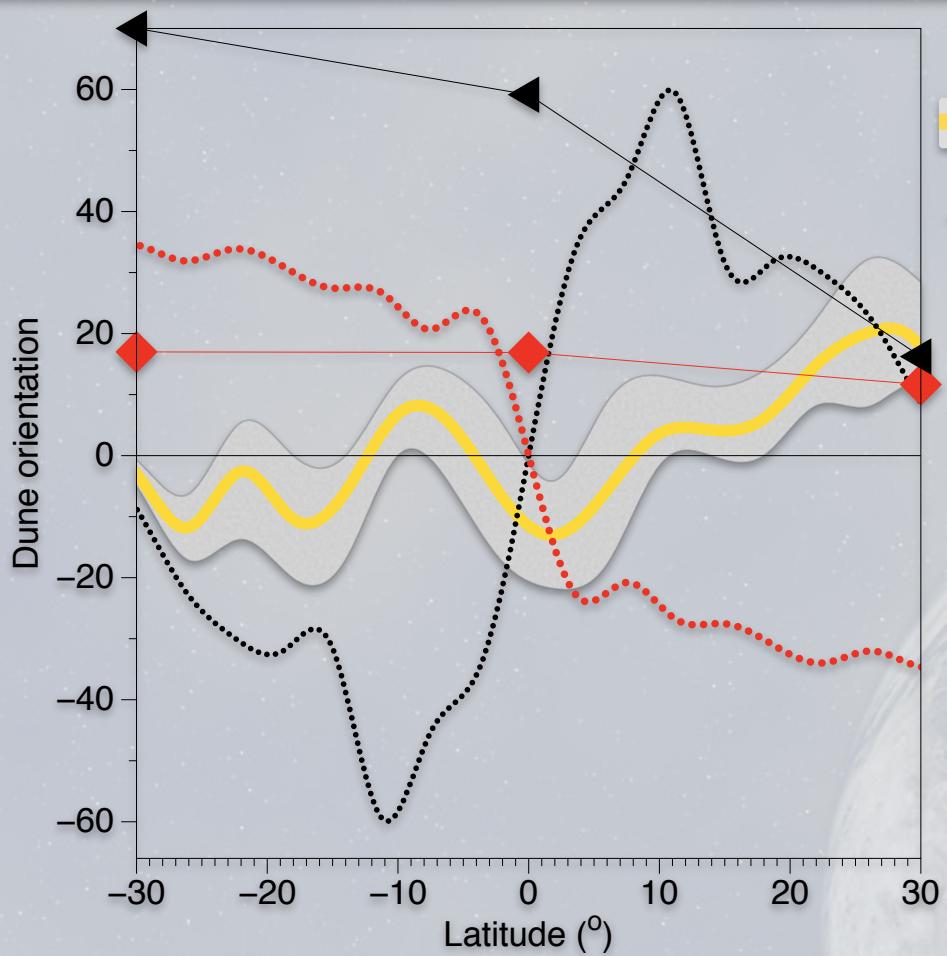
Dune orientation



Observed and Predicted Dune Orientation



Observed and Predicted Dune Orientation



Equinoctial perturbations:

GCM: Tokano, 2010



Bed Instability Mode

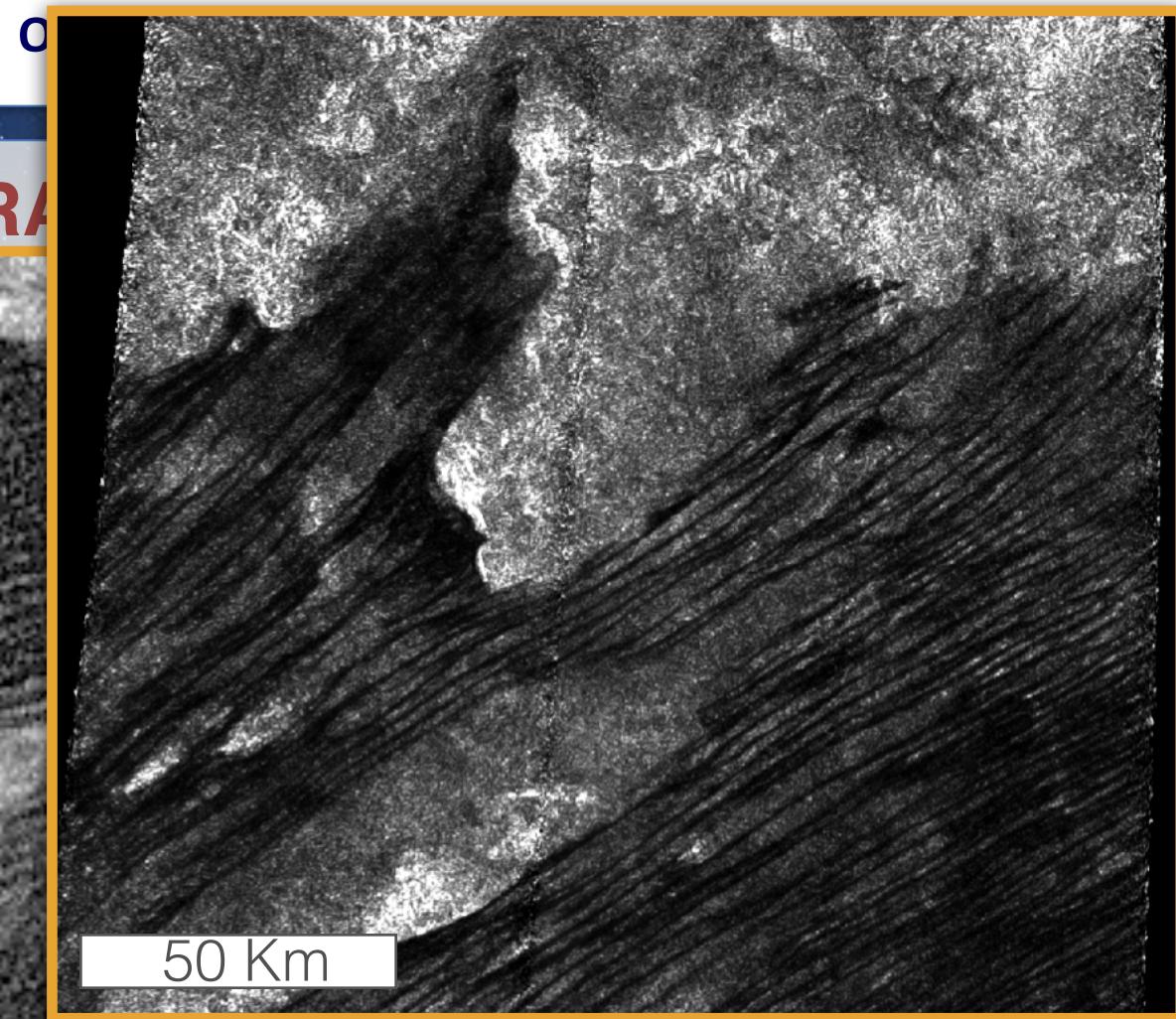


Fingering Mode

We need STILL an additional force leading to eastwards gusts



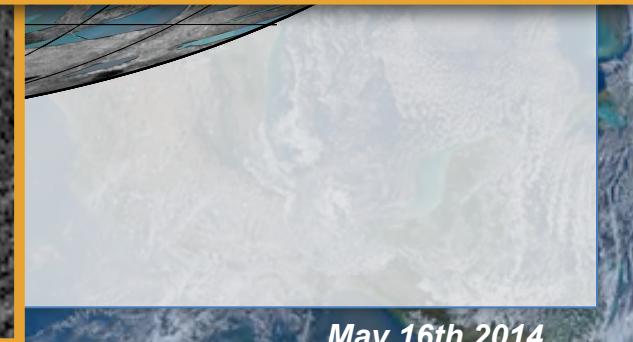
Titan seen by RADAR



50 Km



50 Km

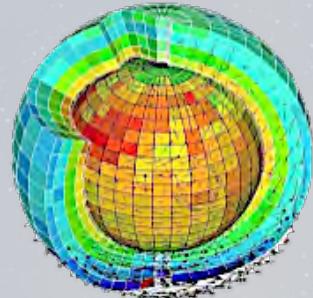


May 16th 2014

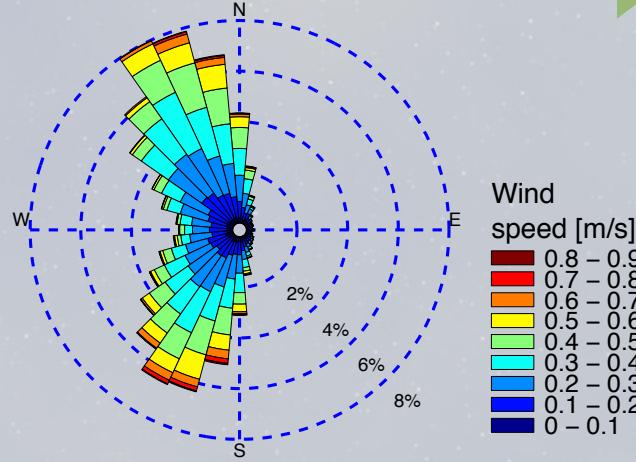
Prediction of Dune Orientation on Titan

Titan's GCM (IPSL-LMD)

Charnay et al., (2013)



Wind Regime

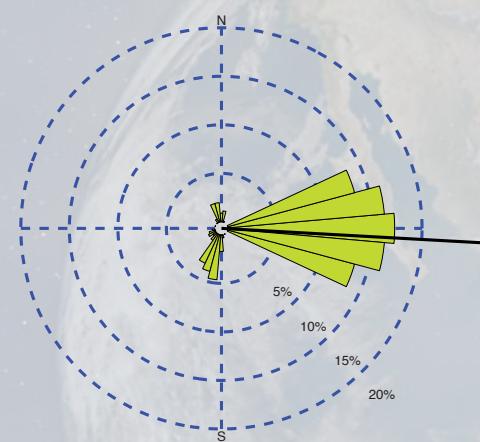


Sediment transport law

Lorenz et al., (1995)

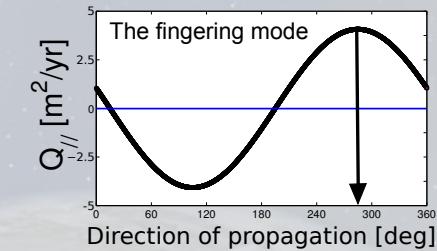
$$Q = f(u, u_{th}^*, d, \rho_s, \rho_f, g)$$

Sediment fluxes



Dune orientation

Courrech du Pont et al., (2014)



Dune orientation

