



Astroparticule et les Geosciences

Quelques pistes

October 26th to 31st

Labex **UnivEarthS**

FALL SCHOOL 2014

Villa Finaly, Florence (Italy)



FROM GAIÀ TO COSMOS
OBSERVING AND MODELLING

Stavros Katsanevas

APC, Astroparticle Physics and Cosmology



Deux mots de caution

Une entreprise interdisciplinaire doit assumer pendant un certain temps une précision moindre que celle revendiquée dans les branches canoniques du savoir. Parmi les particularités des sciences confortées dans leur conscience d'elles-mêmes, il y a le fait qu'elles croient toujours savoir avec bien plus de précision qu'on ne peut savoir. En aucun cas n'a-t-on obtenu le rétrécissement et l'isolement de l'objet d'étude sans contrepartie. Parce qu'elle n'accepte pas la délimitation bien définie et éprouvée de l'objet, l'entreprise interdisciplinaire doit avoir tout d'abord pour effet de décevoir

Hans Blumenberg

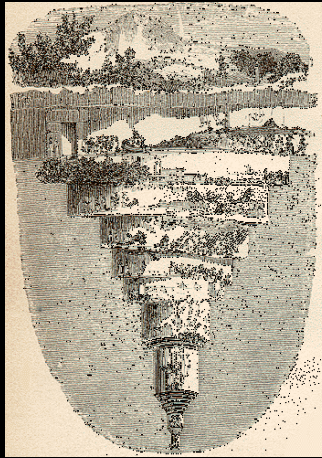
« Il y a deux types de physiciens qui mettent un pied en biologie. Les arrogants et les modestes. Les physiciens arrogants disent que la biologie a besoin de meilleurs concepts et que puisque les physiciens sont bons en concepts, leur tâche est de dire aux biologistes comment ils doivent penser. Les physiciens modestes disent que la biologie a besoin de meilleures machines et que comme les physiciens sont bons en machines, leur tâche est d'inventer des nouveaux outils que les biologistes pourront utiliser. Les physiciens qui inventent des nouveaux outils ont eu le plus grands impact. »

Freeman Dyson

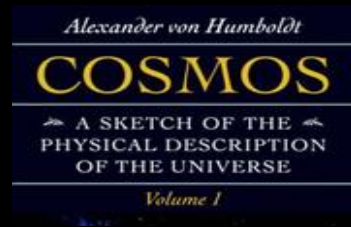
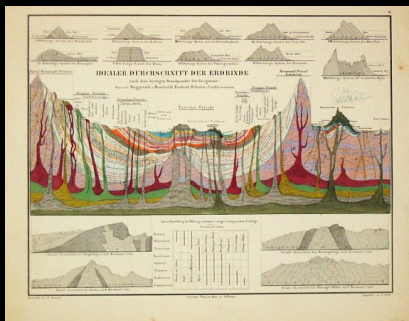


From Gaia to Cosmos

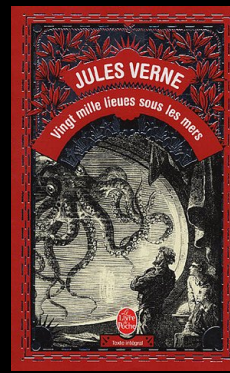
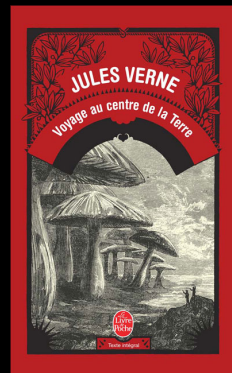
Illustrious predecessors



Galileo Galilei: Two Lectures to the Florentine Academy:
On the Shape, Location and Size of Dante's Inferno
(1588)
Use of scaling arguments



Alexander von Humboldt: Cosmos
(1845-1862)
“Unified” description of the Universe
Re-introduction of the ancient word



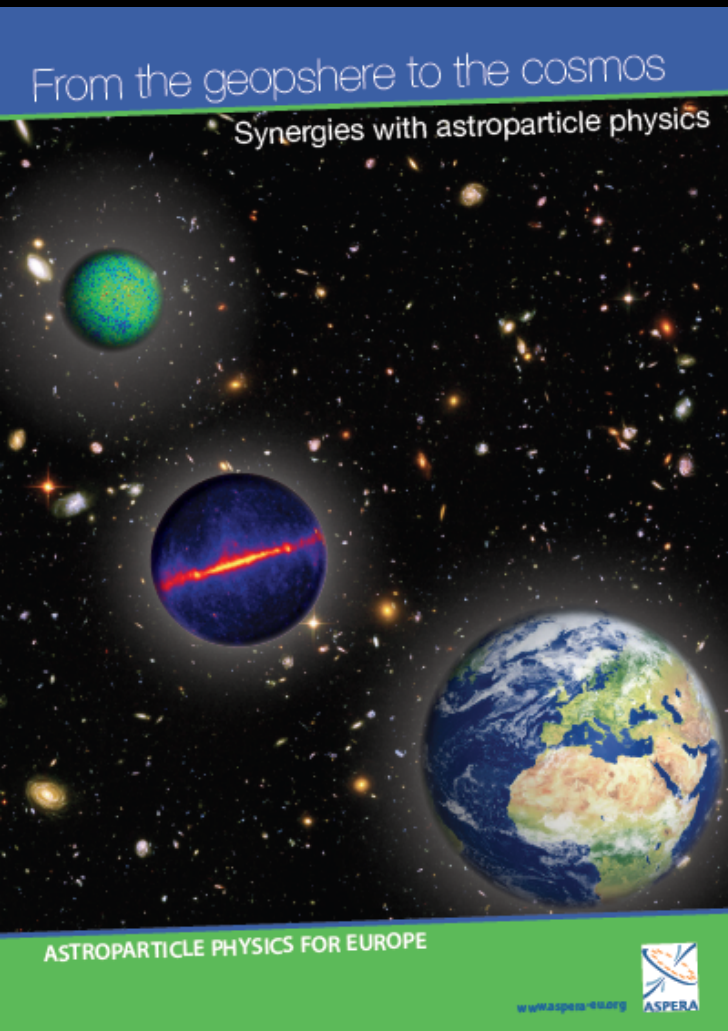
Jules Verne
Voyages Extraordinaires
(1862-1886)
I will use him as “cicerone”



A modern attempt From the Geosphere to the Cosmos

APPEC and ASPERA

<http://www.appec.org/multidisciplinarity.html>



3 workshops:

- Paris 2010 : From the Geosphere to the cosmos
- Amsterdam 2012: Deep Ocean Cabled Observatories
- Durham 2012: Underground synergies

Synergy Geosciences-astroparticle: a new frontier:

- For astroparticle the Geosphere is both the target and the detecting medium: continuous time series data by deploying large (autonomous?) sensor networks in hostile environments (sea, desert, underground)
- Dating is important: radioactivity-free platforms
- High sensitivity instruments for probing very low intensity geological effects, metrology
- Large data manipulation and worldwide networking. Distribution of alerts



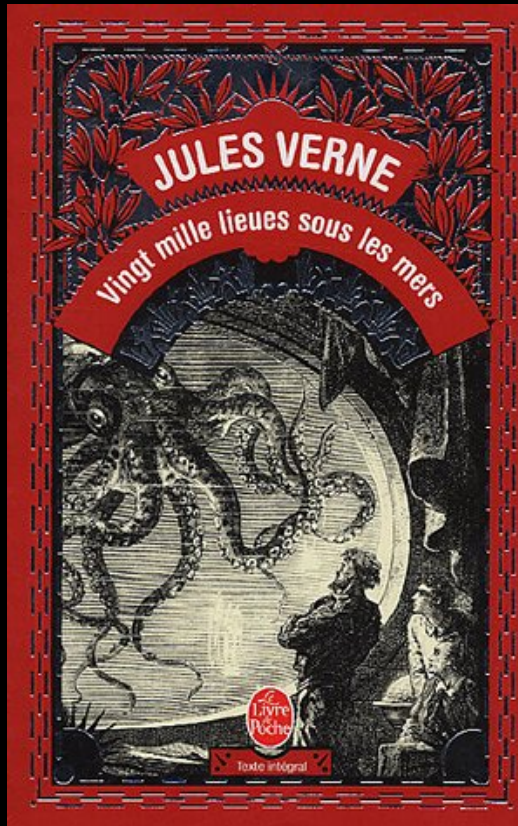
The 2010 census: close to 40 points of synergy

| | ATMOSPHERIC AND TERRESTRIAL | UNDERGROUND | UNDERWATER |
|------------------------------|---|--|---|
| UNDERSTANDING THE ATMOSPHERE | <ul style="list-style-type: none"> 1. SPACE WEATHER (Section 3.1.1) 2. ATMOSPHERIC MONITORING (Section 3.1.2) 3. COSMOCLIMATOLOGY (Section 3.1.3) 4. THUNDERSTORMS and LIGHTNINGS (Section 3.1.5) | <ul style="list-style-type: none"> 1. COSMOCLIMATOLOGY (Section 3.1.3) | <ul style="list-style-type: none"> 1. ATMOSPHERIC TEMPERATURE VARIATION (Section 3.1.4) |
| UNDERSTANDING THE EARTH | <ul style="list-style-type: none"> 1. EROSION RATE CALCULATION (Section 3.2.1) 2. VOLCANO TOMOGRAPHY (Section 3.2.5) | <ul style="list-style-type: none"> 1. COASTAL ROCK CLIFF EROSION (Section 3.2.2) 2. CHRONOLOGY for the PALEOENVIRONMENT (Section 3.2.3) 3. EARTH's INTERIOR - GEONEUTRINOS (Section 3.2.6) | <ul style="list-style-type: none"> 1. PALEOCLIMATE (Section 3.2.4) 2. EARTH RADIOGRAPHY (Section 3.2.7) |
| UNDERSTANDING THE OCEANS | | <ul style="list-style-type: none"> 1. CORAL CHRONOLOGY (Section 3.3.6) | <ul style="list-style-type: none"> 1. CONTINUOUS OCEANOGRAPHIC DATA (Section 3.3.1) 2. SEDIMENT TRANSPORT (Section 3.3.2) 3. OXYGEN DYNAMICS (Section 3.3.3) 4. RADIOACTIVITY (Section 3.3.4) 5. INTERNAL WAVES (Section 3.3.5) |
| UNDERSTANDING EARTHQUAKES | <ul style="list-style-type: none"> 1. EARTHQUAKE MONITORING GRID (Section 3.4.1) | <ul style="list-style-type: none"> 1. SEISMO-ELECTROMAGNETIC COUPLINGS (Section 3.4.4) 2. EARTHQUAKE PRECURSORS (Section 3.4.5) 3. SLOW EARTHQUAKE MONITORING (Section 3.4.6) | <ul style="list-style-type: none"> 1. EARTHQUAKE AND TSUNAMI MONITORING (Section 3.4.2) 2. STUDYING THE LAKE ENVIRONMENT (Section 3.4.3) |
| UNDERSTANDING BIODIVERSITY | | <ul style="list-style-type: none"> 1. IMPACT OF RADIATION (Section 3.5.2) 2. EXTREMOPHILES (Section 3.5.8) | <ul style="list-style-type: none"> 1. UNDERWATER SOUND MONITORING (Section 3.5.1) 2. DEEP SEA BIOLUMINESCENCE (Section 3.5.2) 3. BIODIVERSITY UNDER ICE (Section 3.5.3) 4. BIODEGRADATION (Section 3.5.4) 5. MICROBIOLOGY (Section 3.5.5) 6. BIOFOULING (Section 3.5.6) |
| APPLICATIONS | | <ul style="list-style-type: none"> 1. WINE DATATION (Section 3.6.1) 2. SALT CHARACTERISATION AOC (Section 3.6.2) 3. SOFT ERROR RATE IN ELECTRONICS (Section 3.6.3) 4. ROCK DEFORMATION (Section 3.6.4) | |

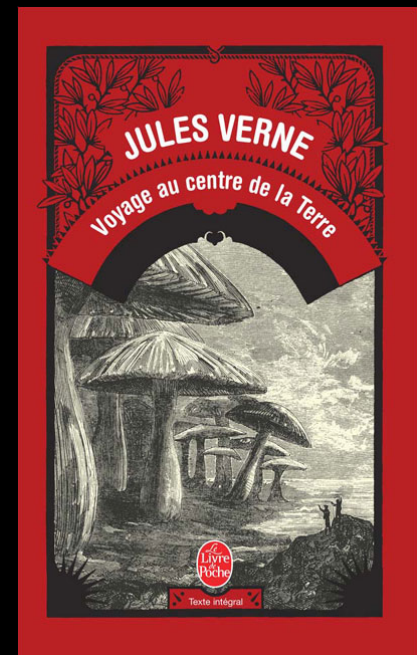


Outline

- Cosmic rays and the atmosphere
 - Atmospheric electricity
 - Meteorites
 - Activating the earth elements
 - Tomography
 - Gravitational wave antenna apps(see JPM)



- Underwater/Under-ice
 - Underwater instrumentation
 - Seismology
 - Oceanography
 - Biodiversity (coupling of life to matter)
 - Probing the earth core (see AT)
- Underground
 - Dating applications
 - Geoneutrinos (see AT)
 - Future





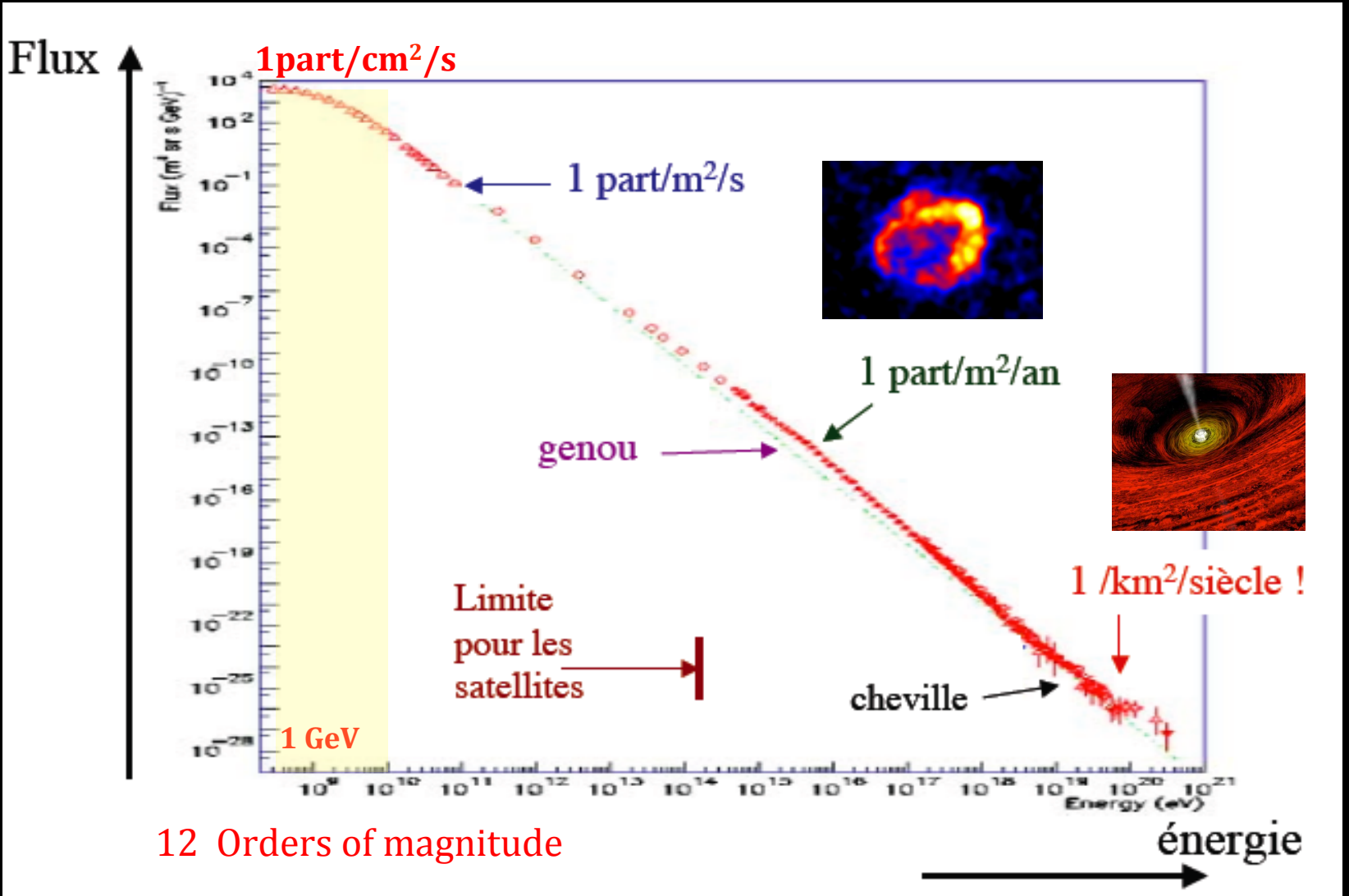
Atmospheric synergies



ASTRI E PARTICELLE - LE PAROLE DELL'UNIVERSO

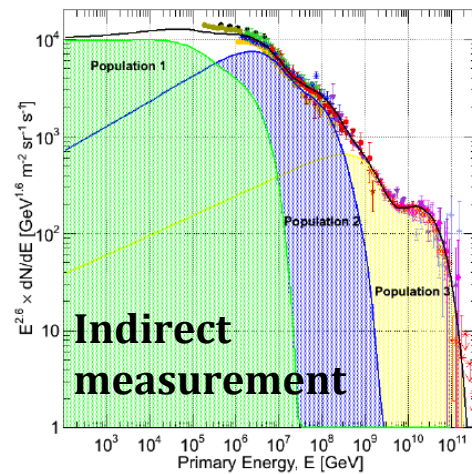
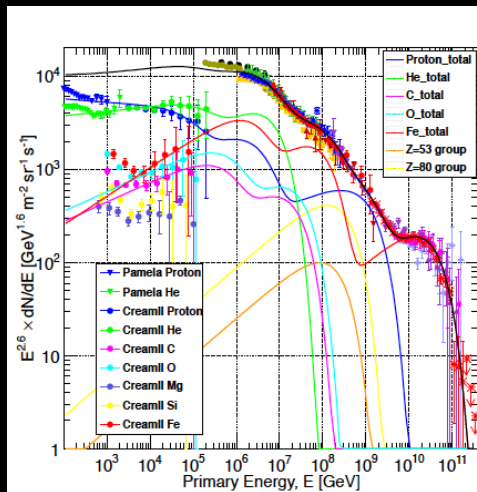
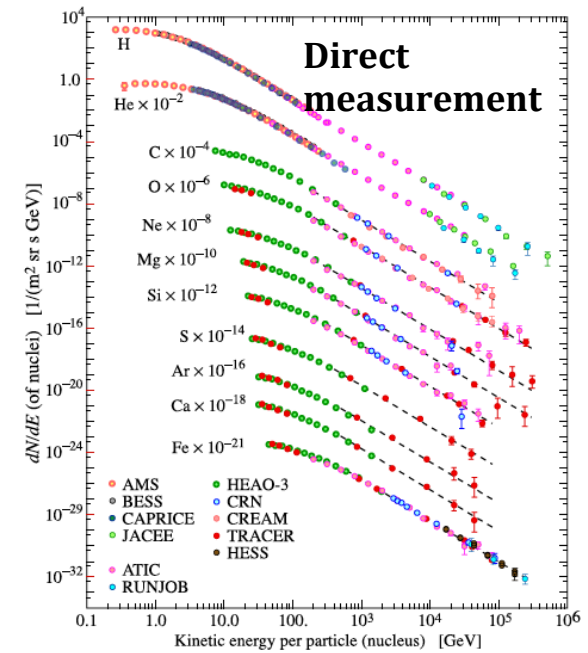
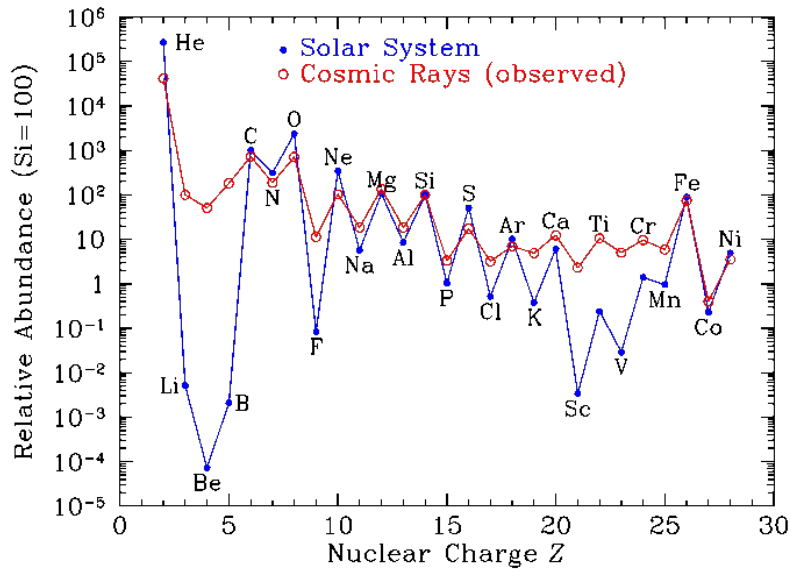


Fluxes



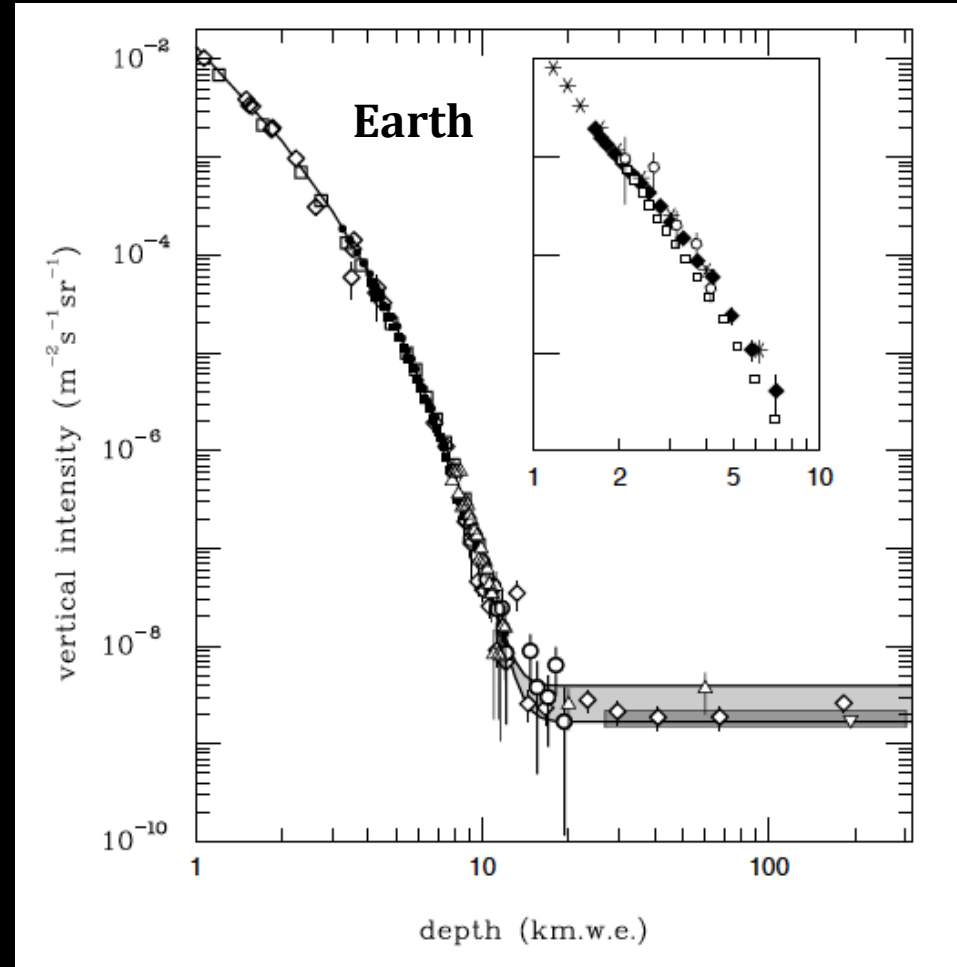
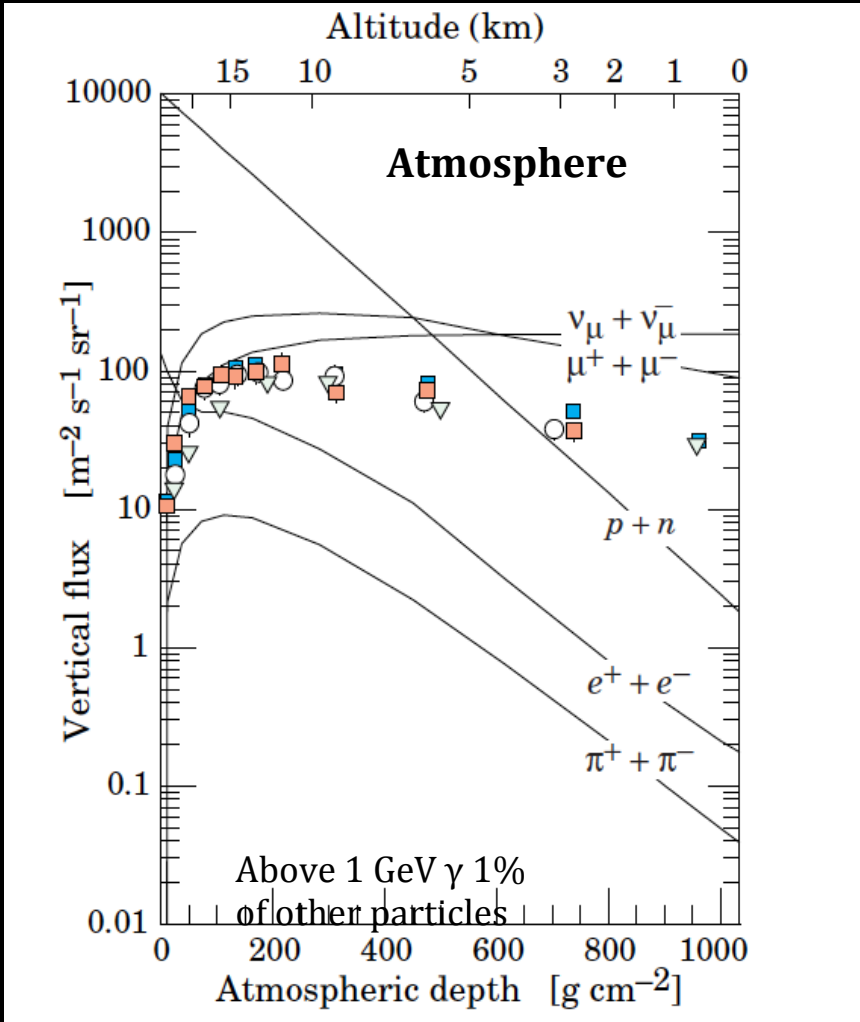


Composition of cosmic rays (nuclei)





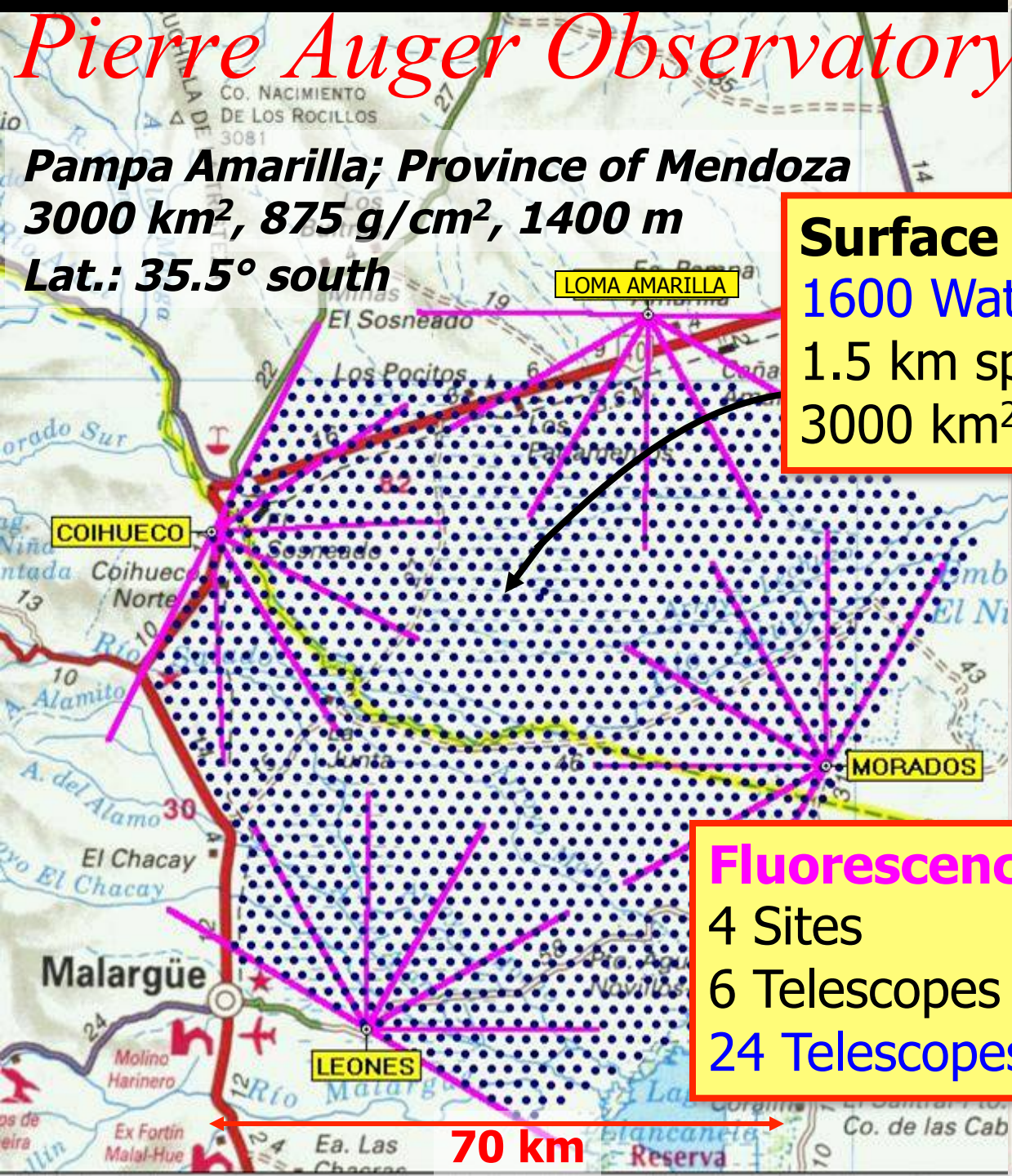
Composition of cosmic rays (particles)



Vertical muon intensity vs depth
(1km w.e = 10^5 g/cm^2 of standard rock)

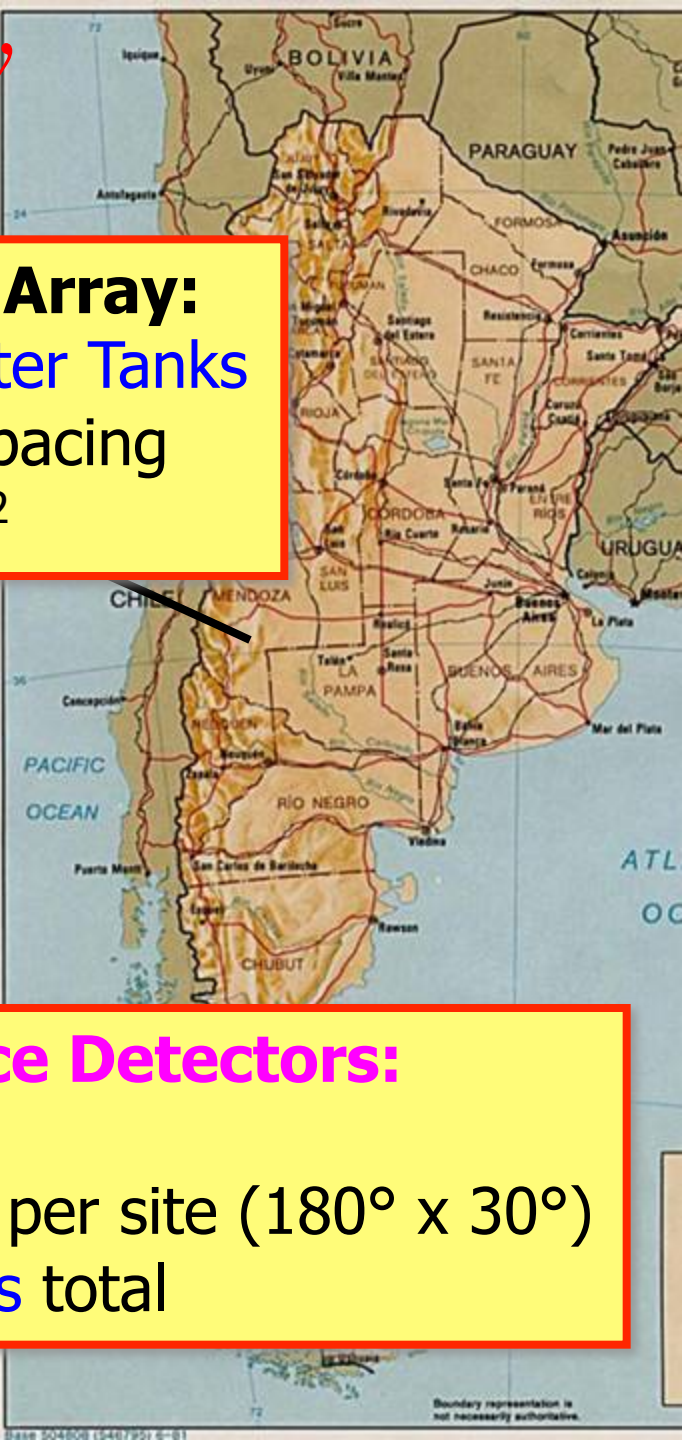
Pierre Auger Observatory

Pampa Amarilla; Province of Mendoza
3000 km², 875 g/cm², 1400 m
Lat.: 35.5° south



Surface Array:
1600 Water Tanks
1.5 km spacing
3000 km²

Fluorescence Detectors:
4 Sites
6 Telescopes per site (180° x 30°)
24 Telescopes total



Water Tank in the Pampa

Communication antenna

GPS antenna

Electronics enclosure
40 MHz FADC, local triggers, 10 Watts

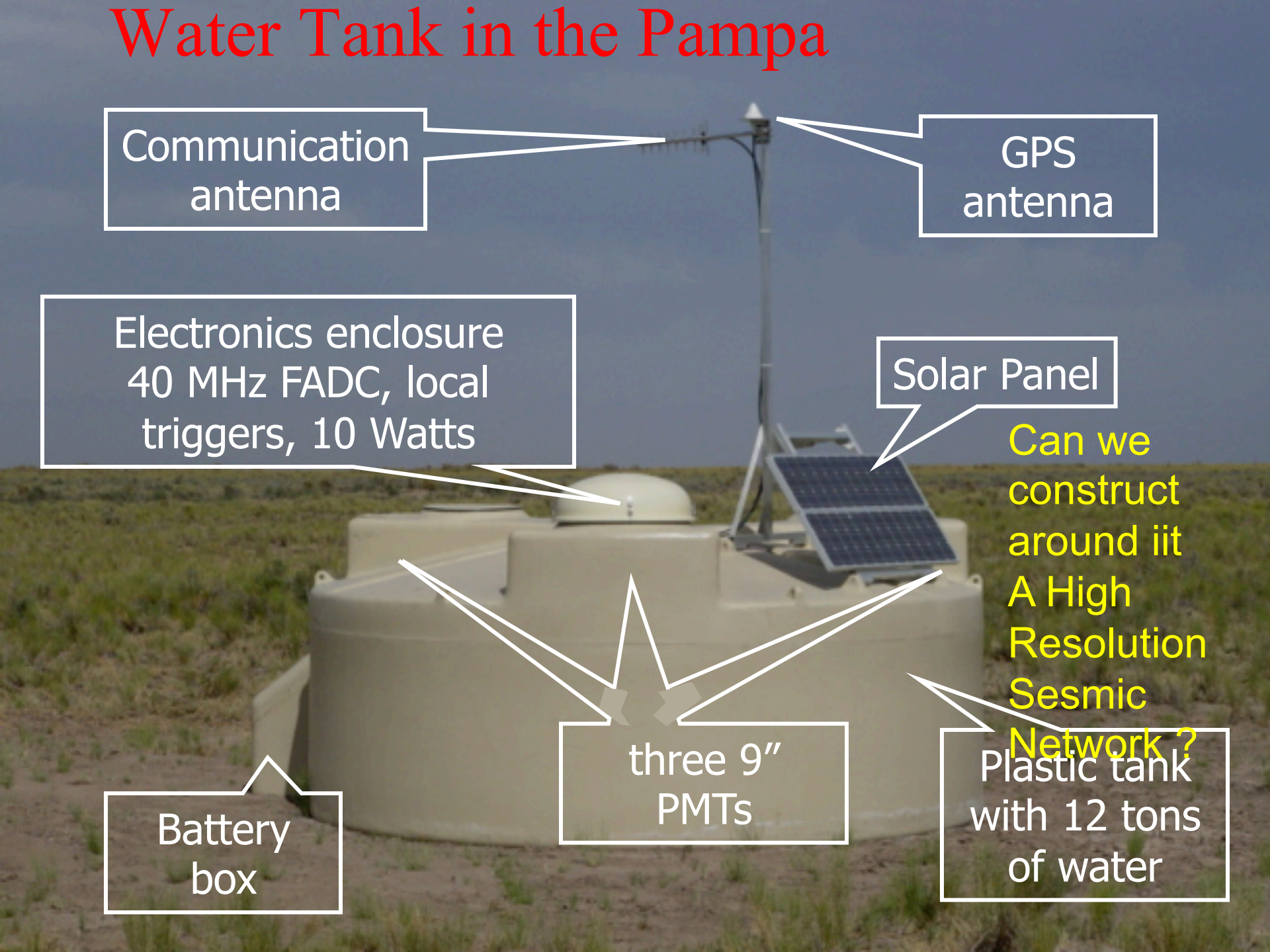
Solar Panel

Can we construct around it
A High Resolution
Sesmic
Network ?

Battery box

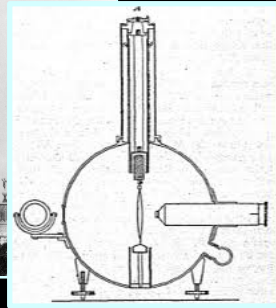
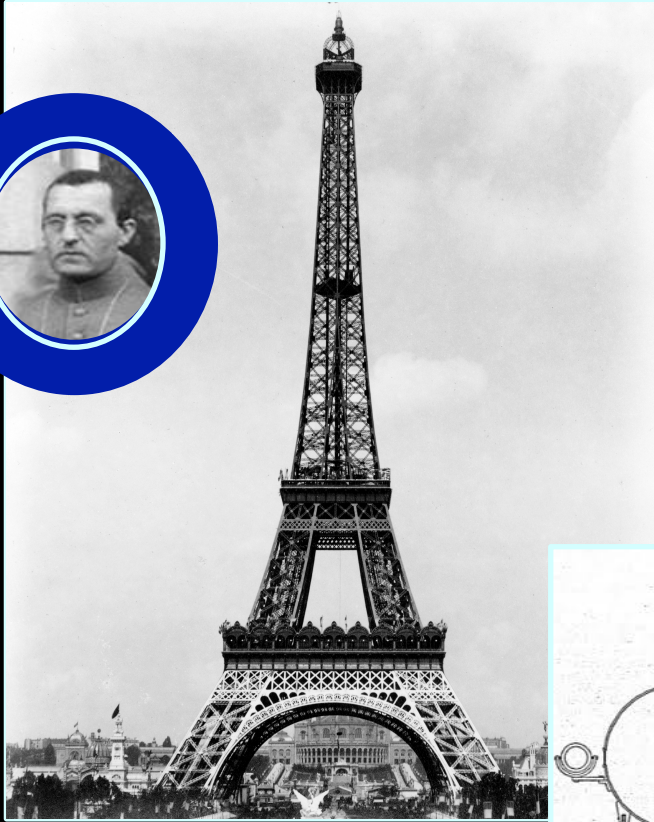
three 9" PMTs

Plastic tank with 12 tons of water



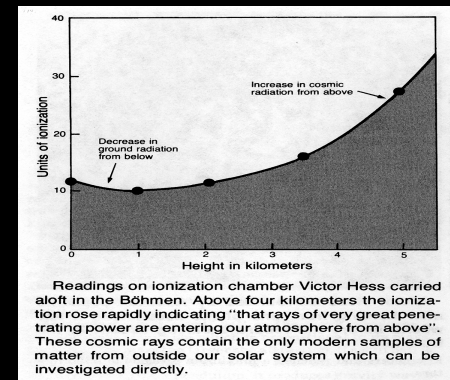


Ionisation of the atmosphere (CRT Wilson) a centenary problem coupling cosmic rays and atmospheric science



6am August 7, 1912
Aussig, Austria

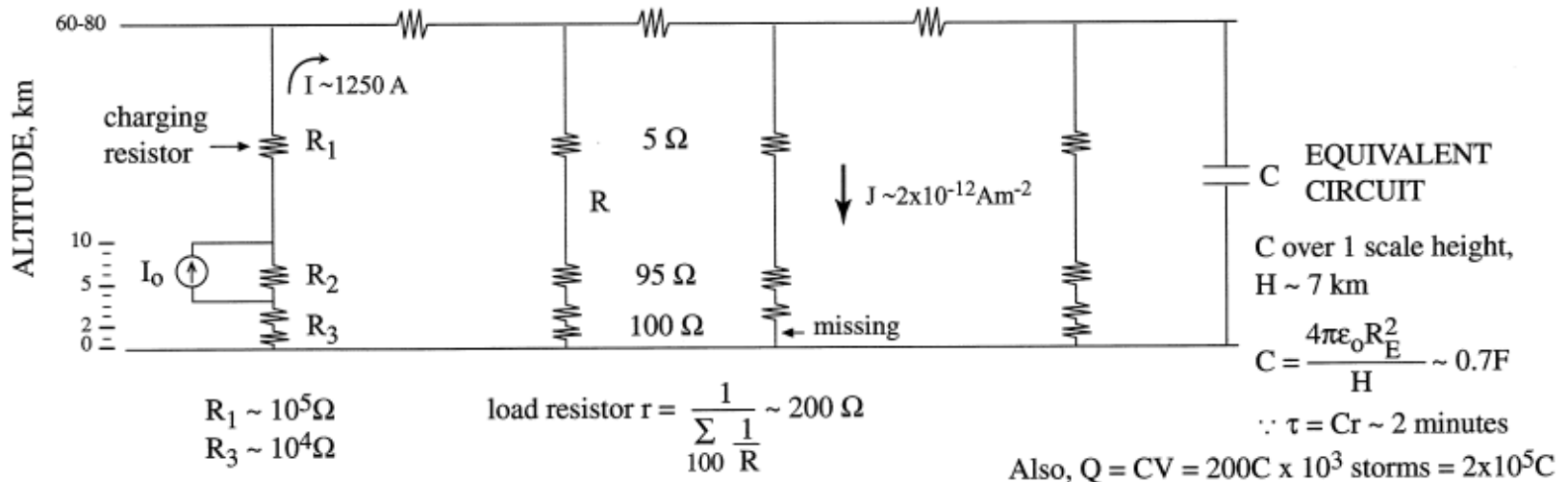
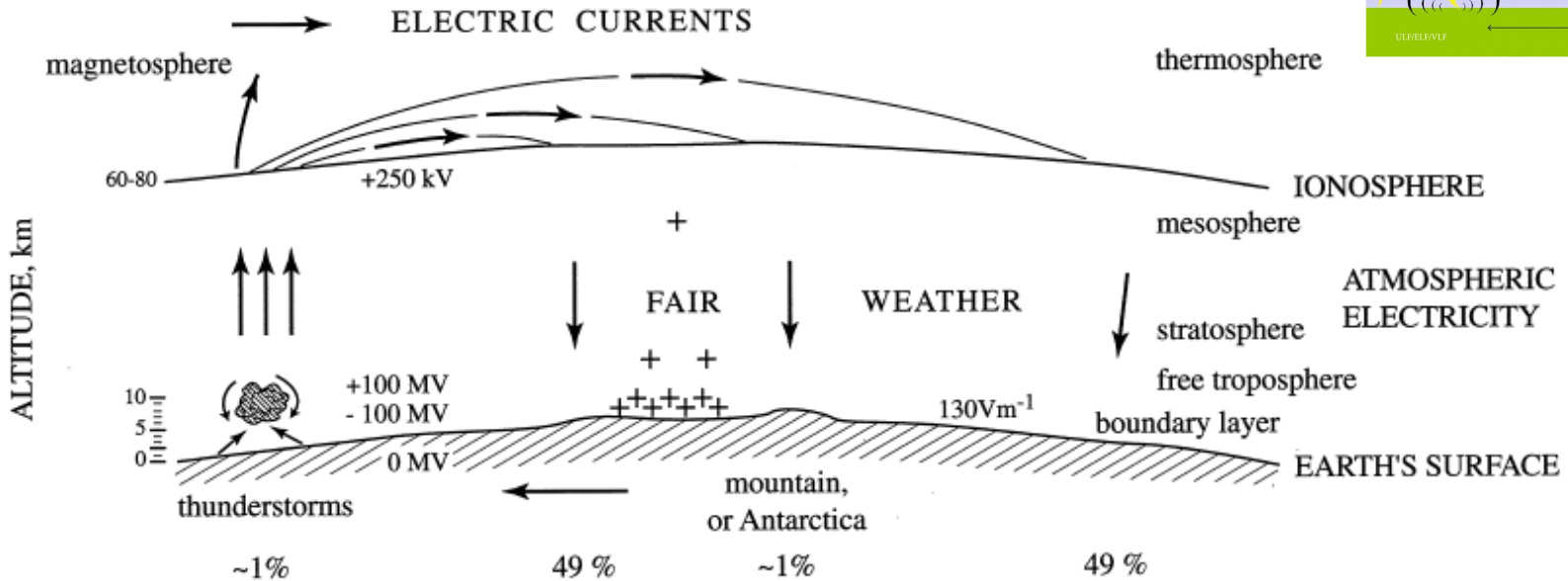
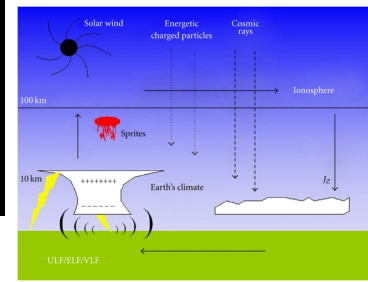
First measures T. Wulff
1909-1910 Tour Eiffel



Effectivement ces particules viennent du cosmos,



The atmosphere as an electric circuit



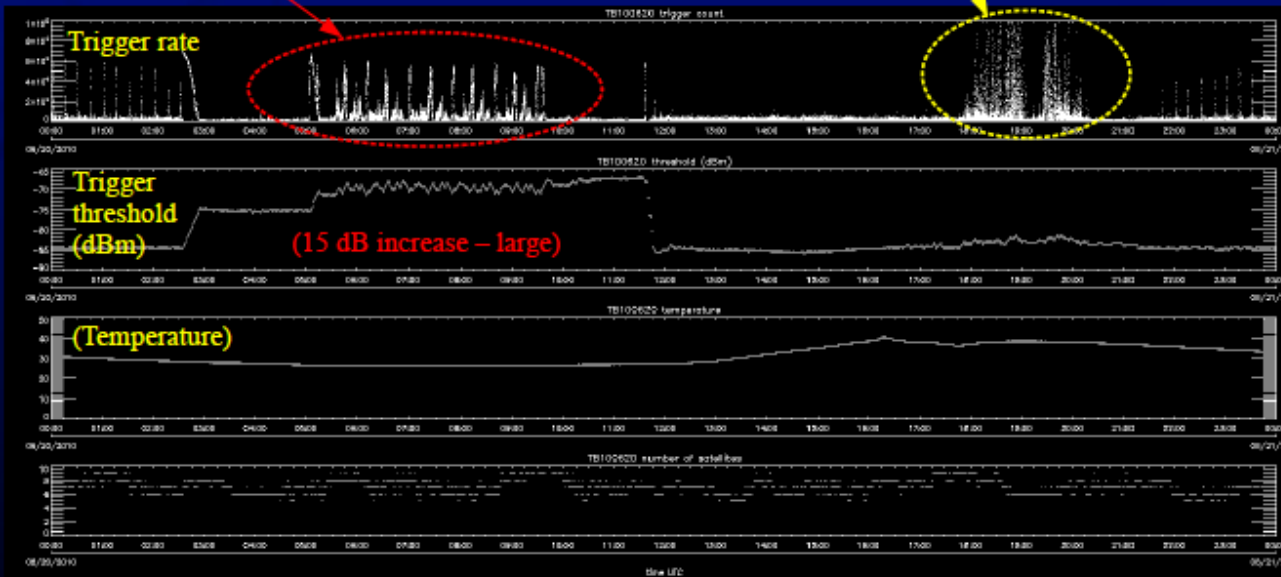


Lightnings and High Energy Cosmic Rays

Raw data from station B:
Man-made and lightning- produced signals

Local radio frequency
interference (RFI) from
Fluorescence Telescopes

Lightning signals from
somewhere over/around
Pierre Auger

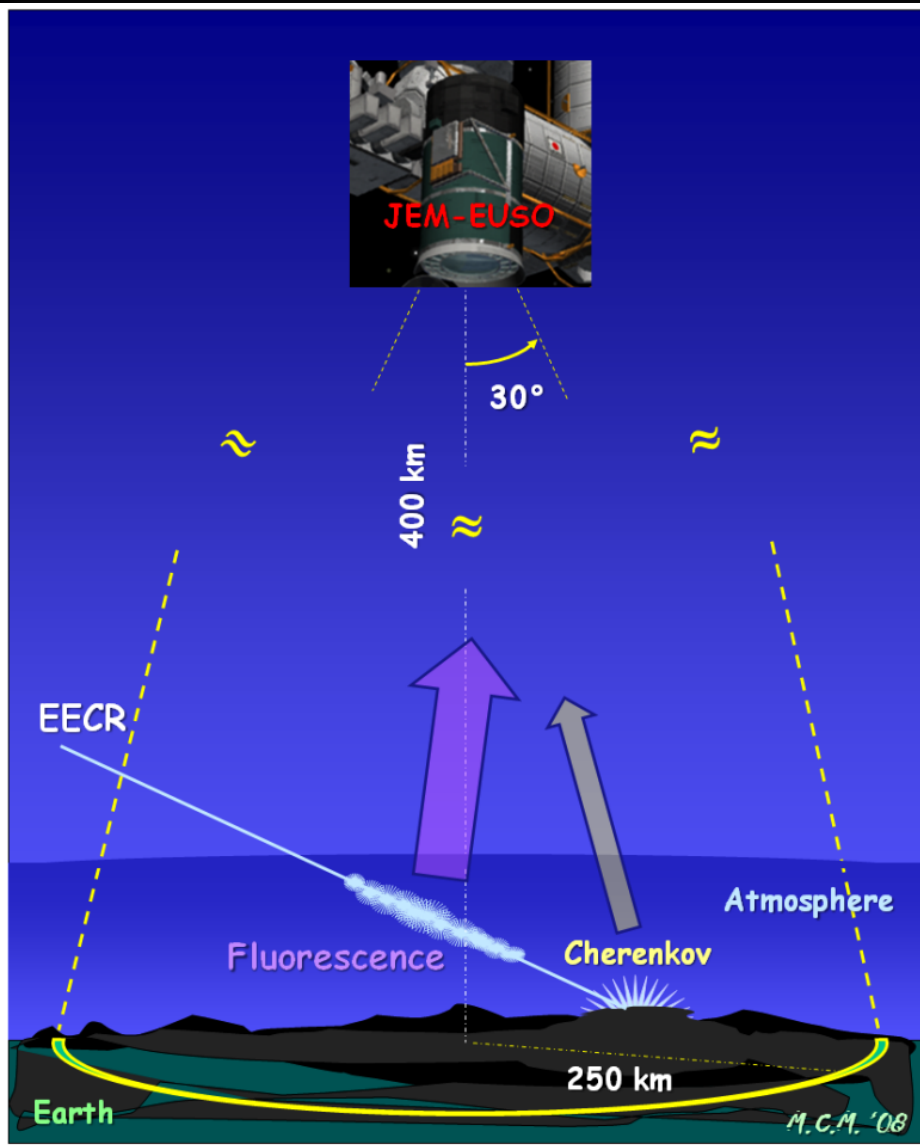


Can distinguish between lightning and local RFI noise by
correlating signals from the two LMA stations

- Expect a couple of geolocated ultra high energy CRs through electrically active storms per year. If individual CRs have an effect on storms these should.



See the atmosphere from above



Énorme champ de vue
(60° d'ouverture)

Caméra ultra-sensible dans
l'UV (photon unique)

Caméra ultra-rapide (2.5 μ s)

300 000 pixels

+ Caméra Infraouge

+ LIDAR

+ High power Laser



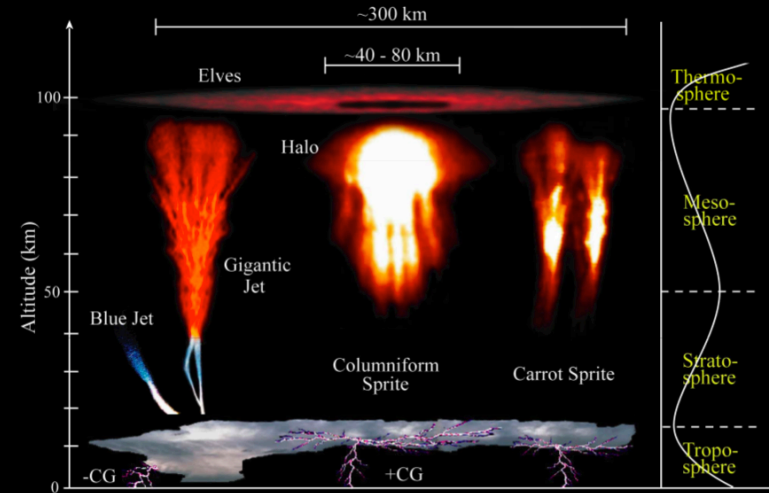
Possible synergies JEM-EUSO I

Atmospheric science

TLEs (sprites, jets, elves, halos...)

NEW : ionospheric airglow:

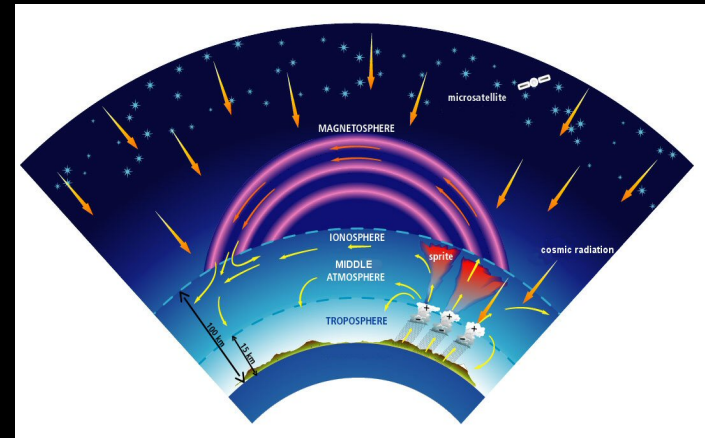
Transient Luminous Events (TLEs)



Imaging and modeling the ionospheric airglow response over Hawaii to the tsunami generated by the Tohoku earthquake of 11 March 2011

J. J. Makela,¹ P. Lognonné,² H. Hébert,³ T. Gehrels,¹ L. Rolland,² S. Allgeyer,³ A. Kherani,⁴ G. Occhipinti,² E. Astafyeva,² P. Coïsson,² A. Loevenbruck,³ E. Clévéde,² M. C. Kelley,⁵ and J. Lamouroux⁶

Received 19 April 2011; revised 25 May 2011; accepted 25 May 2011; published 7 July 2011.





100 μ s

alt. [km]

100

90

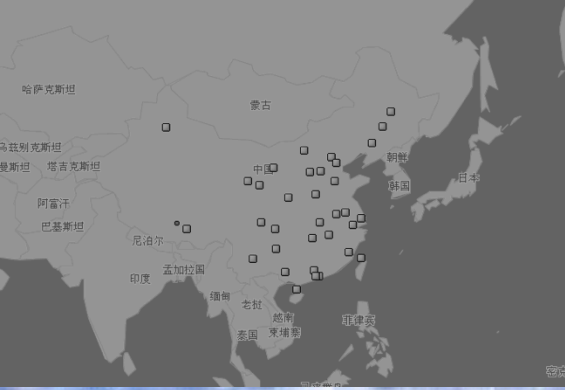
80

70

9 Jul 2005, 0438:00 UT
10,000 fps, 50 μ s exp.

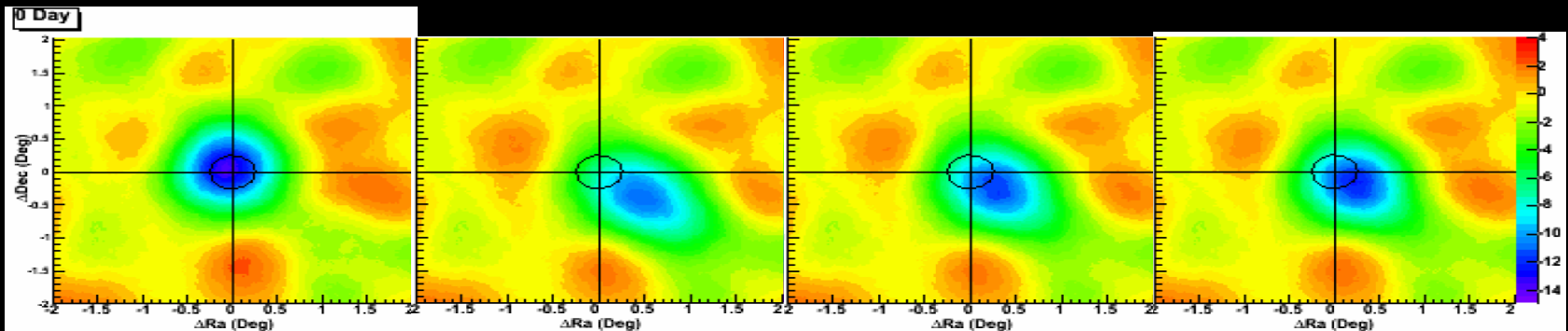
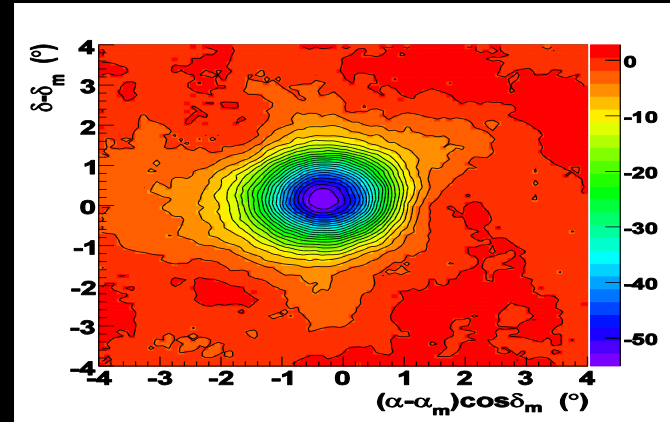
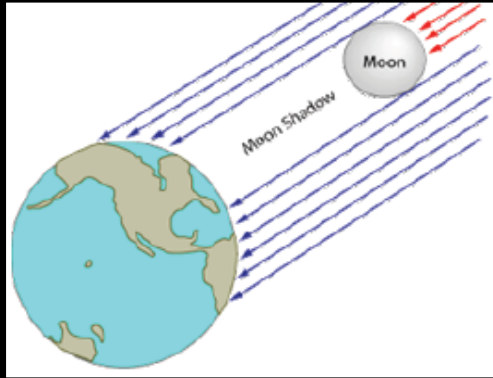
USAFA / UAFGI
NMT Langmuir Lab.

Yangbajing Valley & CR/ γ Observatory



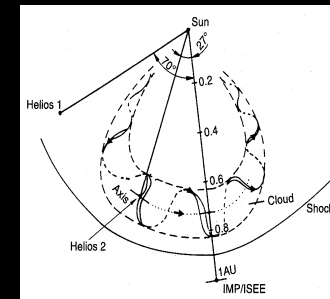
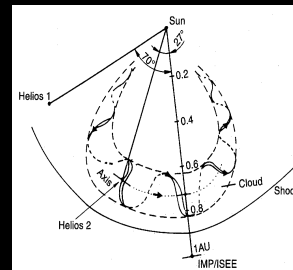
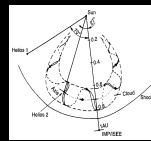


Predicting Interplanetary Coronal Mass Ejection by using the sun's shadow



An ICME of 4 days

Need to detect sun's shadow in 1 day:
LHASSO





Possible synergies JEM-EUSO II

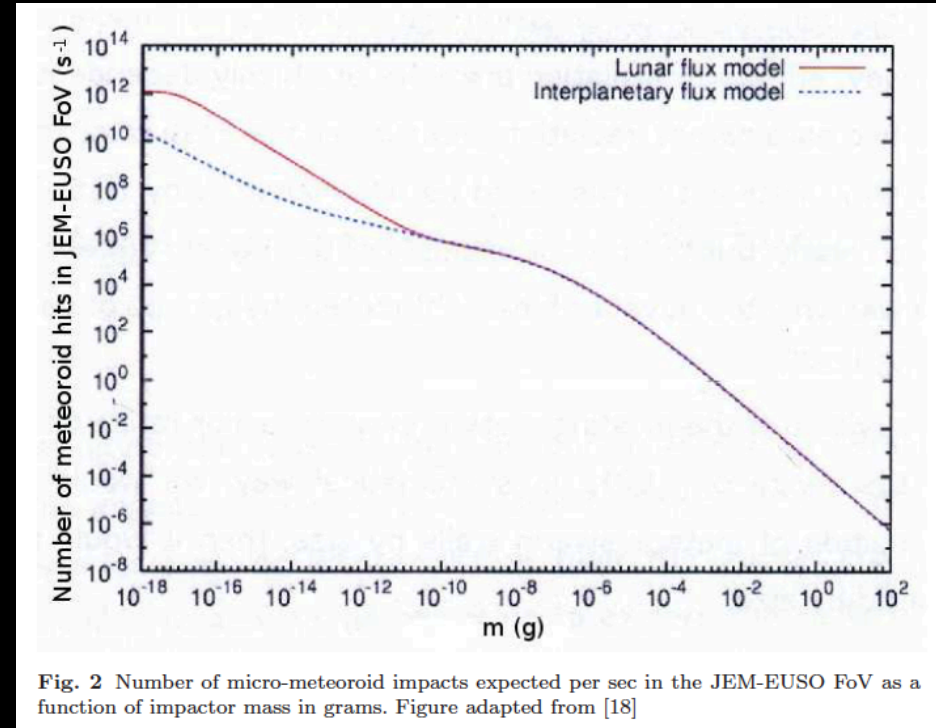
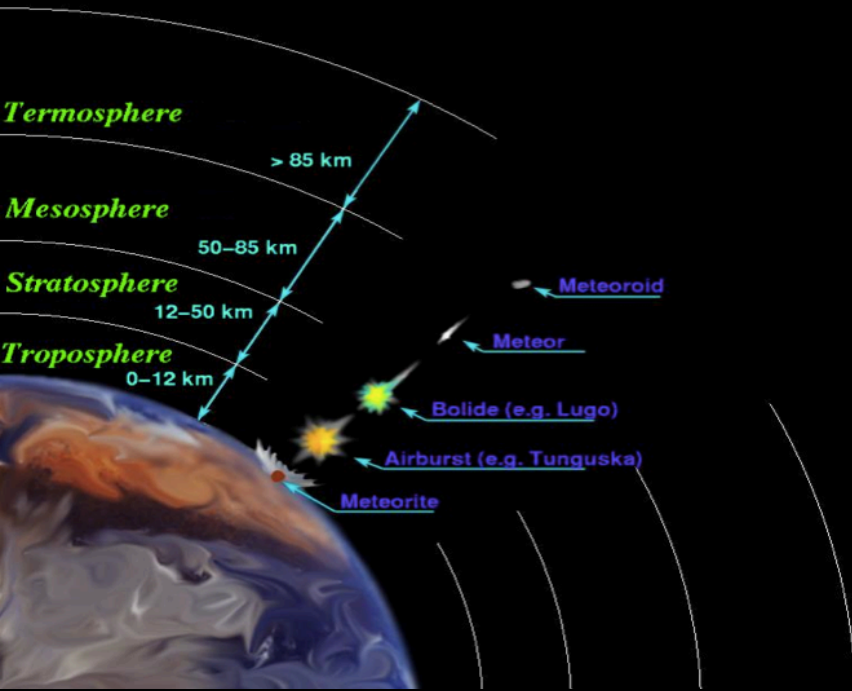


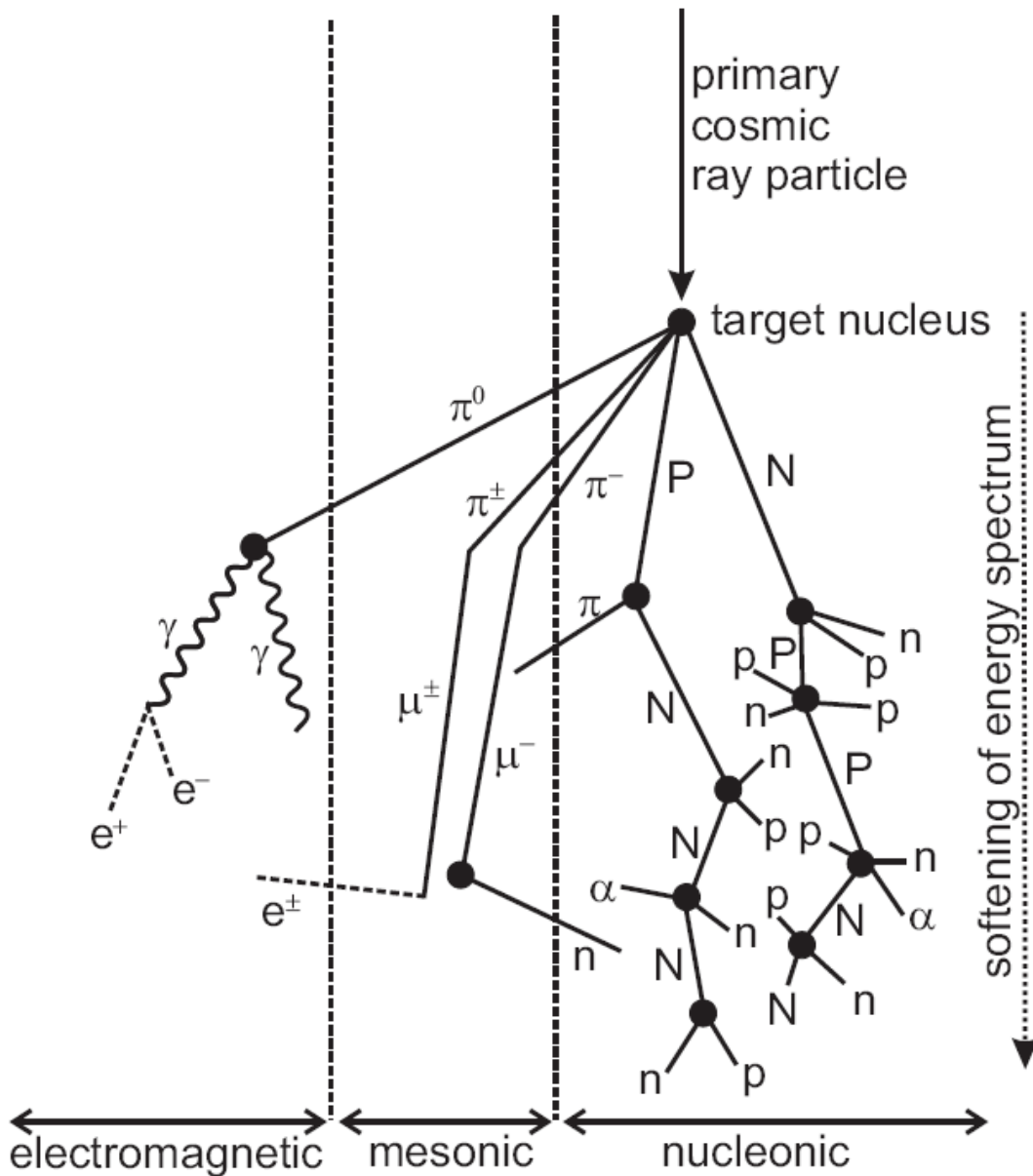
Fig. 2 Number of micro-meteoroid impacts expected per sec in the JEM-EUSO FoV as a function of impactor mass in grams. Figure adapted from [18]

Statistiques

Pour les météorites : détermination la trajectoire et la zone d'impact.

→ récupérer avant la « pollution » par l'environnement terrestre

NEW: Associé à un laser puissant nettoyer les débris spatiaux



Transmutations in the atmosphere: The nuclear cascade

- The secondary cosmic ray flux decreases exponentially with atmospheric depth
- At the Earth's surface it are the neutrons that remain (~90 % of the nucleon flux)
- They activate nuclei up to approximately 3m of depth

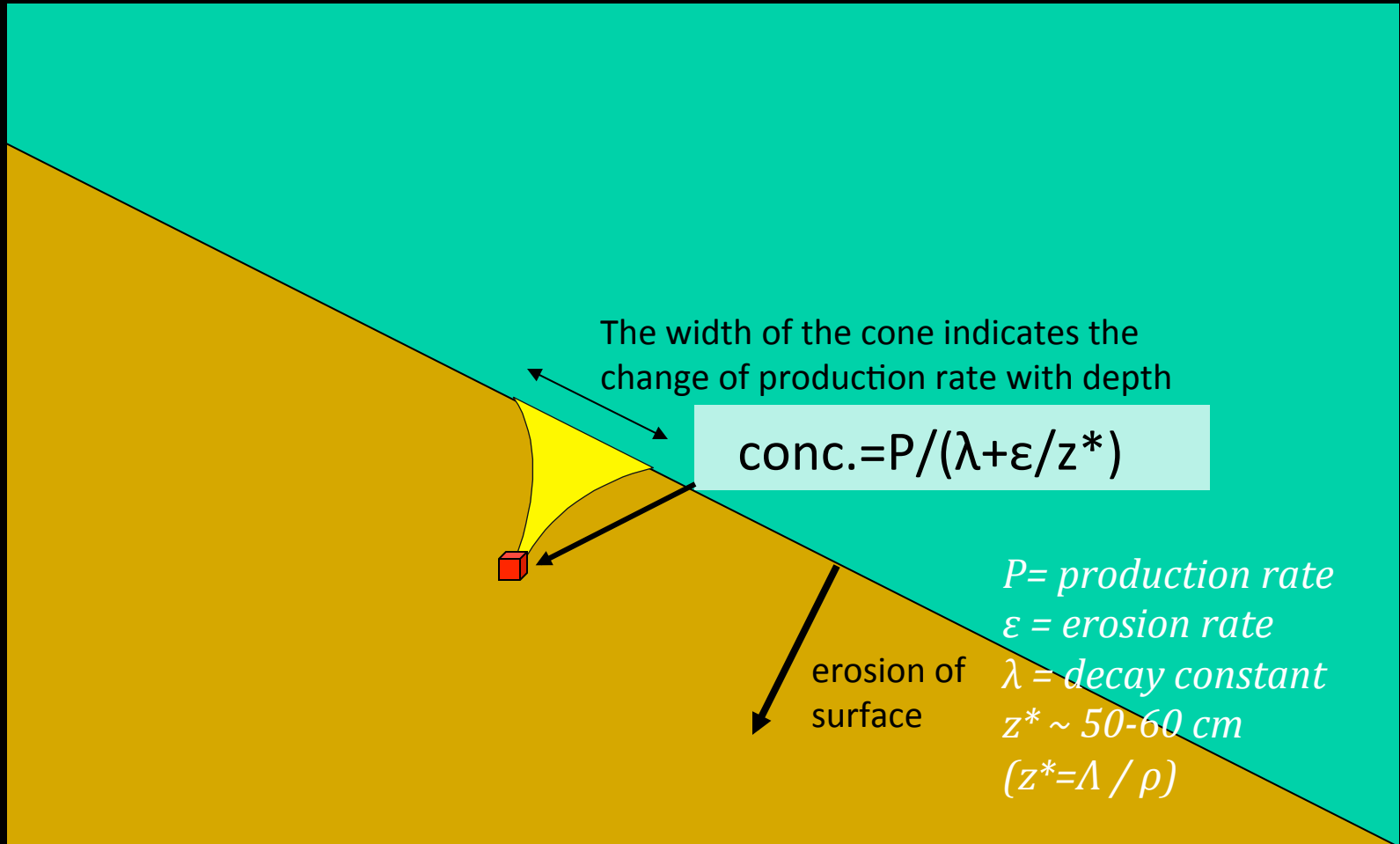
Nuclides used:

^3He , ^{10}Be , ^{21}Ne , ^{26}Al , ^{36}Cl

➔ Measurement of erosion



The concentration of cosmogenic nuclides is a function of the erosion rate



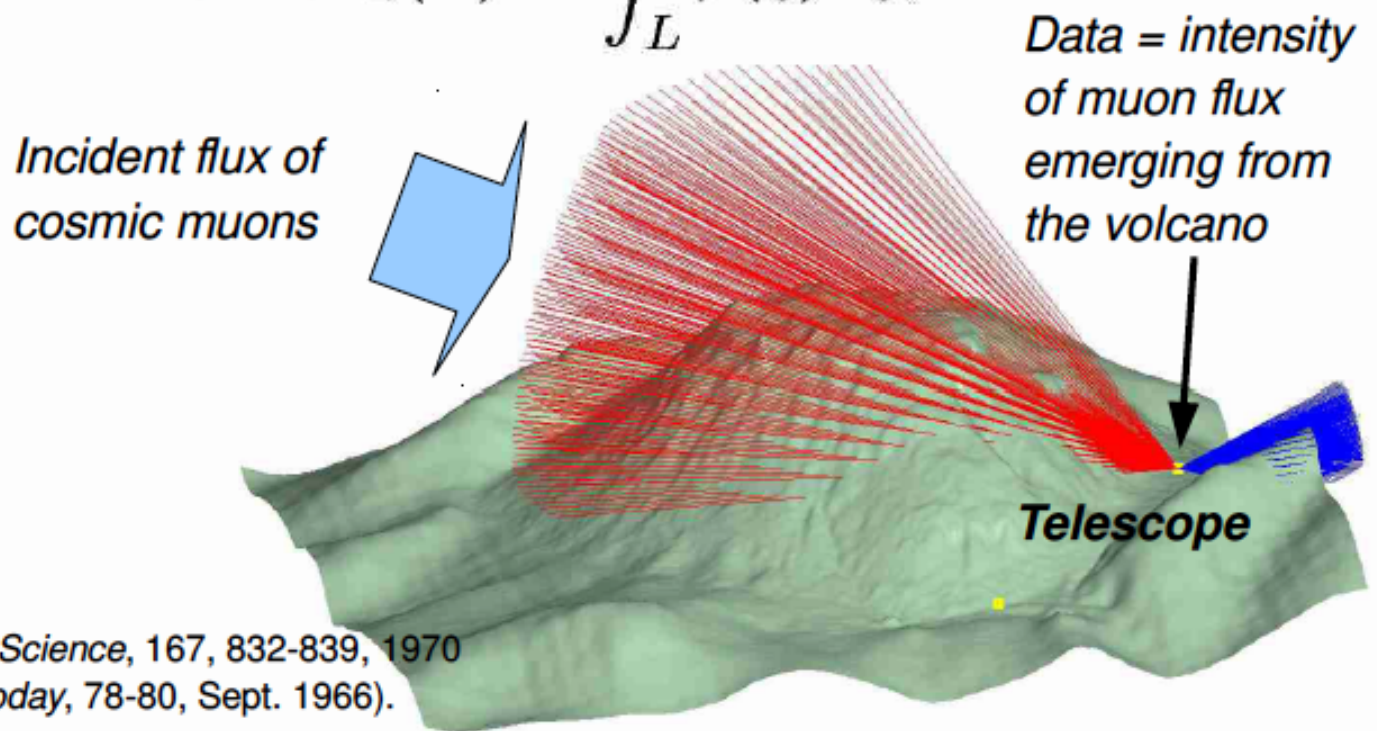


Muon tomography

Muon tomography

Measure the flux of muons across the volcano to determine its density structure

- Basic information = opacity $\varrho(L) \equiv \int_L \rho(\xi) d\xi,$



Precursory work:

Alvarez, L.W. et al., *Science*, 167, 832-839, 1970
(see also *Physics Today*, 78-80, Sept. 1966).

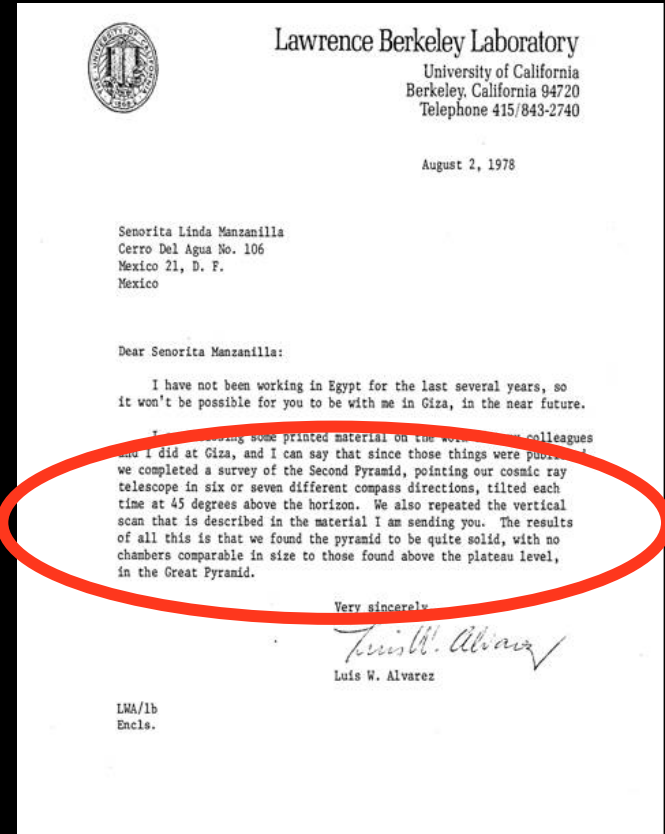
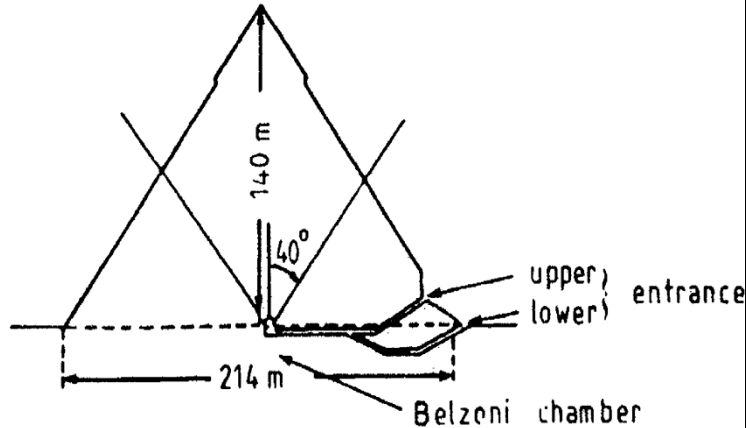
Renew of interest:

Nagamine, K. et al., *NIMA*, 356, 585-595, 1995.

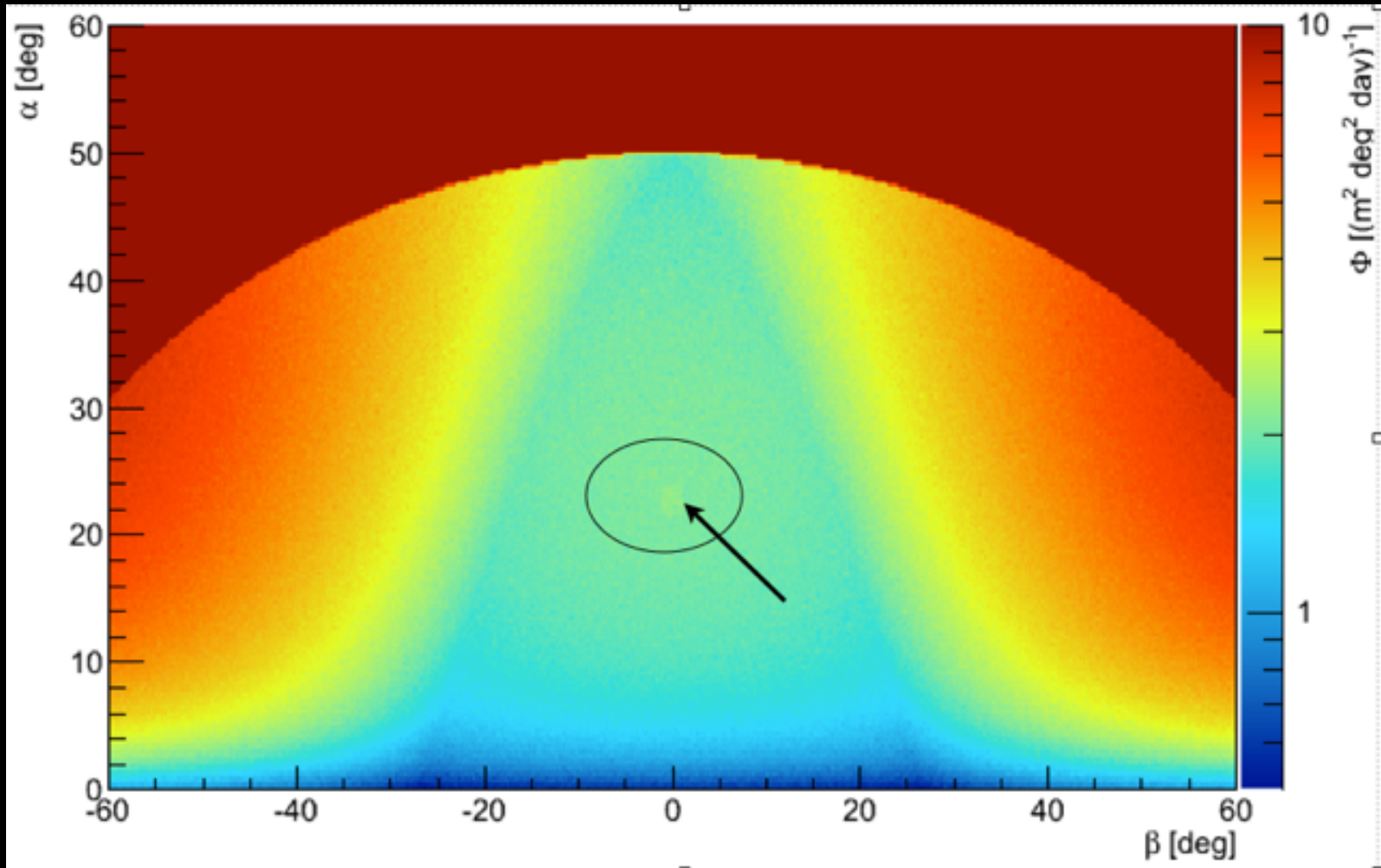
The Alvarez experiment, cosmic ray search for hidden chamber on the Chephren Pyramid



Pirámide de Kefrén



Yes yes the one who working with his son Walter, a geologist at Berkley, developed the hypothesis that an asteroid hit the earth and produced a cloud that covered the planet for a long period of time, blocking sunlight, and eventually leading to the extinction of dinosaurs.



Today's simulation:
A pocket of air of 1m^3 can be seen in a month

Muon tomography

Advantages

- Determination of density *via* a straight-ray geometry
 - Simple inverse Radon transform
 - Curved paths seismic and electrical tomography => non-linear inverse problem
 - Volume integral in gravity measurements
- Wide angle remote imaging => study of active volcanoes

Limitations

- Cannot see « below » => restricted to the top part of volcanoes
- Need to have sufficient statistics => « slow » monitoring with a weekly or monthly timescale

Field telescopes

Photovoltaic panels, wind turbine, fuel cells

Total mass: 200 to 600 kg (w/ shielding)

Angular aperture/resolution: $30^\circ - 60^\circ / 1^\circ - 2^\circ$



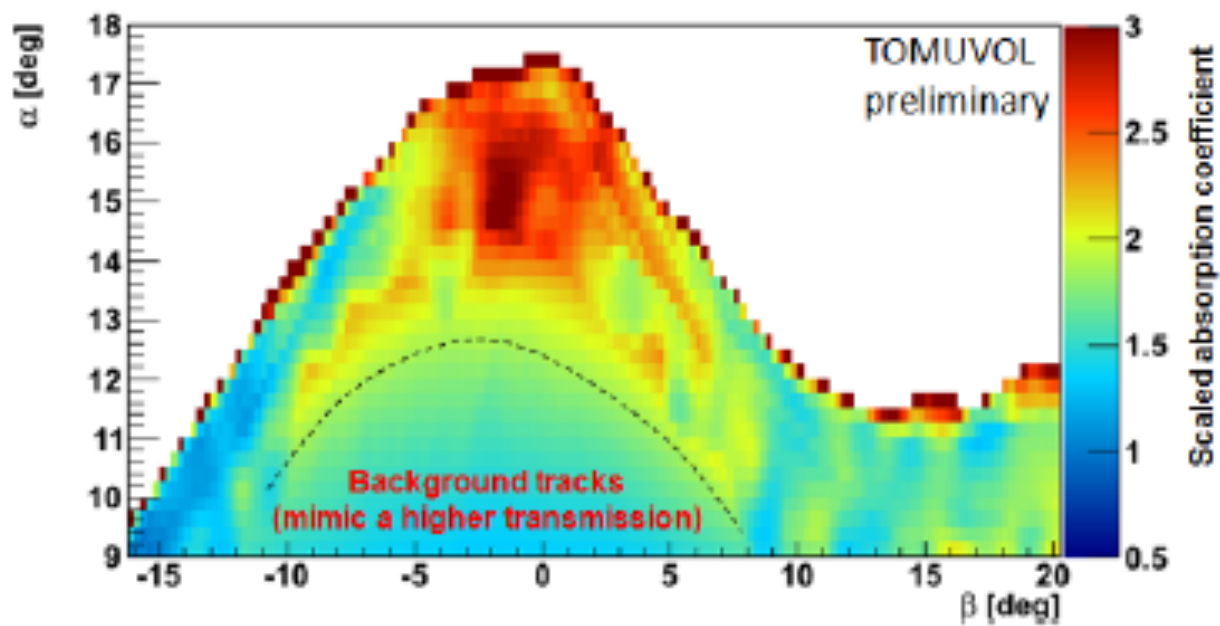
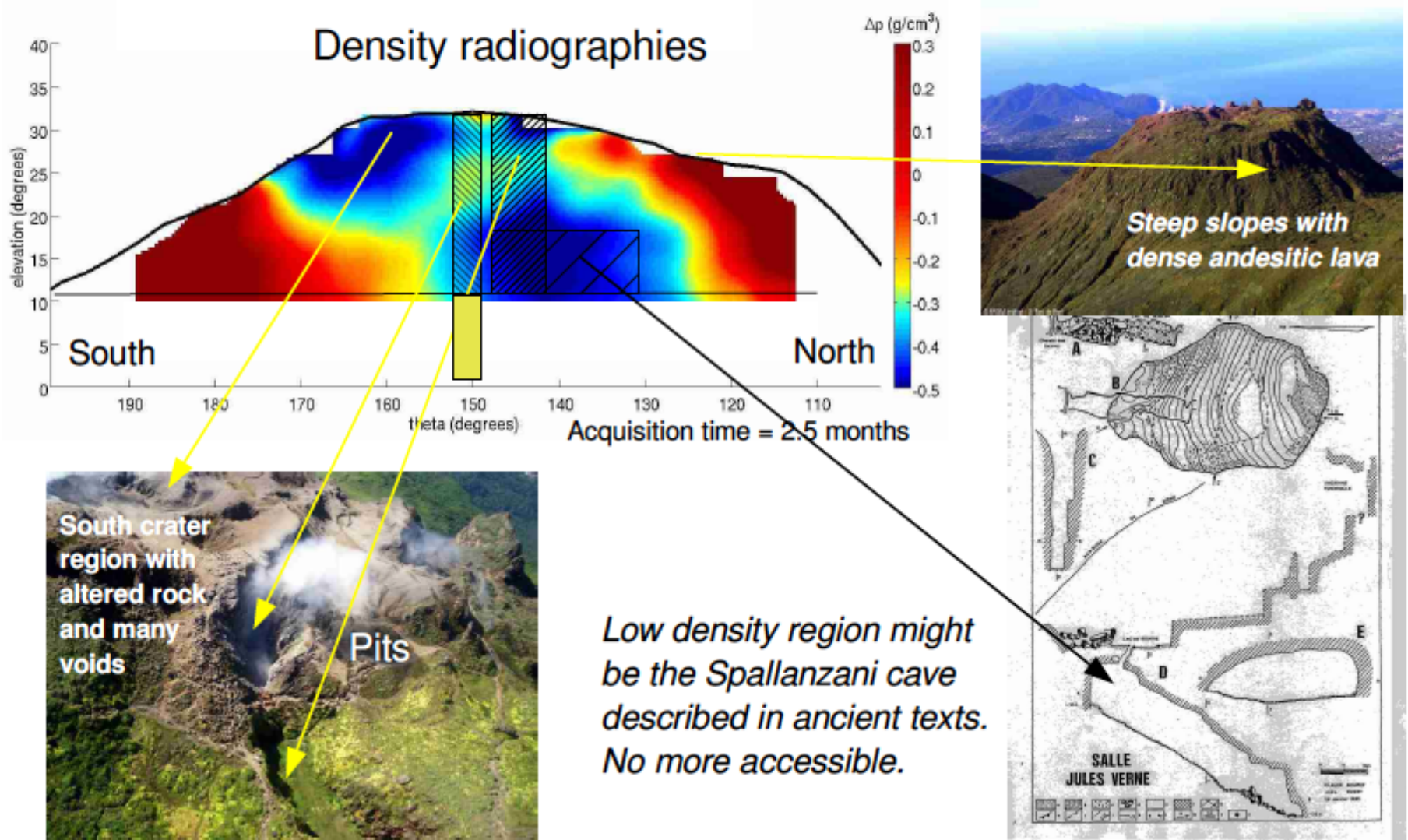
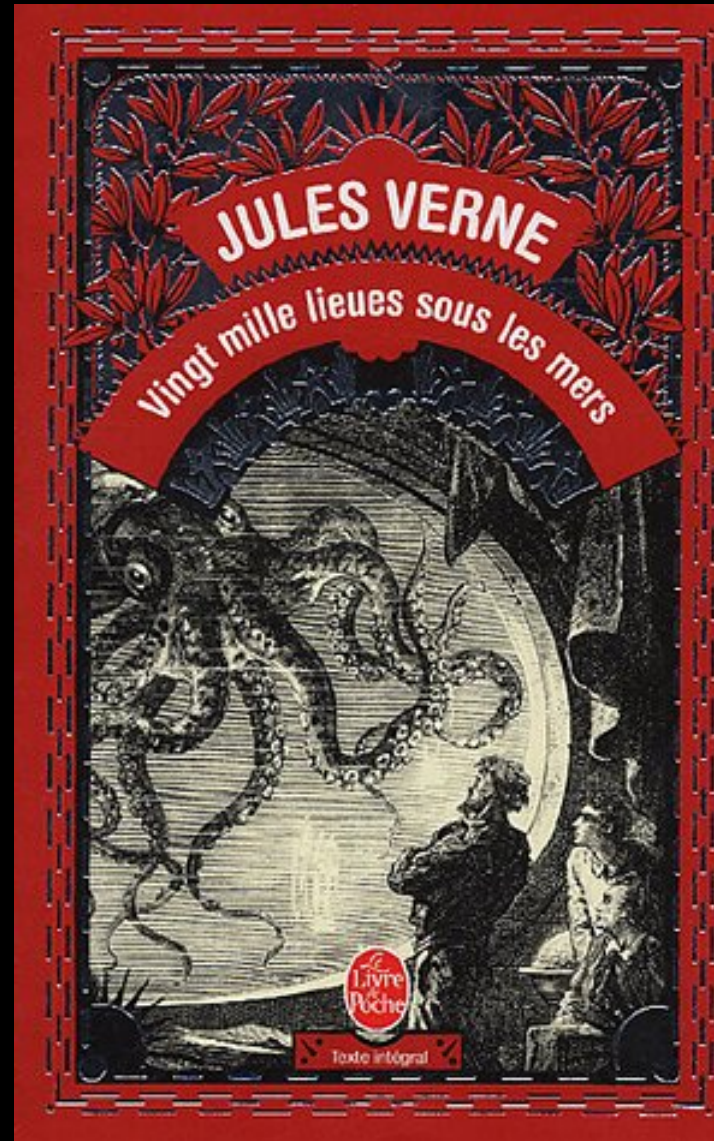


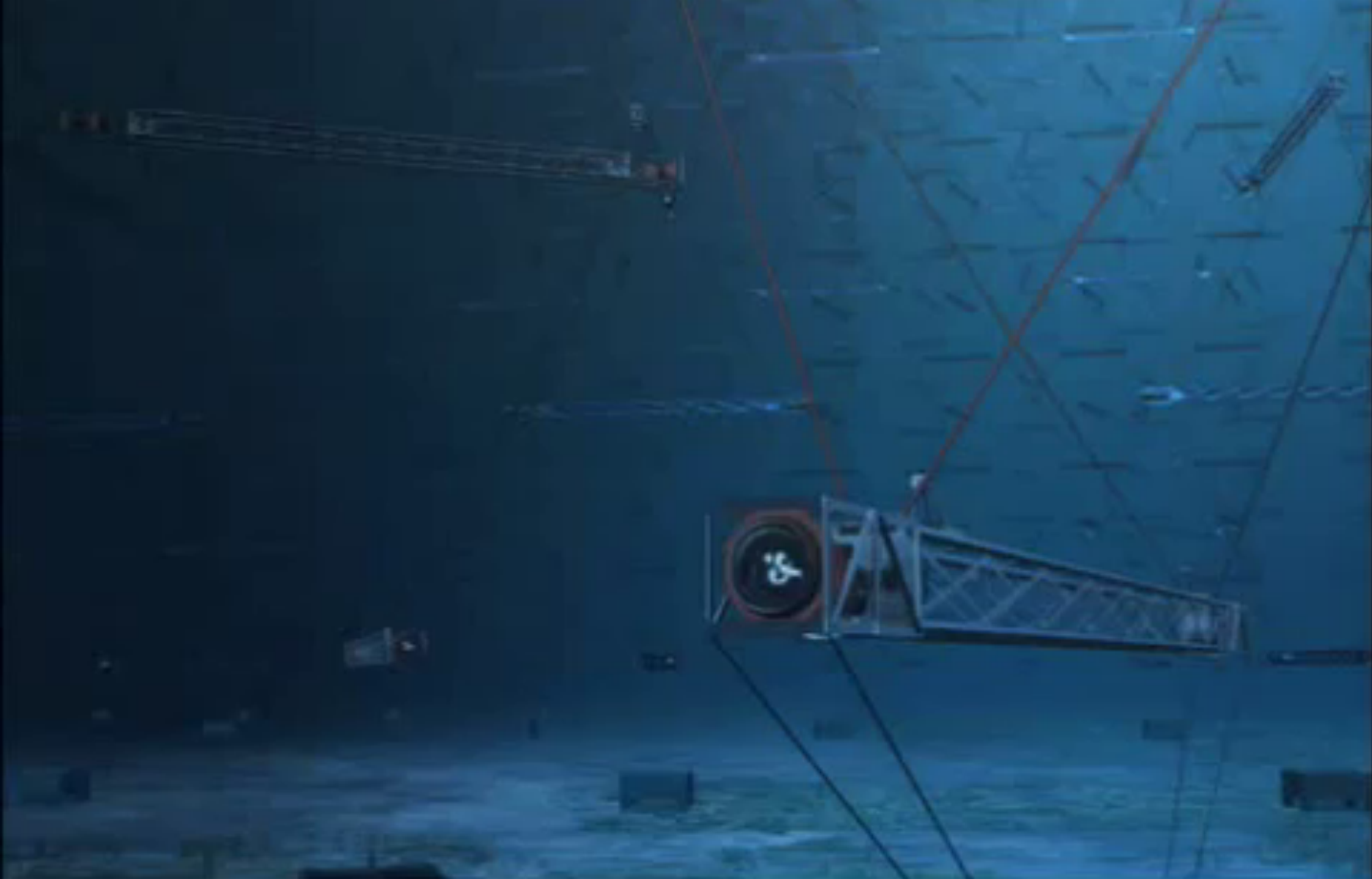
Fig. 3. Map of the scaled transmission through the Puy de Dôme as seen over seven months from the Grôte de la Taillerie with a $\sim 1/6\text{m}^2$ detector

La Soufrière structural imaging





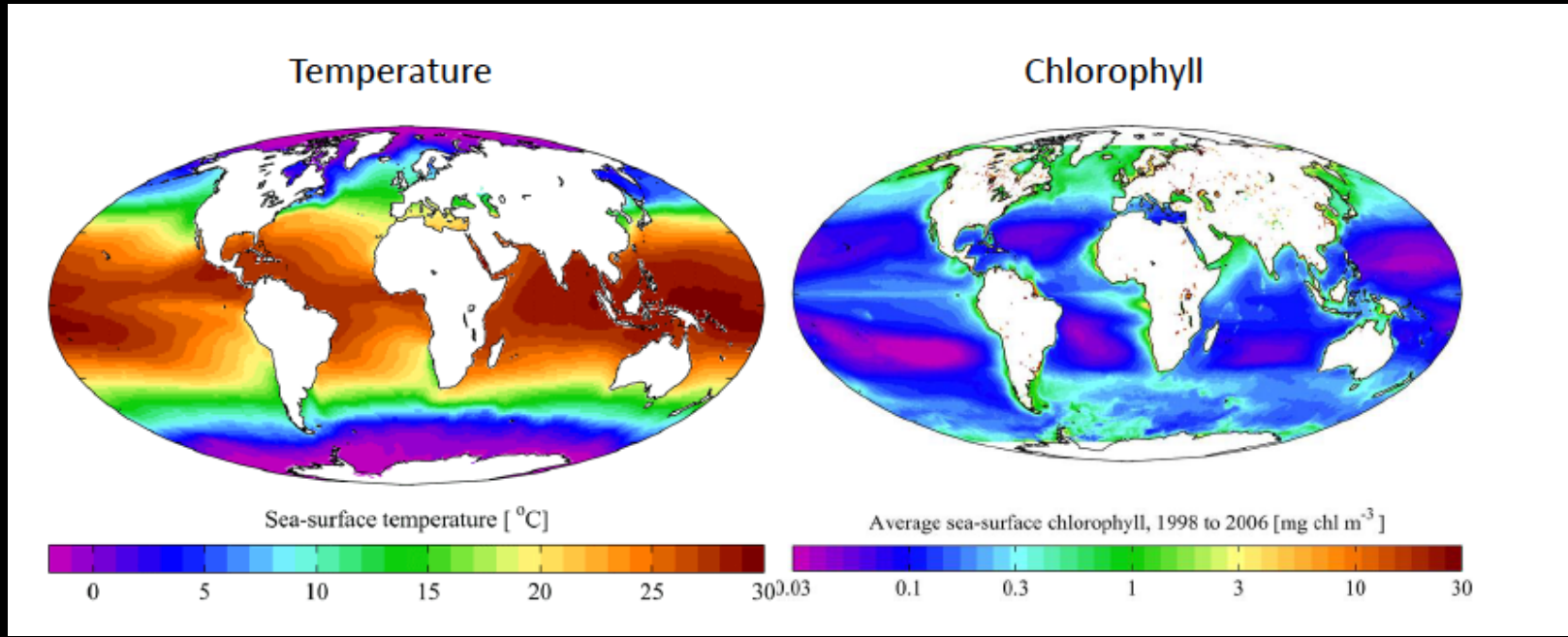
Under-water/ice synergies



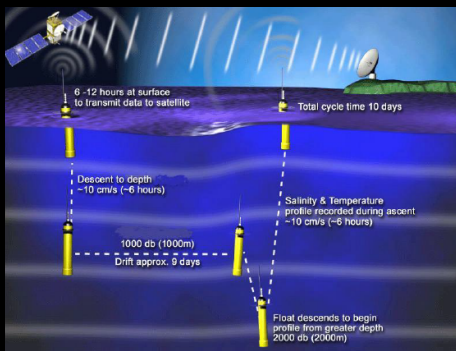
ASTRI E PARTICELLE - LE PAROLE DELL'UNIVERSO



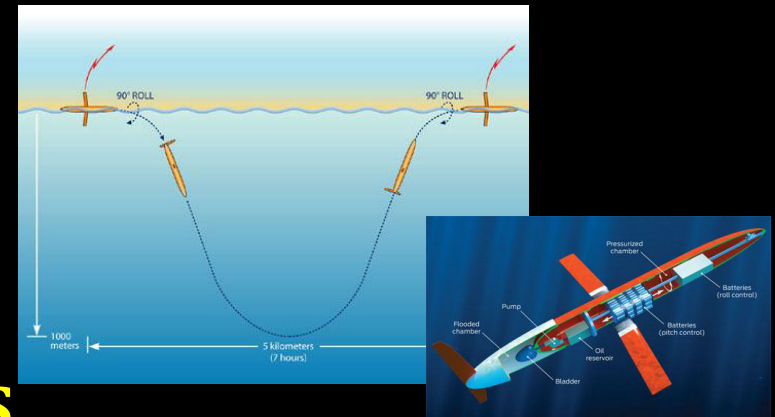
Satellite remote sensing has revolutionised oceanography but measurements are still required from the Ocean's interior



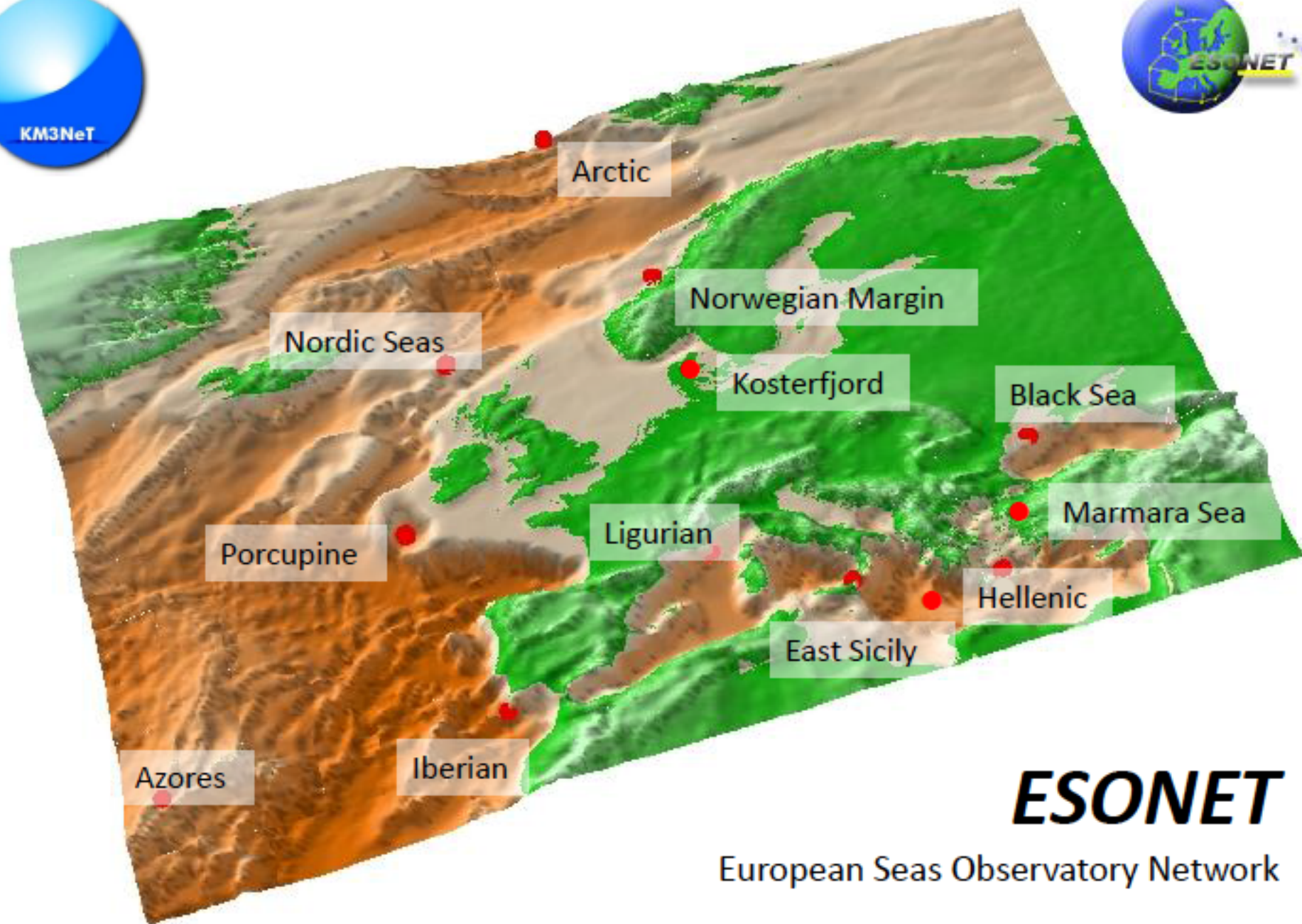
ArRGO floats, controlled buoyancy
emerge every 10 days



Gliders: controlled buoyancy +
wings +GPS



NEED DEEP OCEAN OBERVATORIES

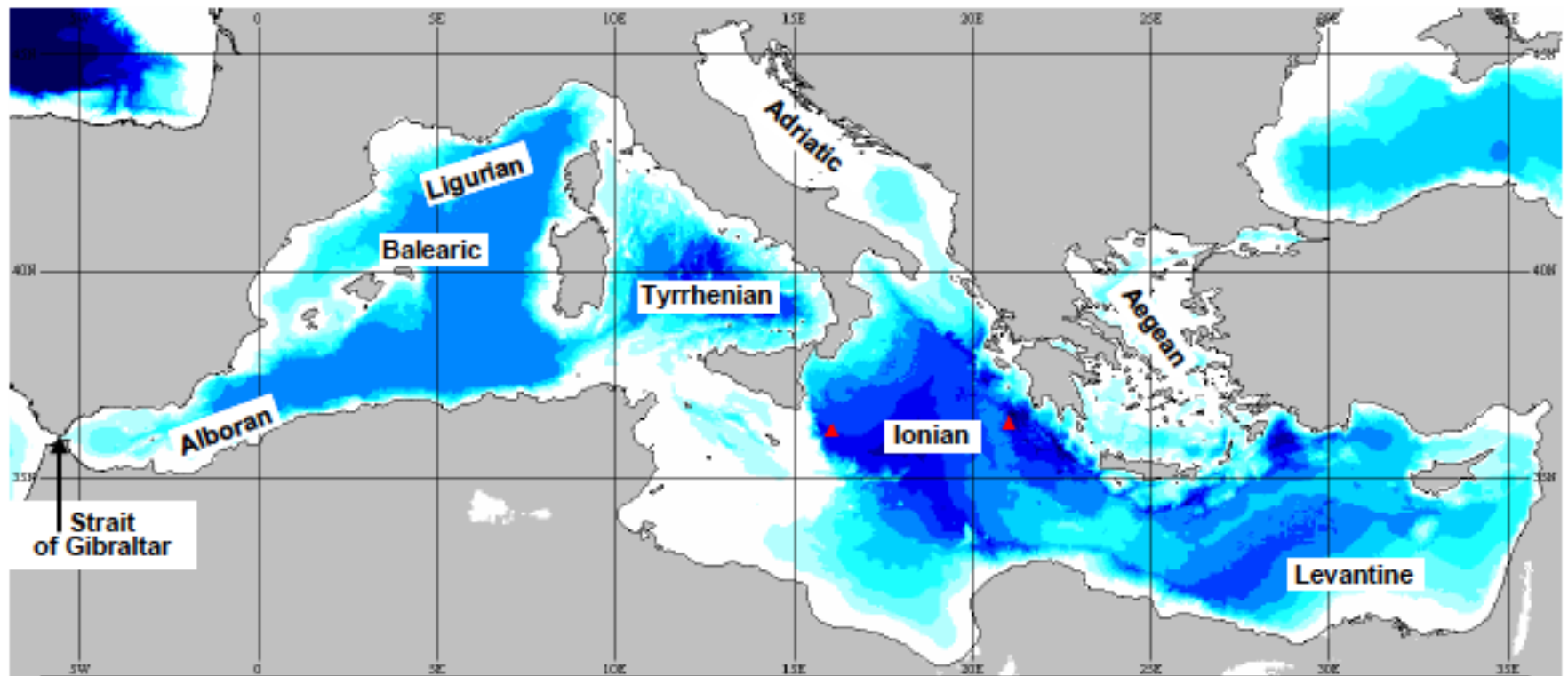


ESONET

European Seas Observatory Network



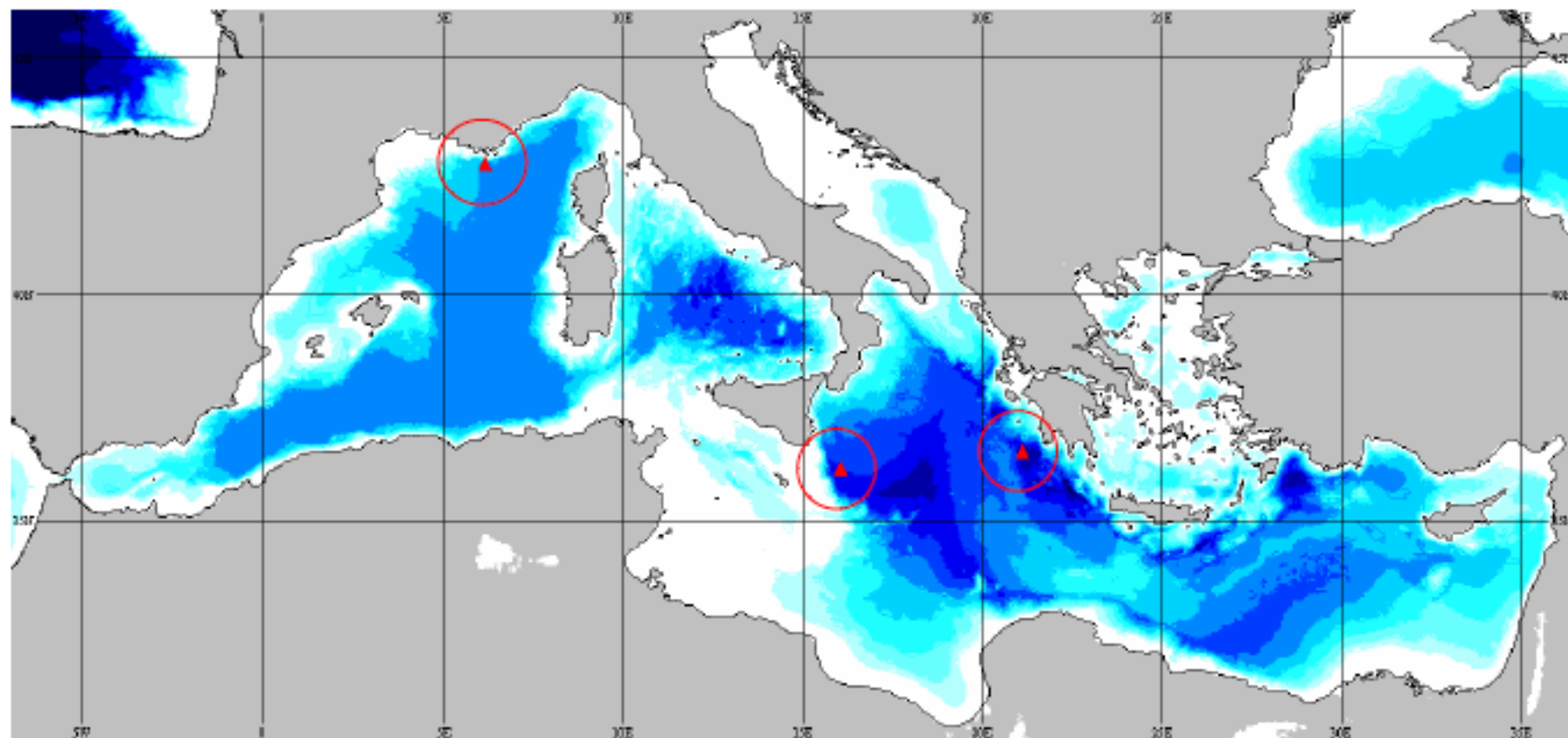
Mediterranean Regional Seas



Colour depth contours at 500m increments.



KM3NeT Sites 100km radius Potential Footprints



ANTARES- Ligurian Sea
2475 m

NEMO - Ionian Sea
3350m

NESTOR - Ionian Sea
4500/5200m



Advantages of **cabled** observatories:

Real-time

High bandwidth

High frequency

Continuous

Long term

Oceanography (water circulation, climate change):

Current intensity and direction, water temperature, water salinity, oxygen, radionuclides...

Geophysics (geohazard):

Seismic phenomena, low frequency passive acoustics, magnetic field variations,...

Biology (micro-biology, cetaceans,...):

Passive acoustics, biofouling, bioluminescence, video, water samples analysis,...

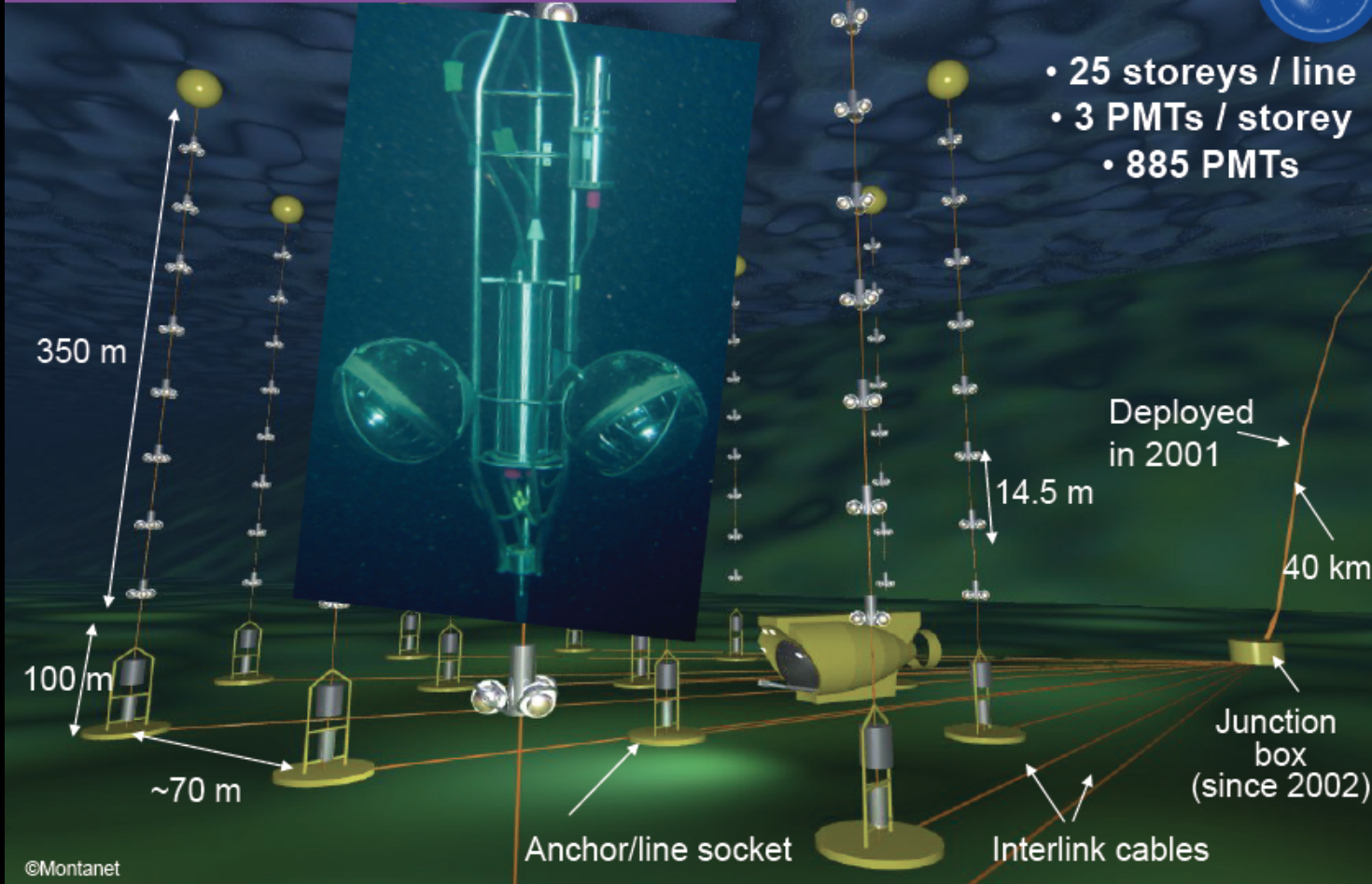


The ANTARES neutrino telescope

Detector completed in May 2008



- 25 storeys / line
- 3 PMTs / storey
- 885 PMTs

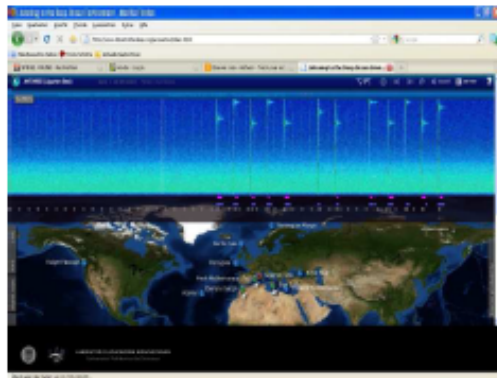


The position of the PM determined with cm precision S and P detectors ?
F. Cornet

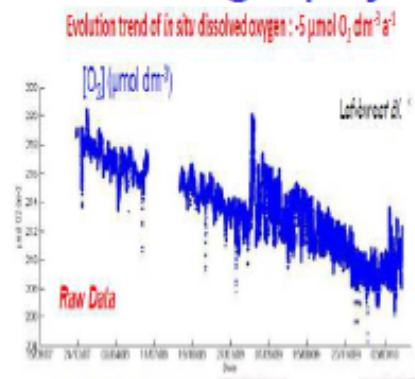
Neutrino Telescopes: Marine Sciences

Deep Ocean Cabled Observatories Aspera Workshop, Amsterdam 2011-
<https://indico.cern.ch/conferenceDisplay.py?ovw=True&confId=165389>

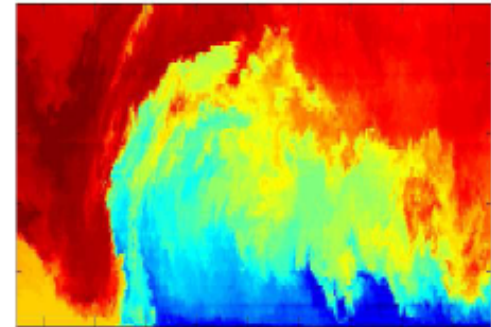
Marine Acoustics



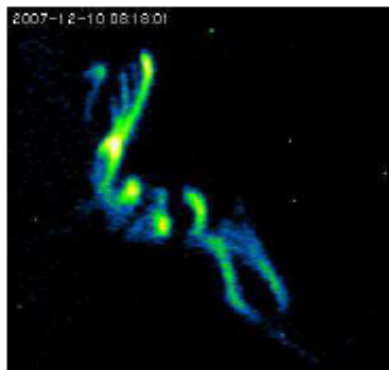
Oceanography



Internal wave dynamics

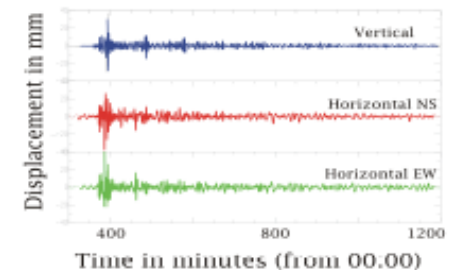


Bioluminescence



Seismology

Japan earthquake 2011 March 11
at Antares site

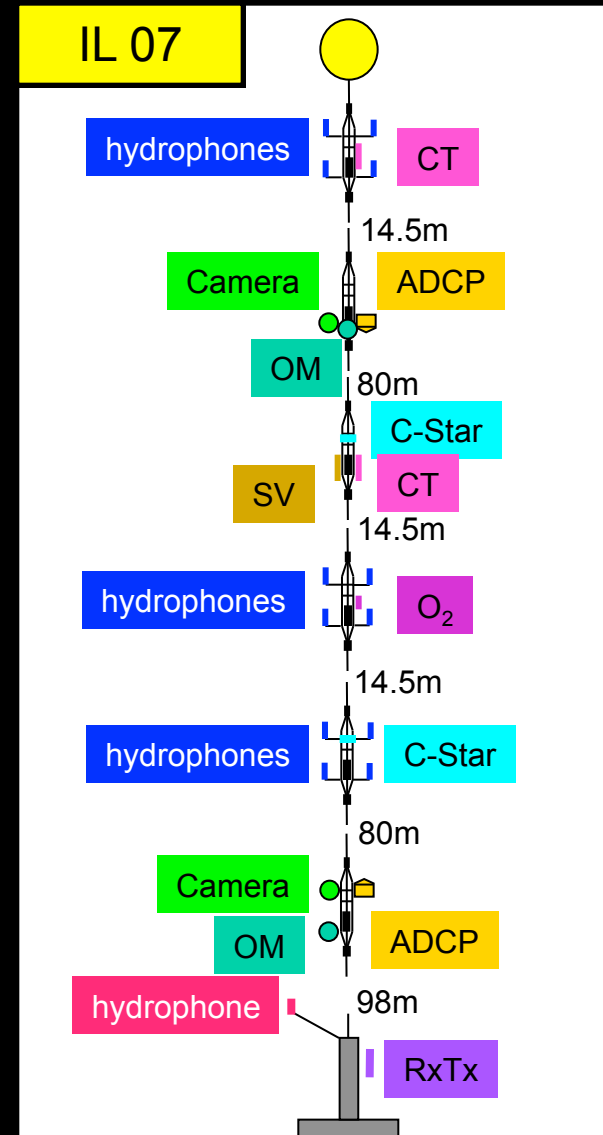
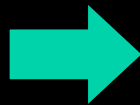




ANTARES: Instrumentation Line

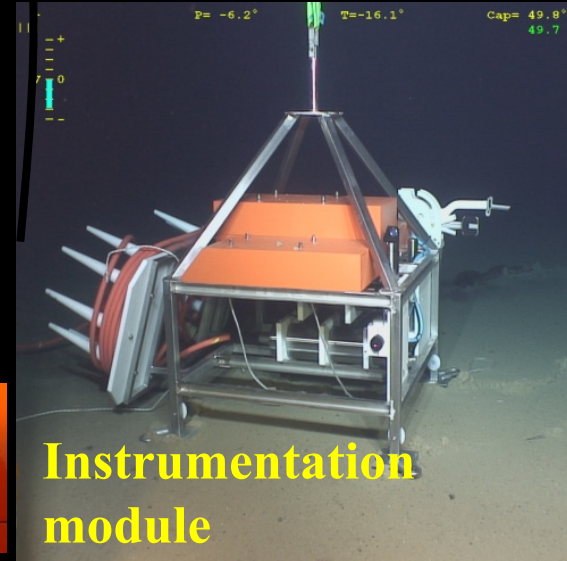
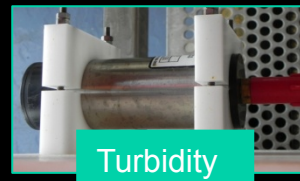
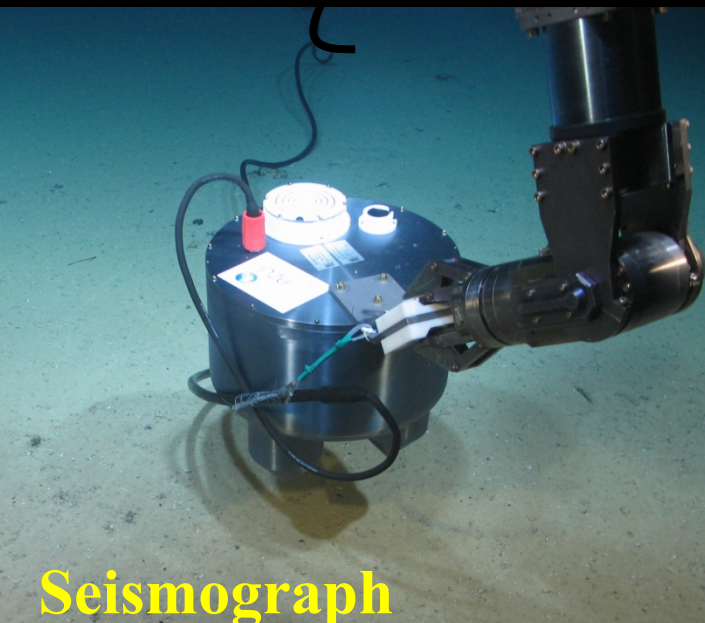
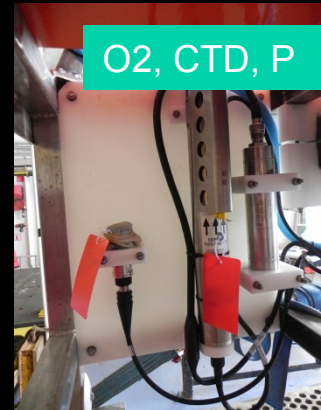
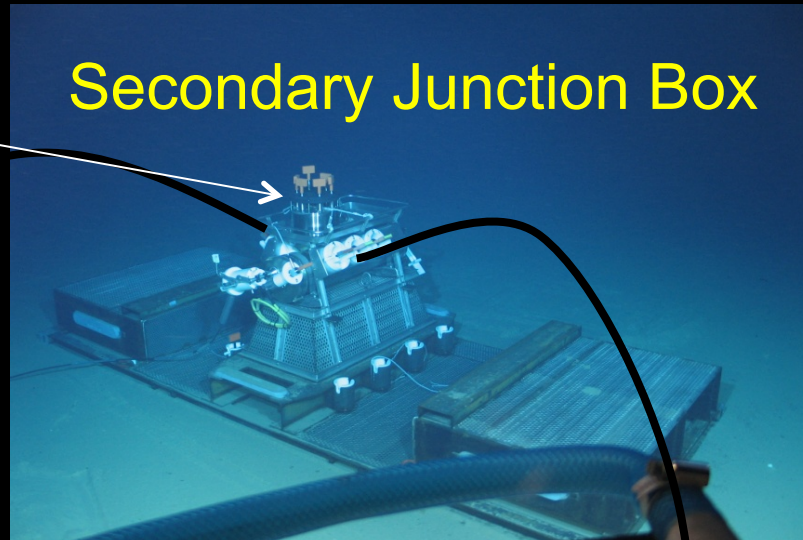
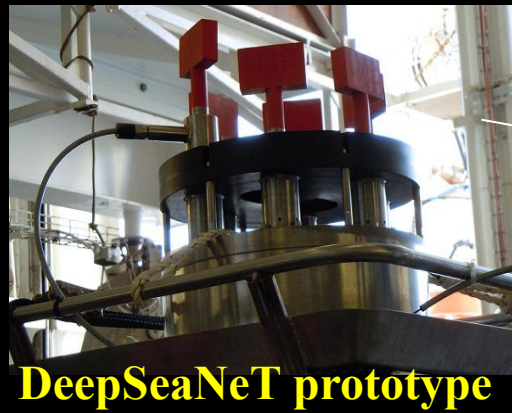
Monitoring of Environmental Parameters (reconnection planned later this year)

- **CSTAR** light transmission
- **CT** = Conductivity-Temperature
- **SV** = sound velocity
- **ADCP** = Current meter
- GURALP seismometer
- 2 Optical Modules
- Acoustic positioning RxTx & Rx
- Oxygen meters
- 2 cameras
- +
- 3 storeys of UHE neutrino acoustic detectors





Secondary Junction Box

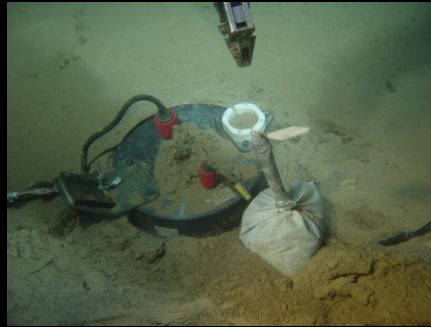




Seismology



In laboratory

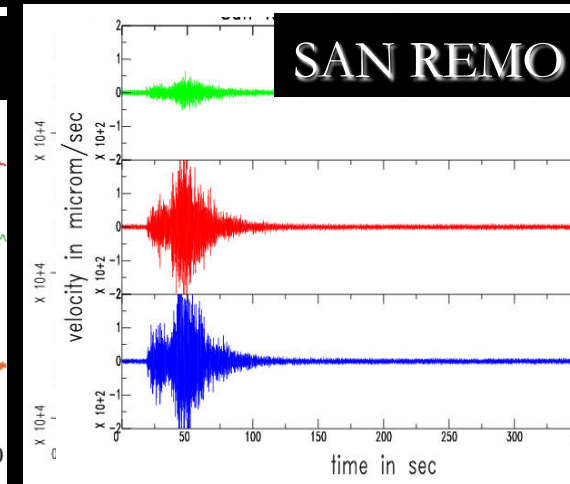
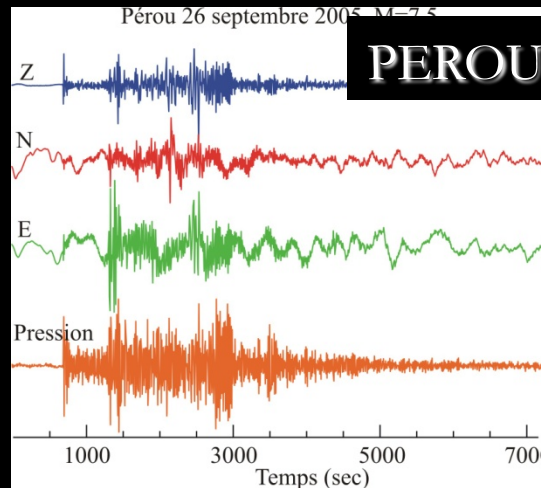


deployment



Buried at site Antares
(gain 20 dB of noise)

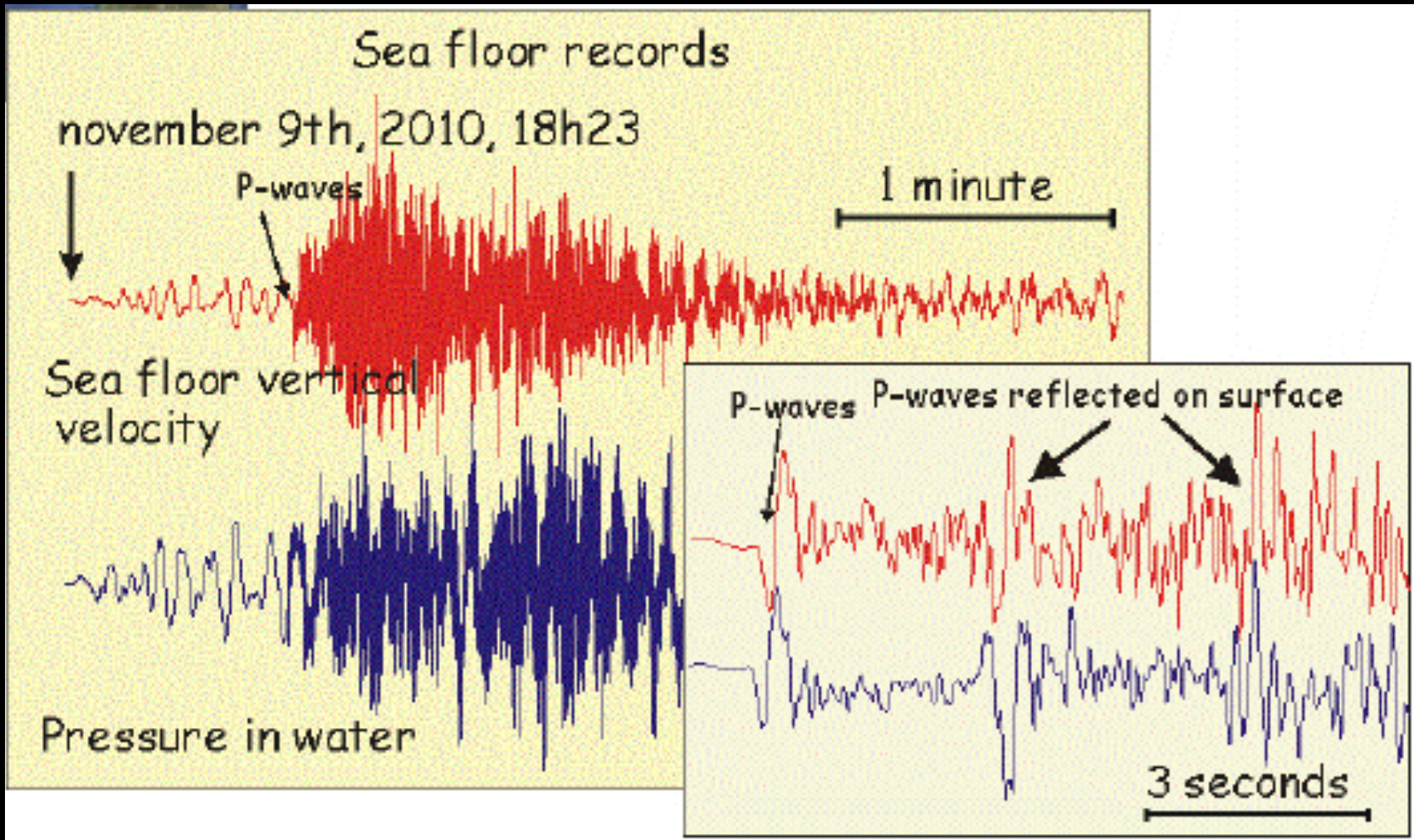
The seismograph ANTARES is part of seismic monitoring network, complementary to the terrestrial stations.



Collaboration with Laboratoire Géosciences Azur

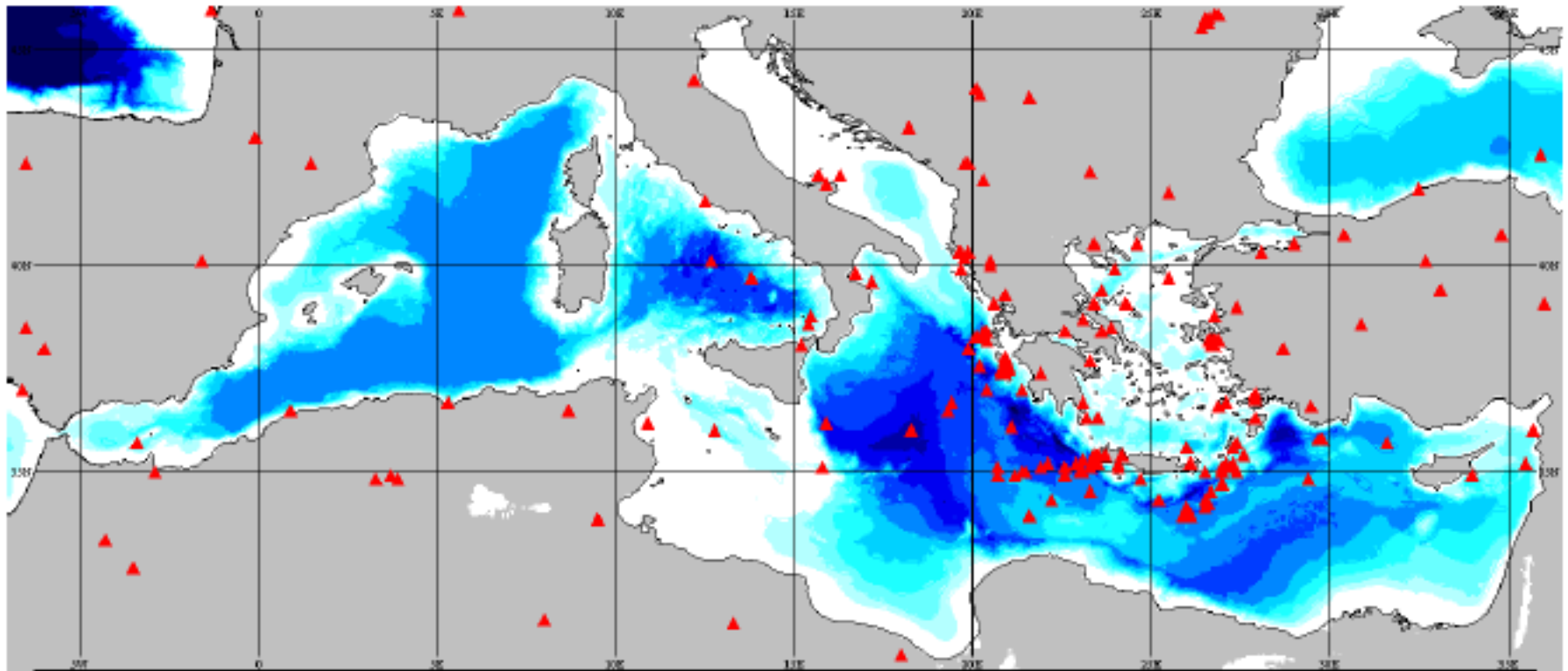


Nov 9th, 2010: Magnitude=3.6-4.2





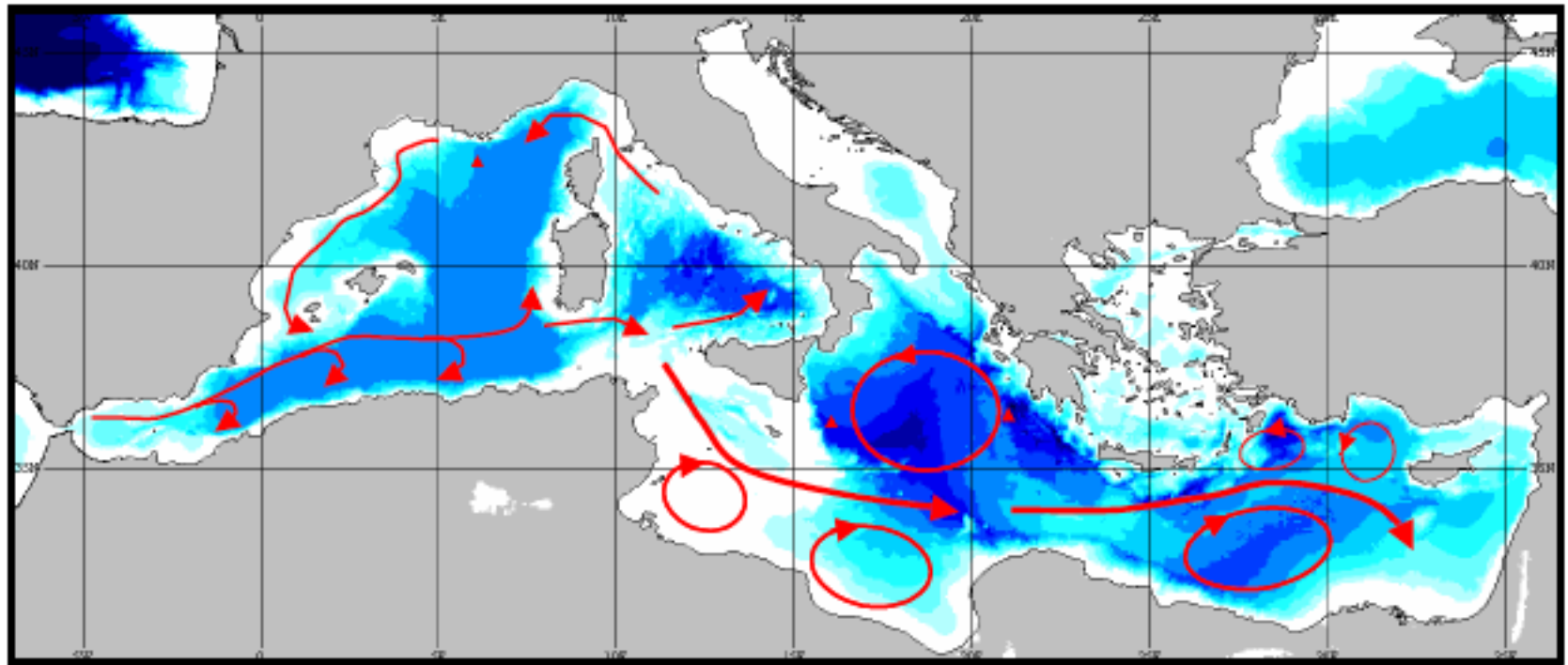
Earthquakes during 2005 and 2006.



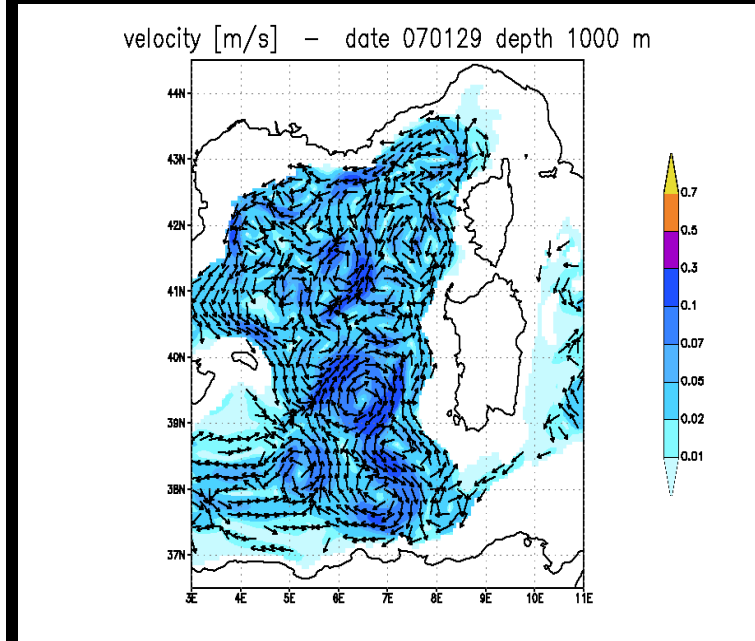
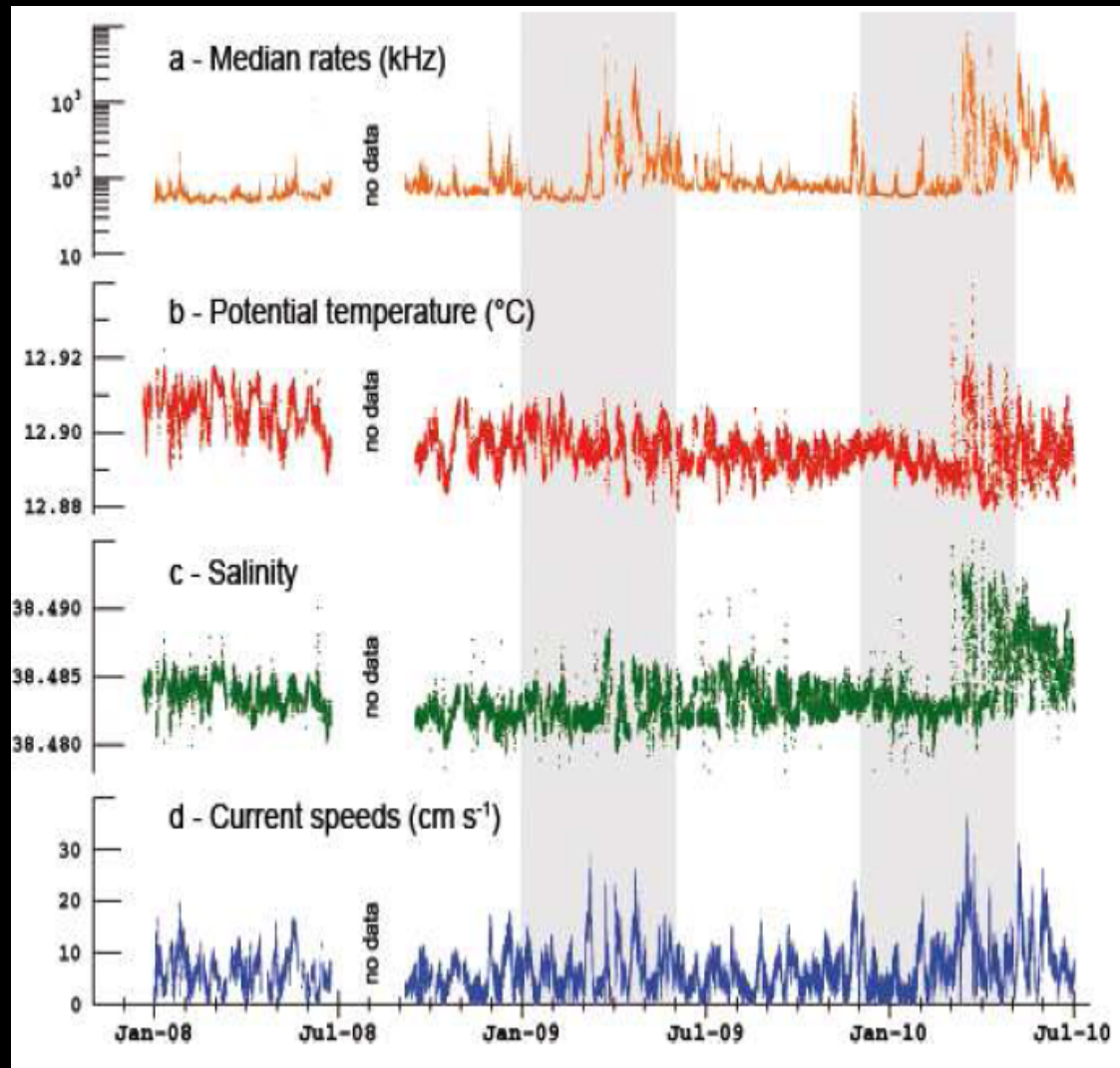
Can we instrument large areas of the Mediterranean with seismometers ?



Surface circulation



ANTARES: Long Term Oceanographic Parameters

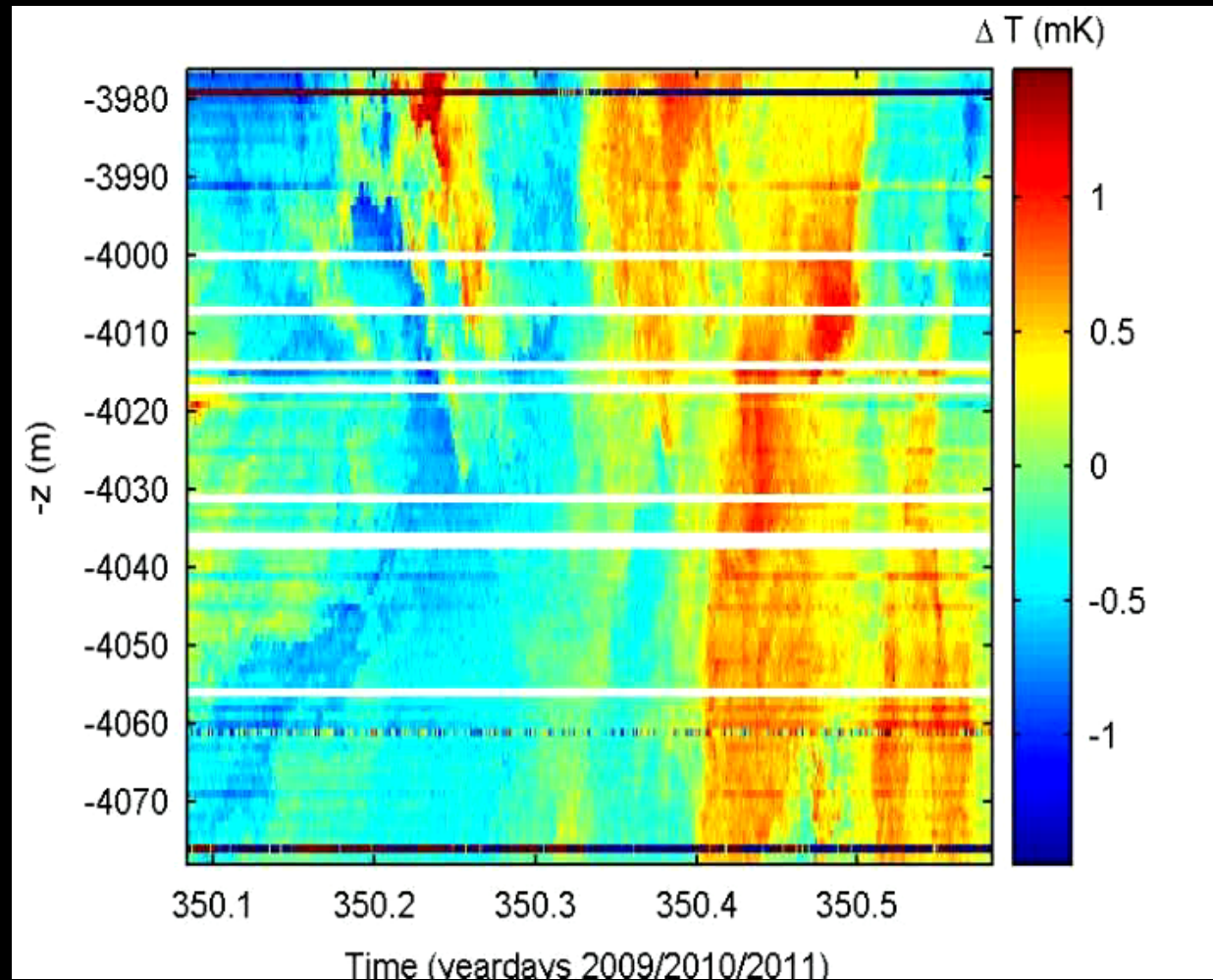


Part of hydrochanges
Mediterranean network

Member of Eurosites



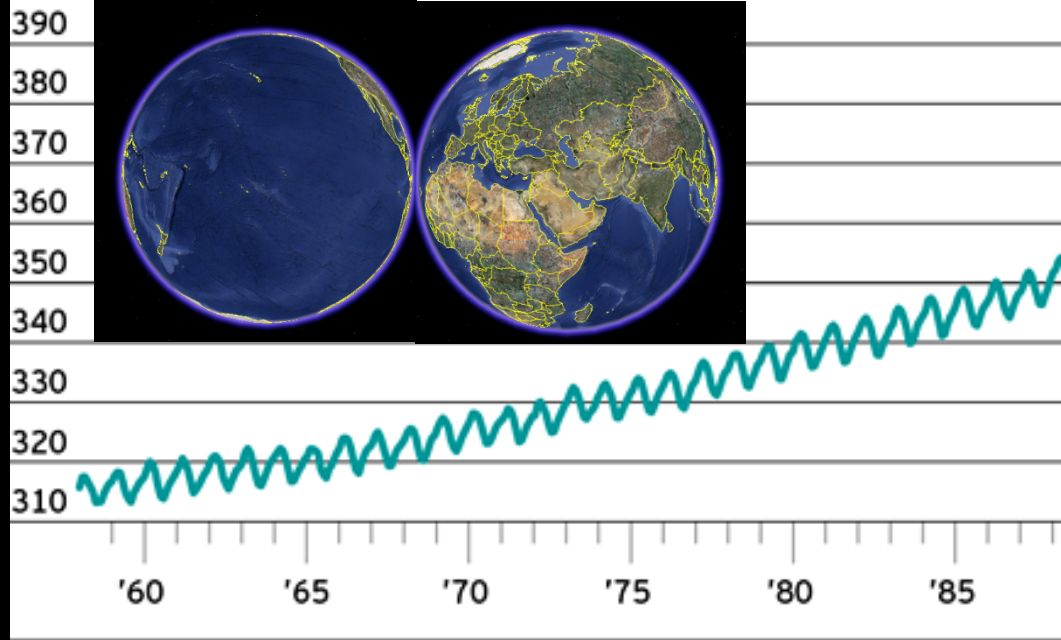
Monitor smaller scale variability , turbulence



The Breathing Planet

ATMOSPHERIC CARBON DIOXIDE

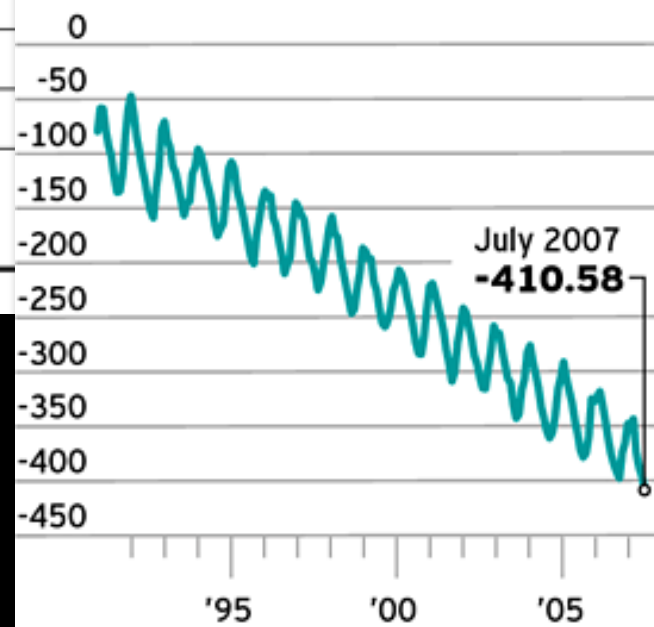
Carbon dioxide concentration (parts per million by volume)



SOURCE: Scripps Institution of Oceanography

ATMOSPHERIC OXYGEN

Oxygen concentration per meg*



*A loss of 100 per meg corresponds to losing 100 out of 1 million oxygen molecules in the atmosphere

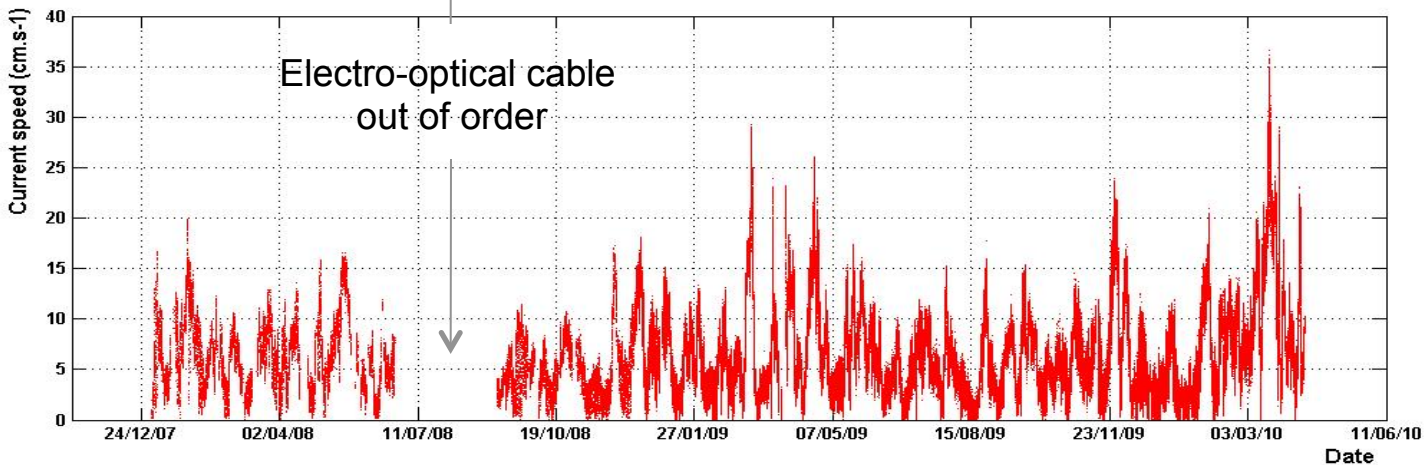
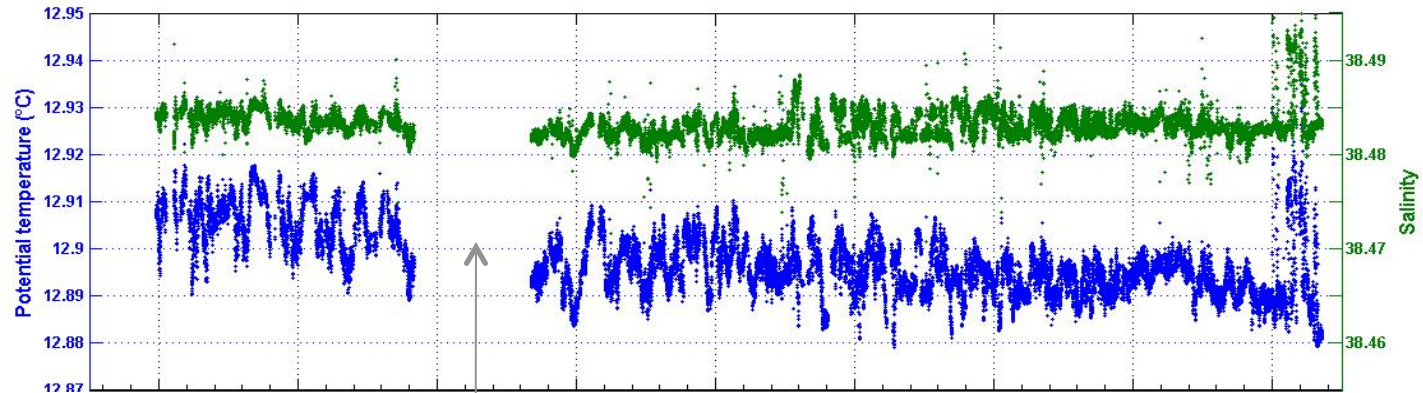
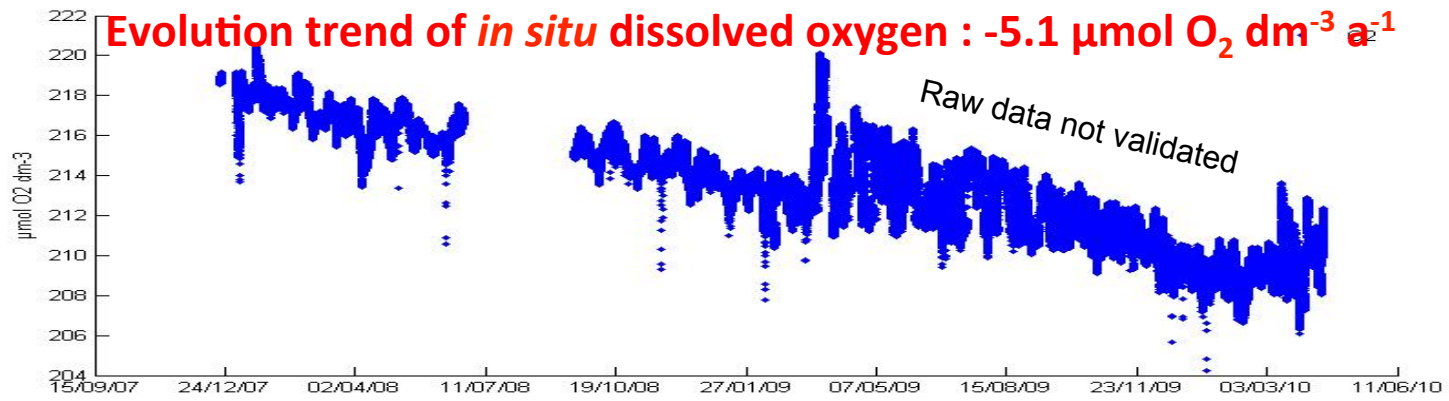
SOURCE: Scripps Institution of Oceanography

CO₂ and O₂ atmospheric trends

Planet earth : A living organism



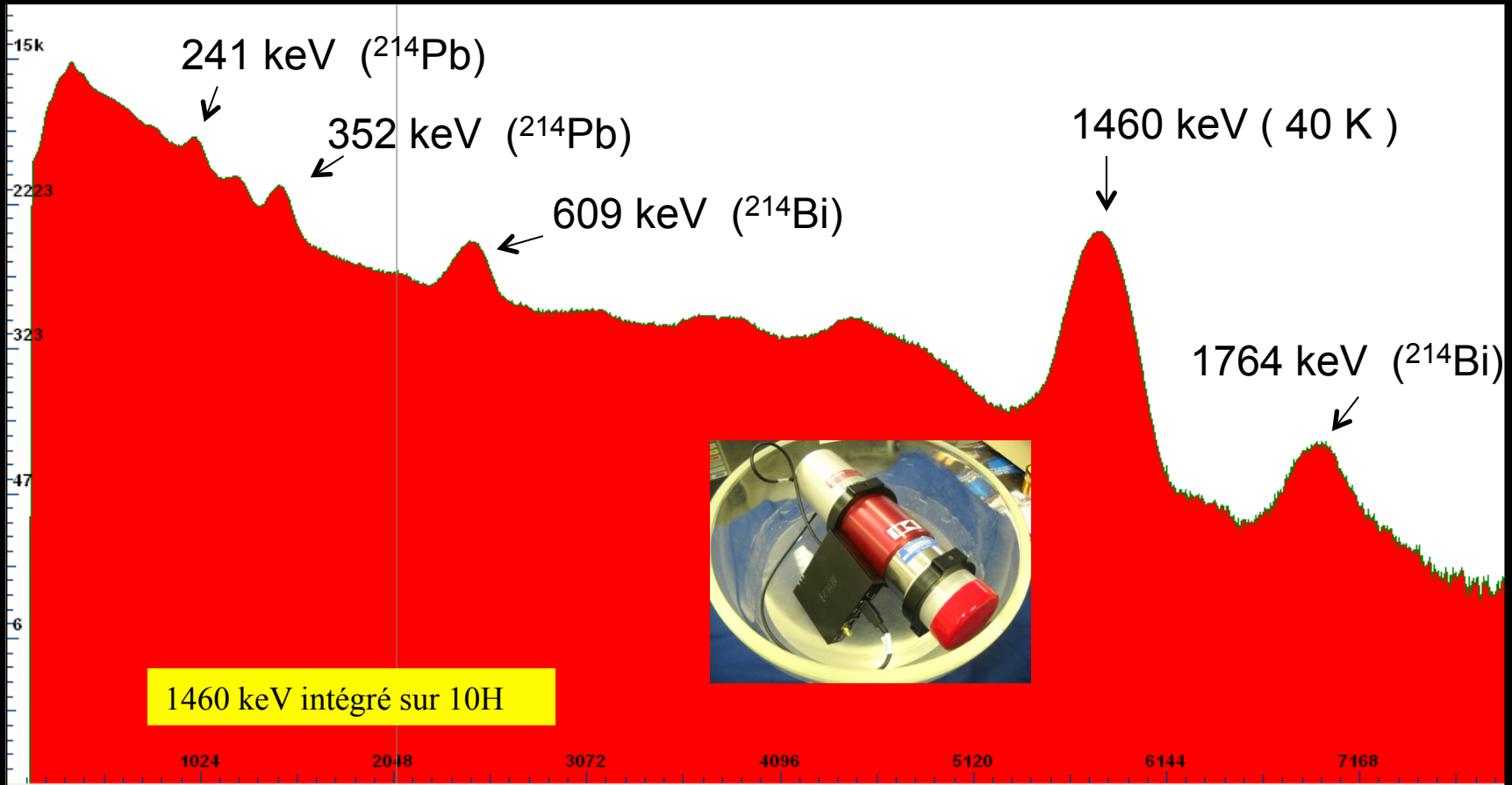
Temperature, Salinity, O₂ time-series (IL07)





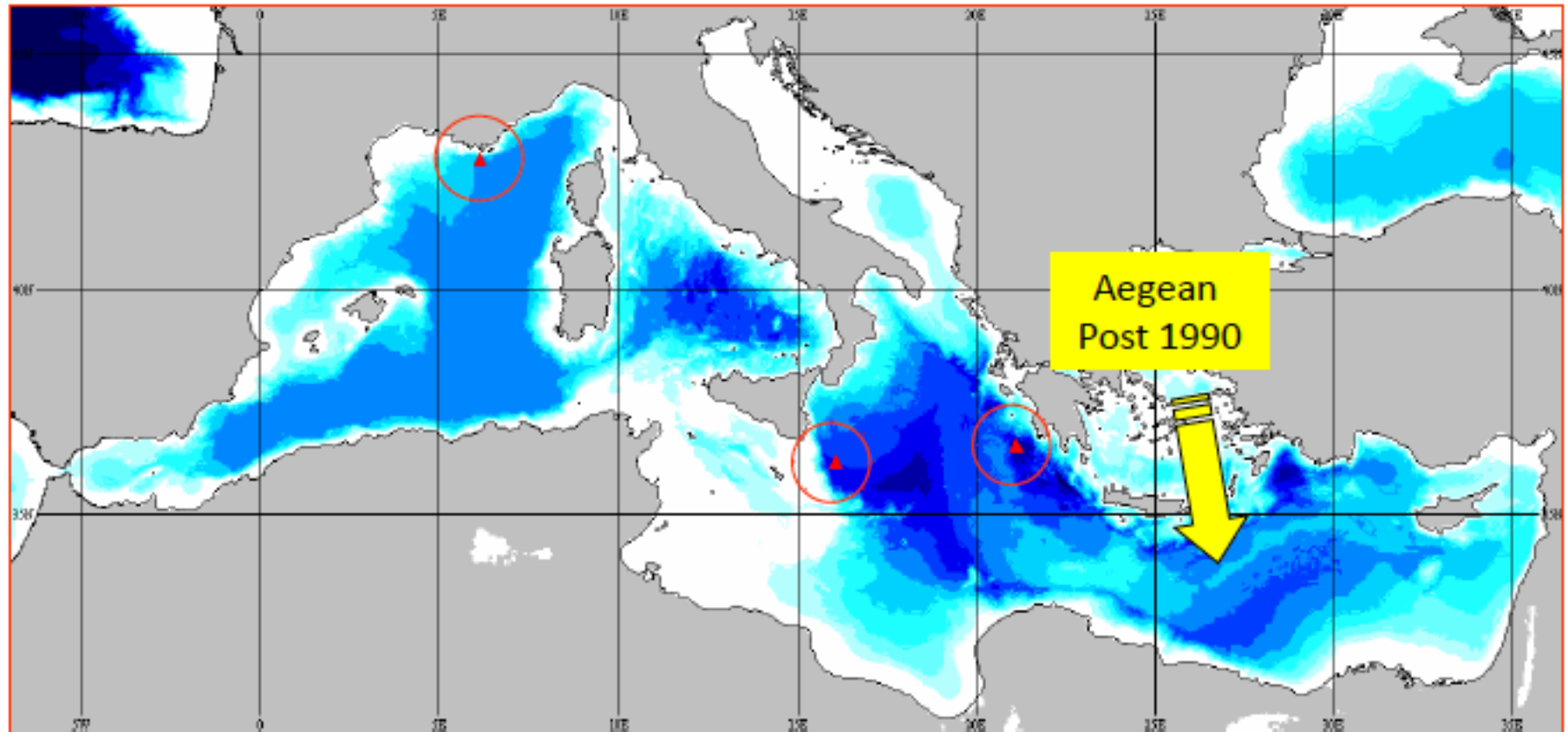
Gamma Spectra in Deep Sea

Log scale



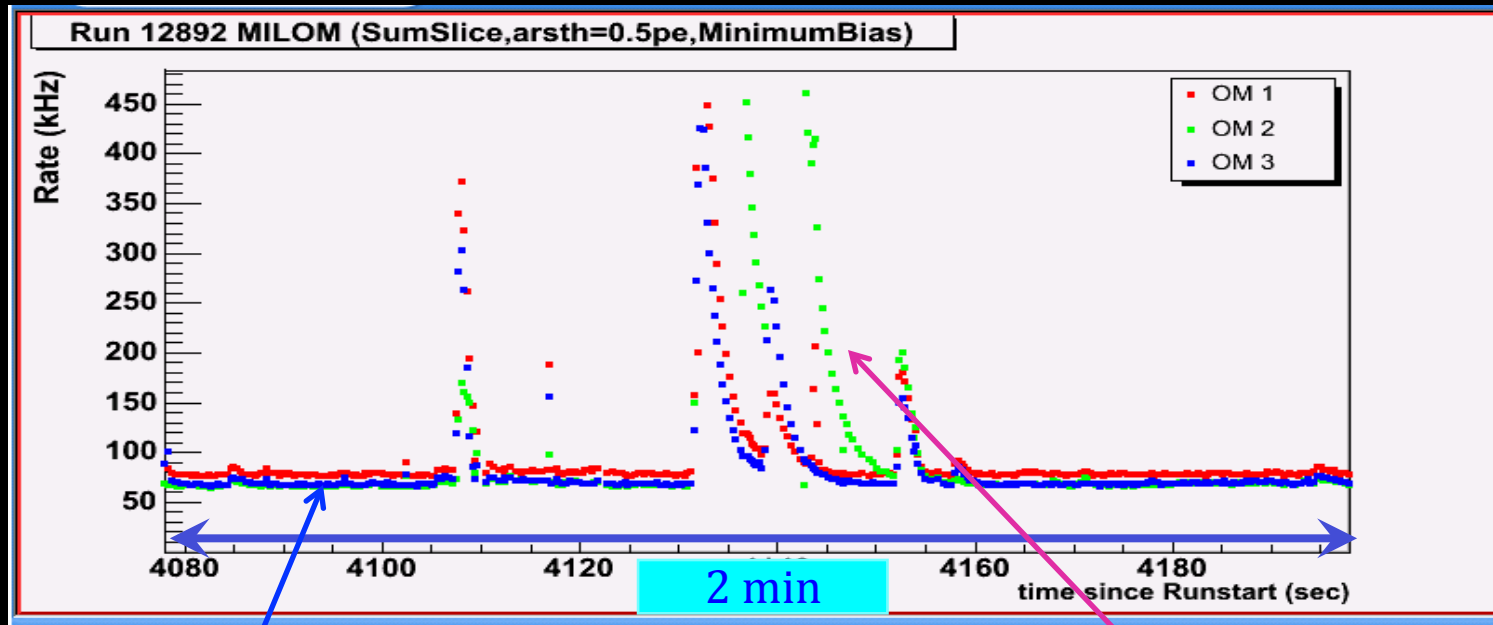


Deep Water Formation Eastern Mediterranean Transient 1990





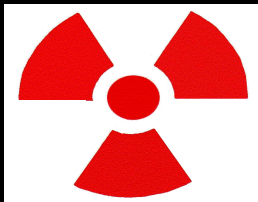
Counting Rates on Optical Modules



Continuous baseline:
Radioactivity in the sea (^{40}K)
+ bioluminescent bacteria?

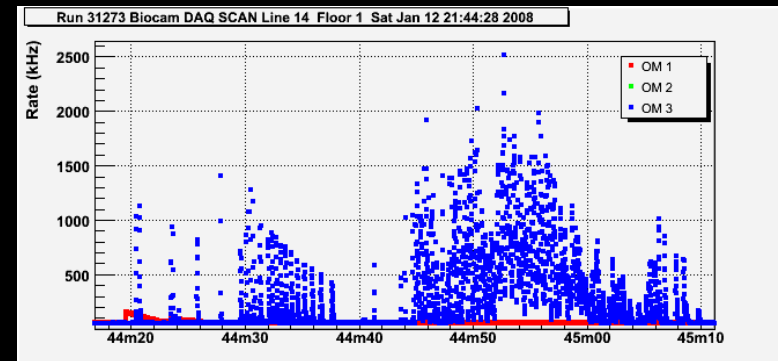
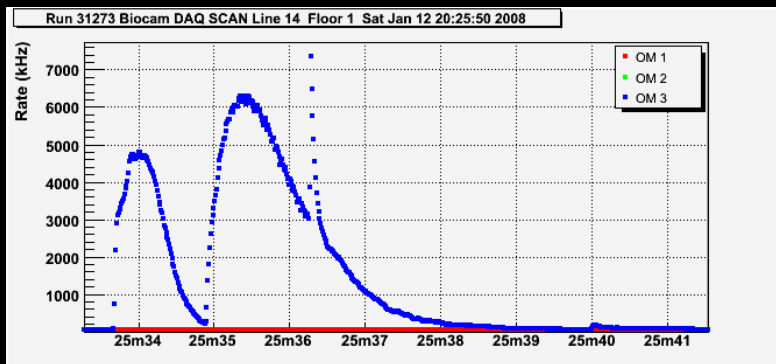
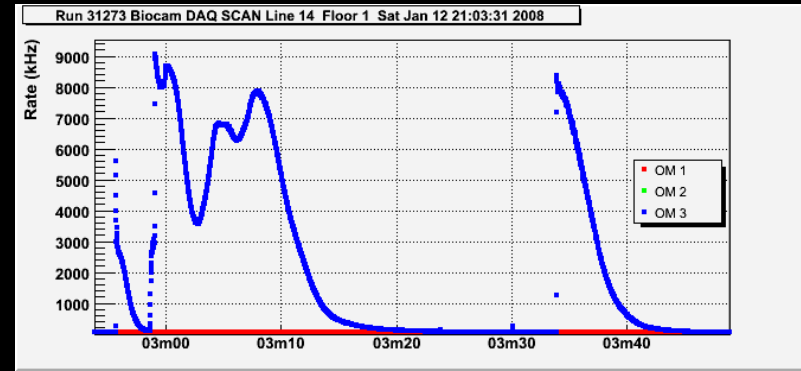
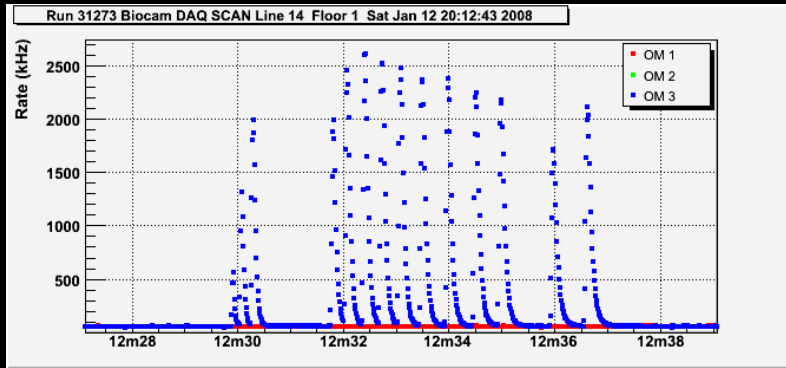
Bursts:
bioluminescence from
Macroscopic organisms

^{40}K

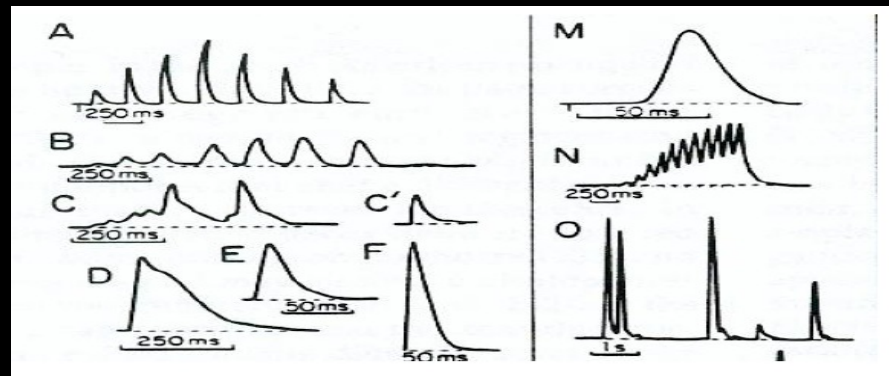




Example Bioluminescence Evt's Observed on PMTs

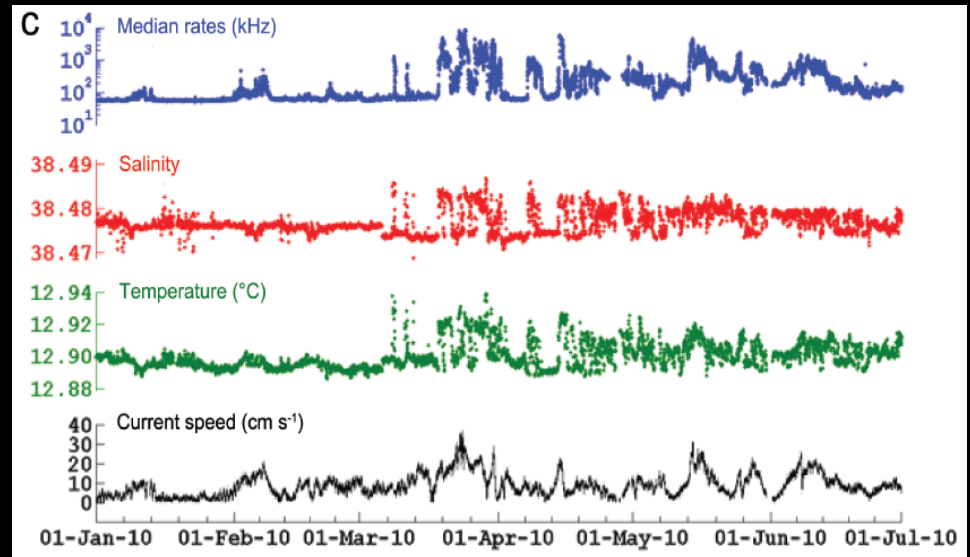
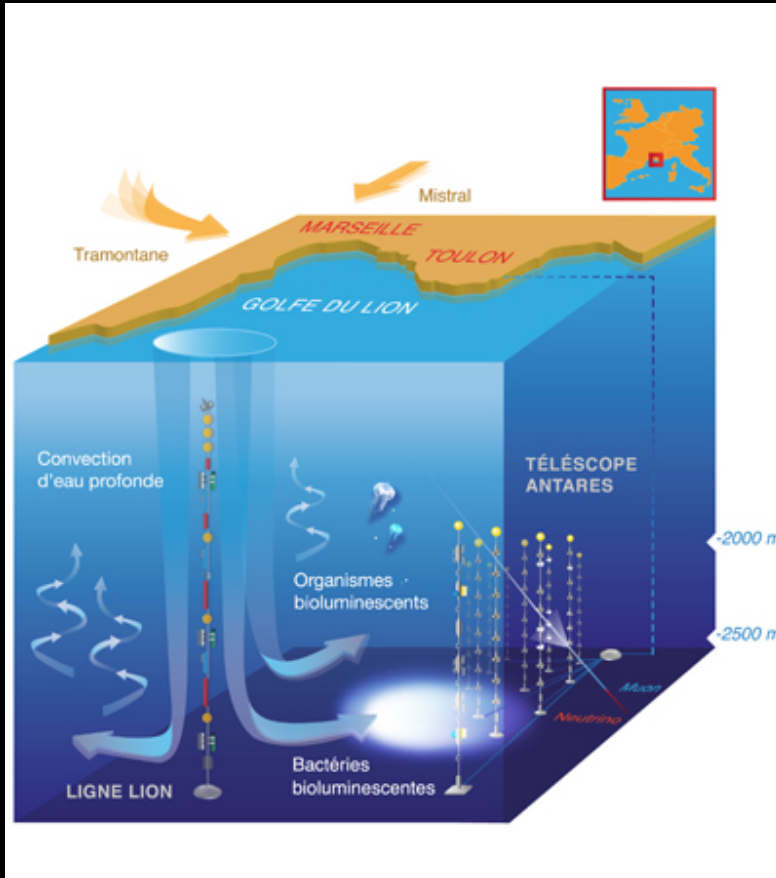


versus literature





Life as proxy of deep water formation



Prix de la Recherche 2014
Catégorie « coup de cœur »

Deep-Sea Research I 58 (2011) 875–884

Contents lists available at ScienceDirect

Deep-Sea Research I

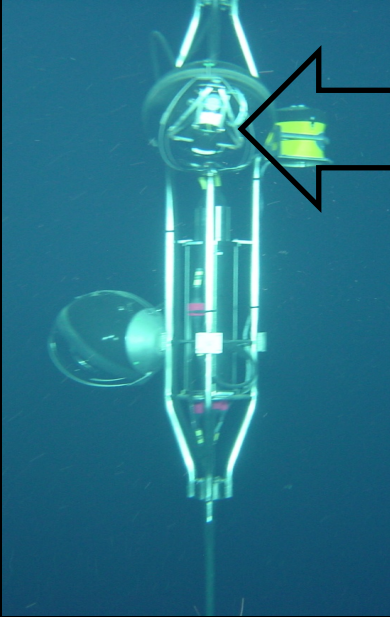
ELSEVIER journal homepage: www.elsevier.com/locate/dsri

Acoustic and optical variations during rapid downward motion episodes in the deep north-western Mediterranean Sea

H. van Haren^{z,*}, I. Taupier-Letage^{ah,1}, J.A. Aguilar^a, A. Albert^b, M. Anghinolfi^c, G. Anton^d, S. Anvar^e, M. Ardid^f, A.C. Assis Jesus^g, T. Astraatmadja^{g,2}, J.-J. Aubert^h, R. Auer^o, B. Baretⁱ, S. Basa^j, M. Bazzotti^{kl}, V. Bertin^h, S. Biagi^{kl}, C. Bigongiari^a, M. Bou-Cabo^l, M.C. Bouwhuis^g, A. Brown^h, I. Brunner^{h,3}, I. Busto^h

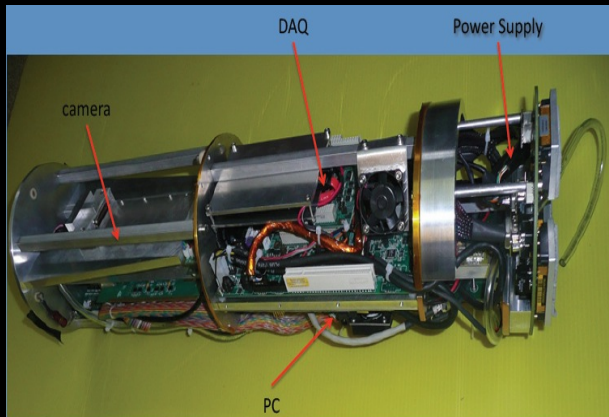


Also World's Deepest Real Time Cameras



2009-06-01 01:46:22

AXIS221 vidéosurveillance
Sensitivity: 0.1 lux
Field of view: up to 90 degrees
Infra red night vision



ebcmos technology
Imaging with
Single photon sensitivity
Very fast frame rate



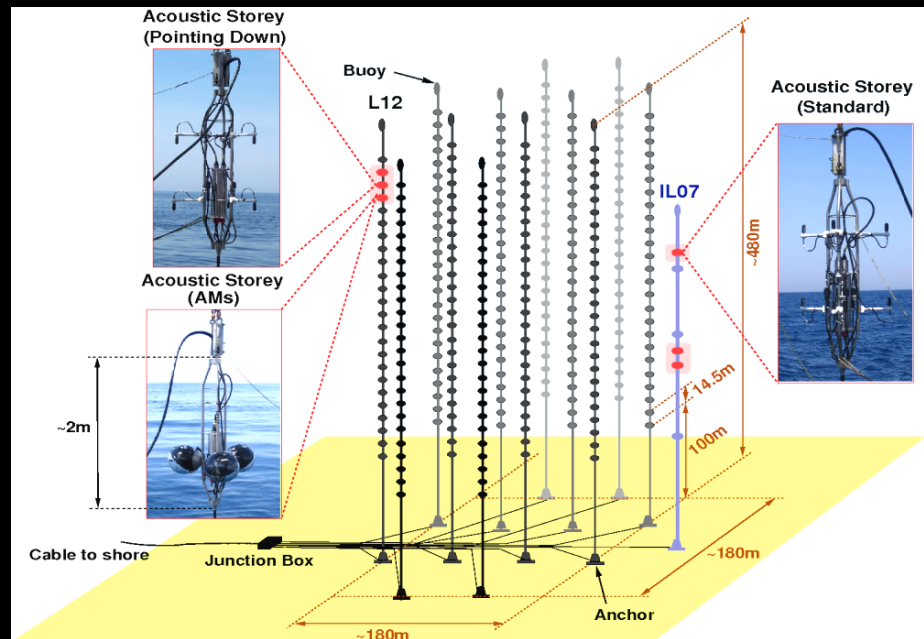
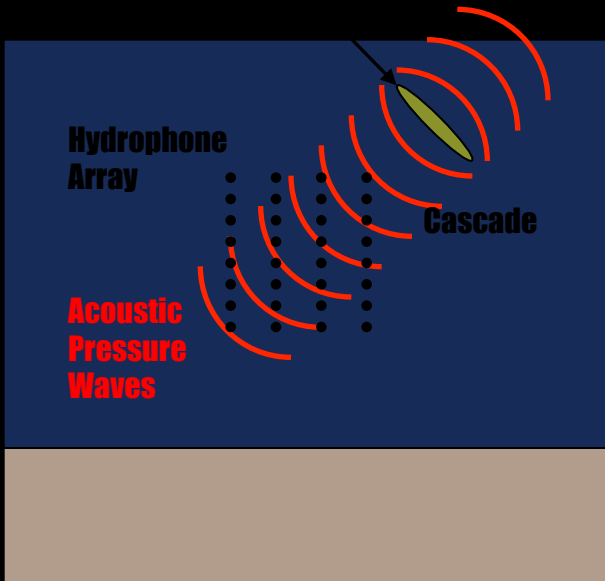
What are these things ?



They even talk to each other

AMADEUS: Acoustic Detection of Neutrinos

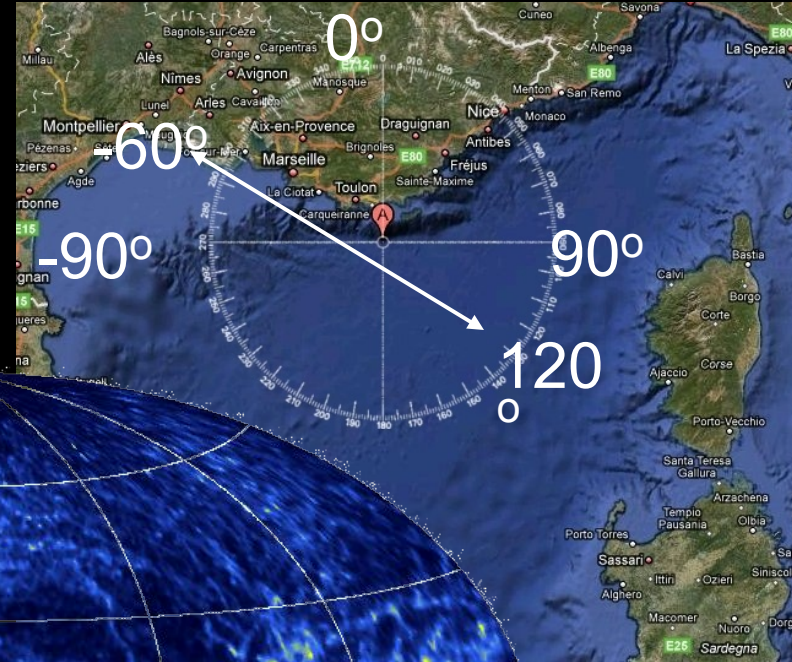
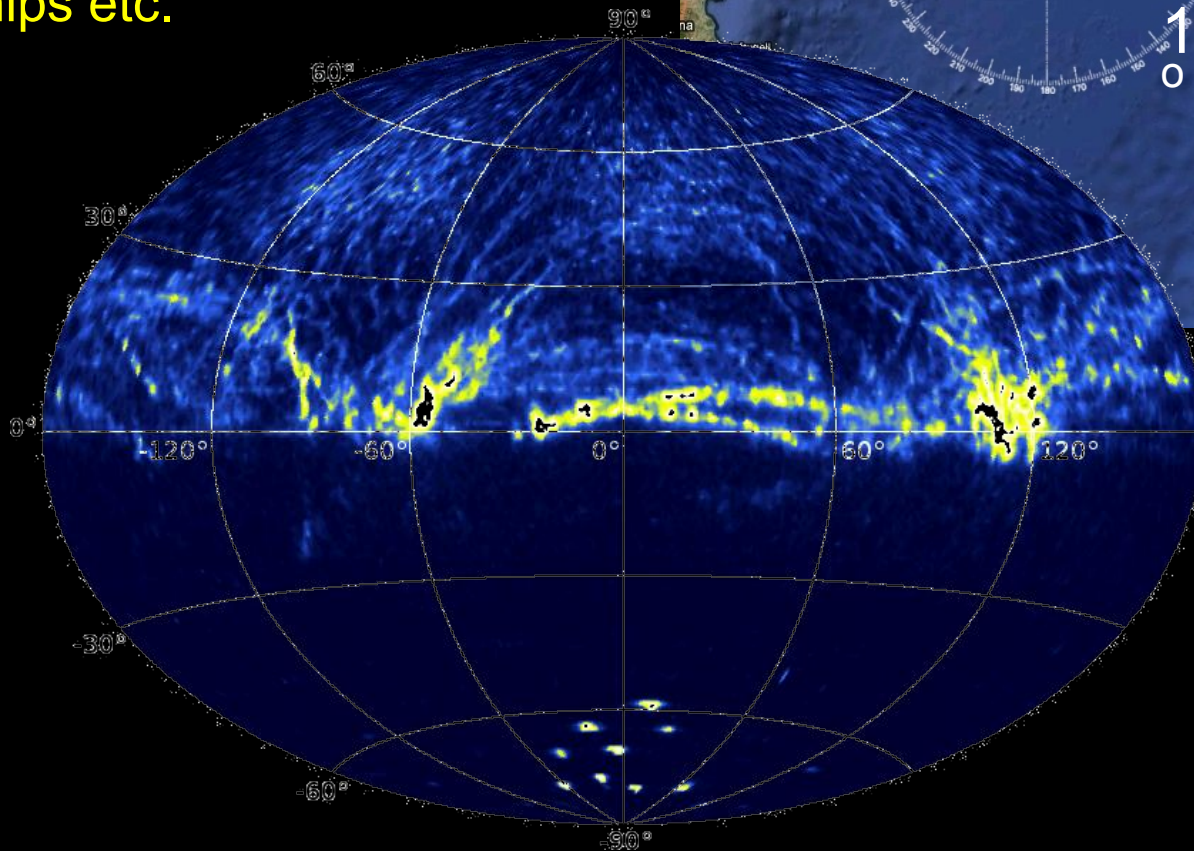
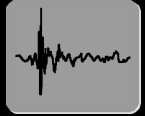
neutrino



Acoustic Source Direction Distribution

Direction reconstruction from one storey on L12

All types of transient signals included, sea mammals, ships etc.





<http://www.listentothedeep.com>



Lido
LISTENING TO THE DEEP-OCEAN ENVIRONMENT



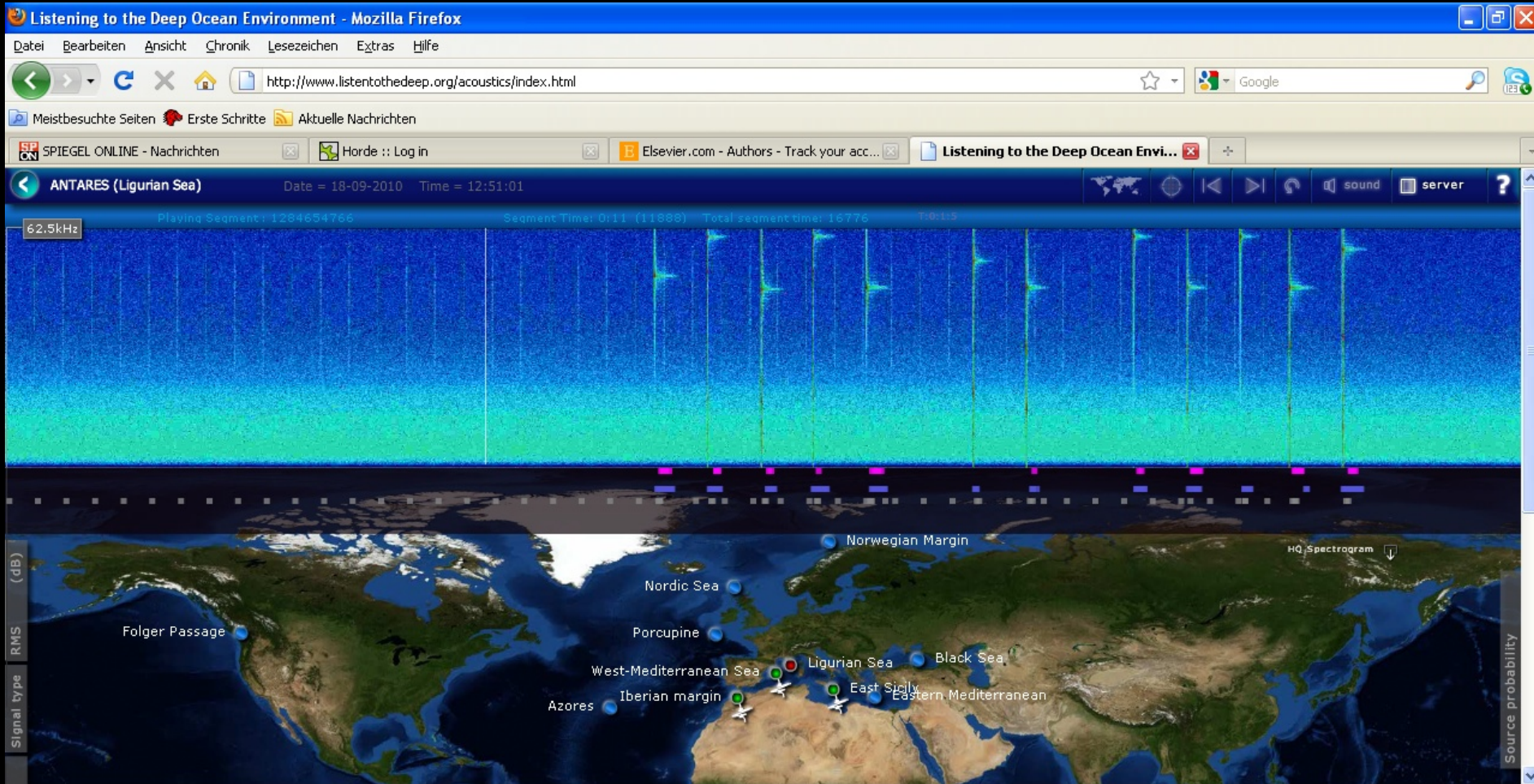
LABORATORI D'APLICACIONS BIOACUSTIQUES
Universitat Politècnica de Catalunya



LISTEN TO ANTARES LIVE

<http://www.listentothedeep.org/>

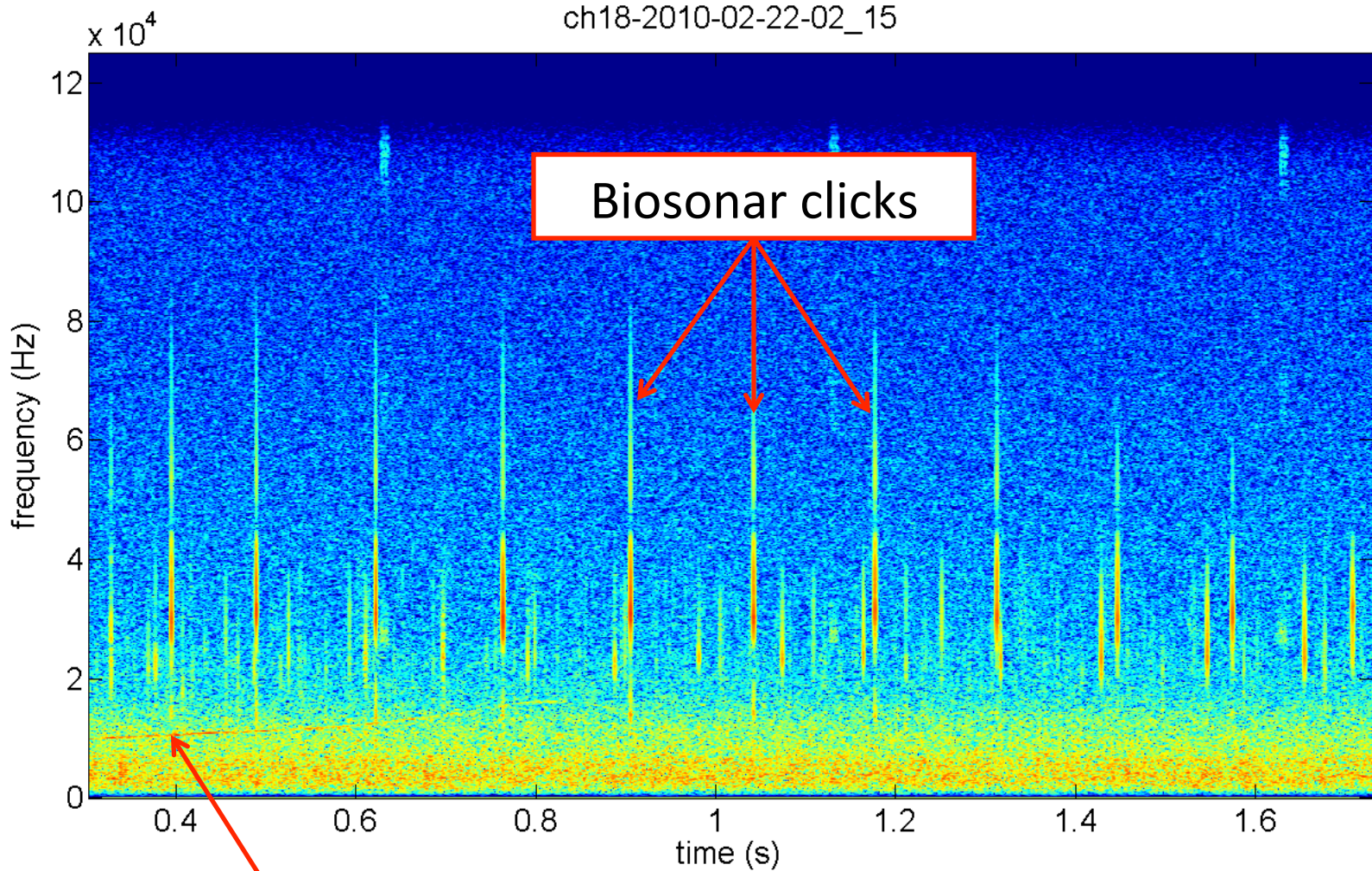
(click on “Enter the bioacoustics page”, then “Ligurian Sea”)



LABORATORI D'APLICACIONS BIOACUSTIQUES
Universitat Politècnica de Catalunya



Dense series of ultrasonic clicks, (biosonar) from beaked whales or dolphins

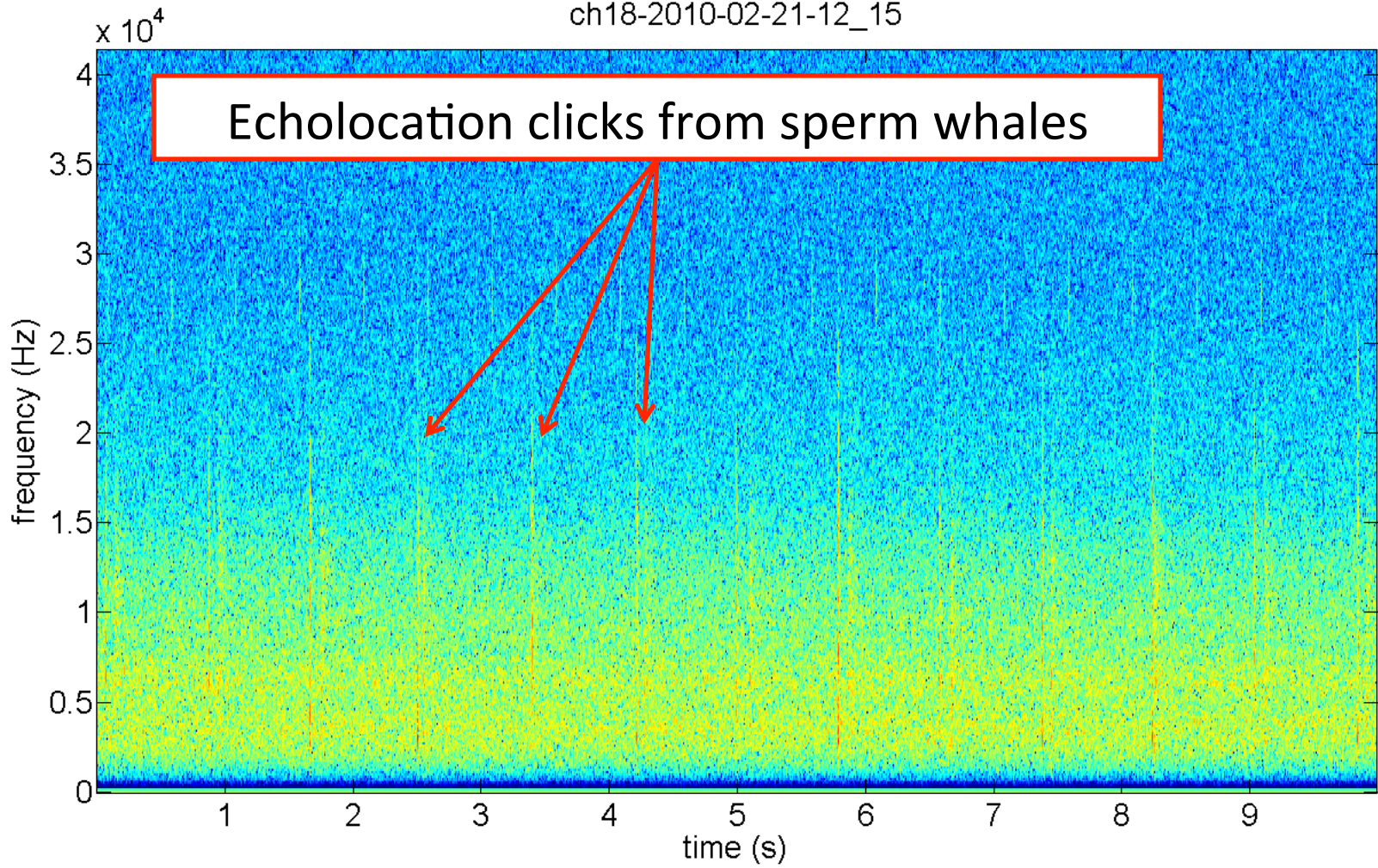


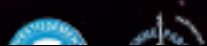
Biosonar clicks

Dolphin whistle



ch18-2010-02-21-12_15





LISTENING TO THE DEEP OCEAN ENVIRONMENT



Presentation

Partners

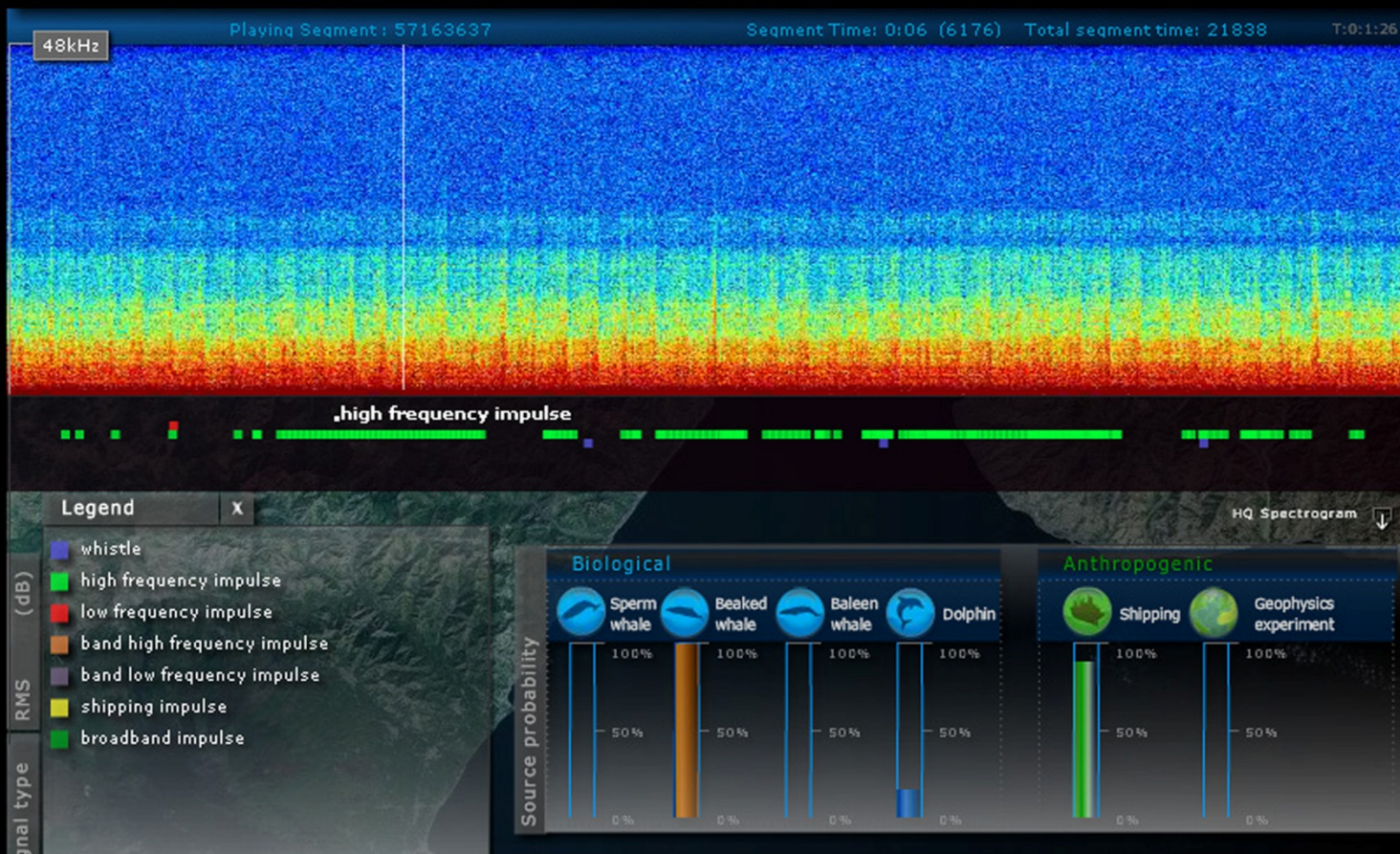
Bioacoustics

Listen on Site

Sound Library

Statistical Analysis

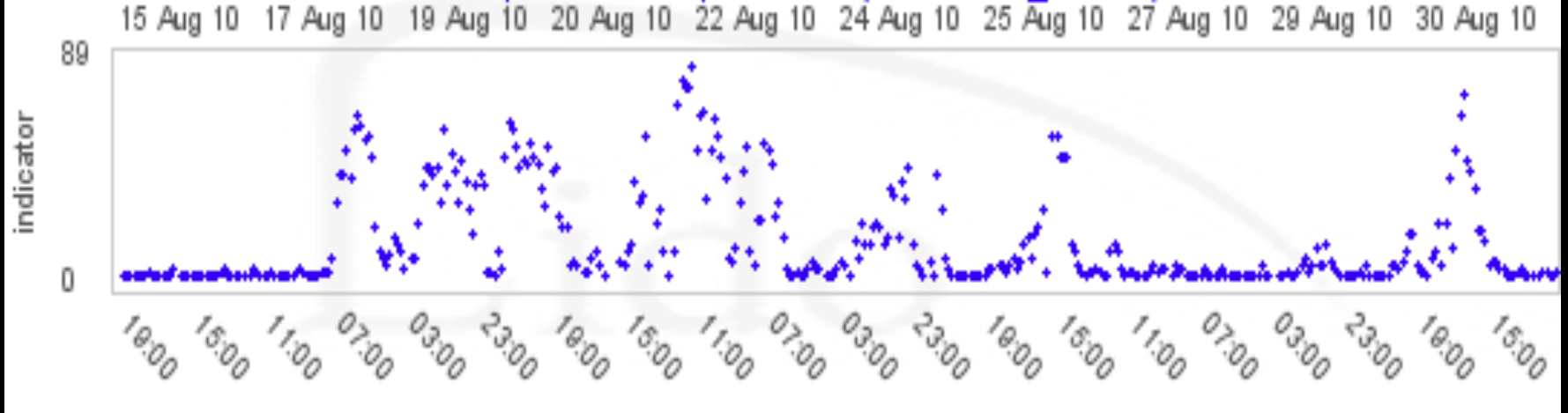
Help



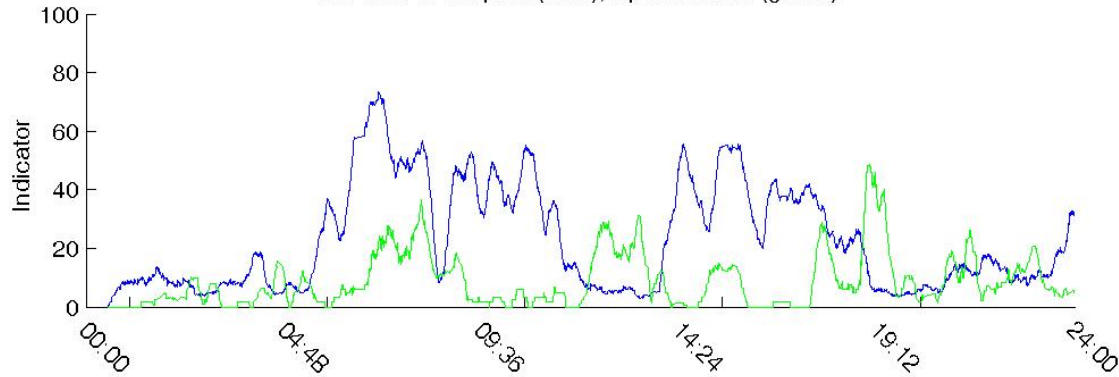


Cetaecean Detections with ANTARES

Sperm whale presence (ANTARES_HOUR)



2010/07/05 Dolphin (blue); Sperm whale (green)



Next step
trajectory
tracking



Up the dolphins and whales

<http://www.listentothedeep.org/>

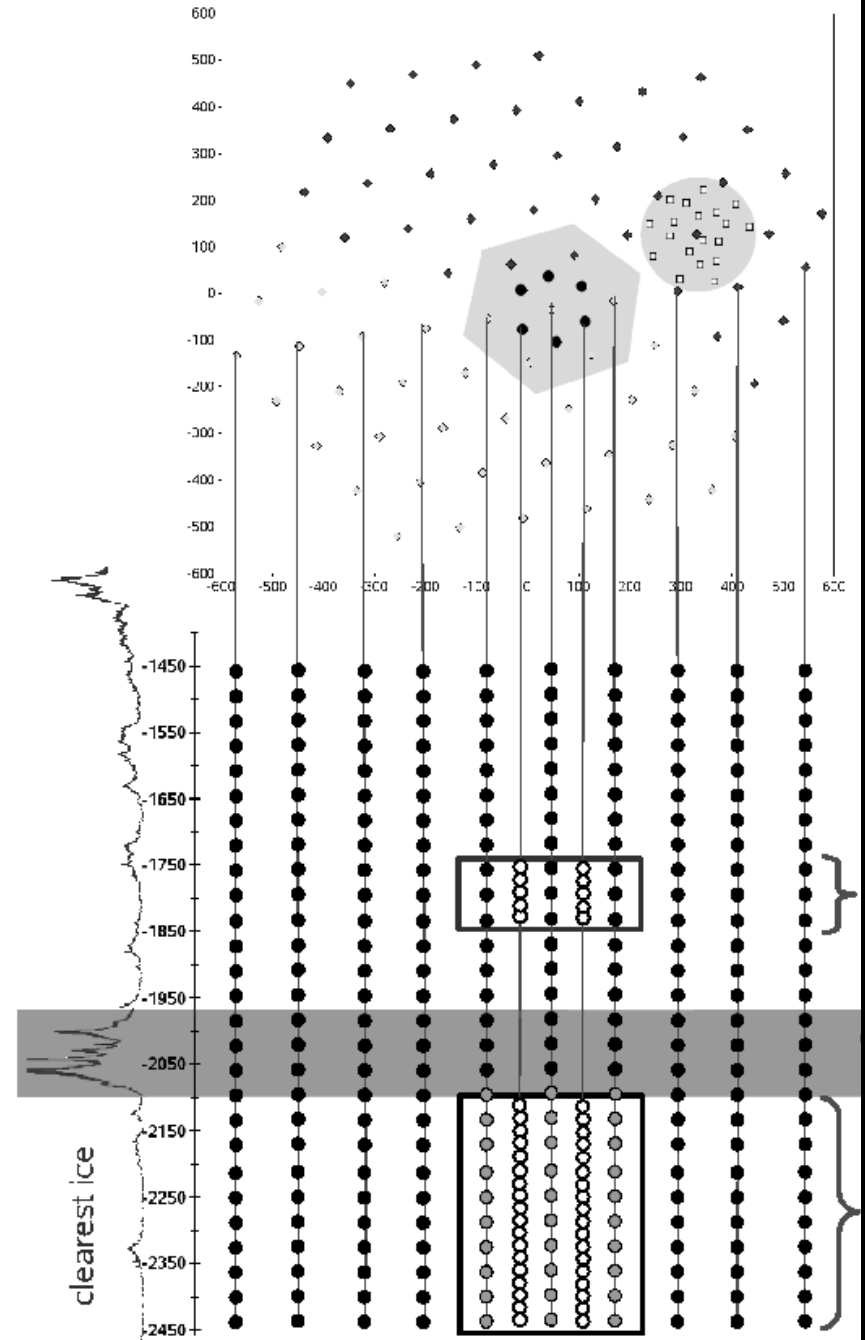
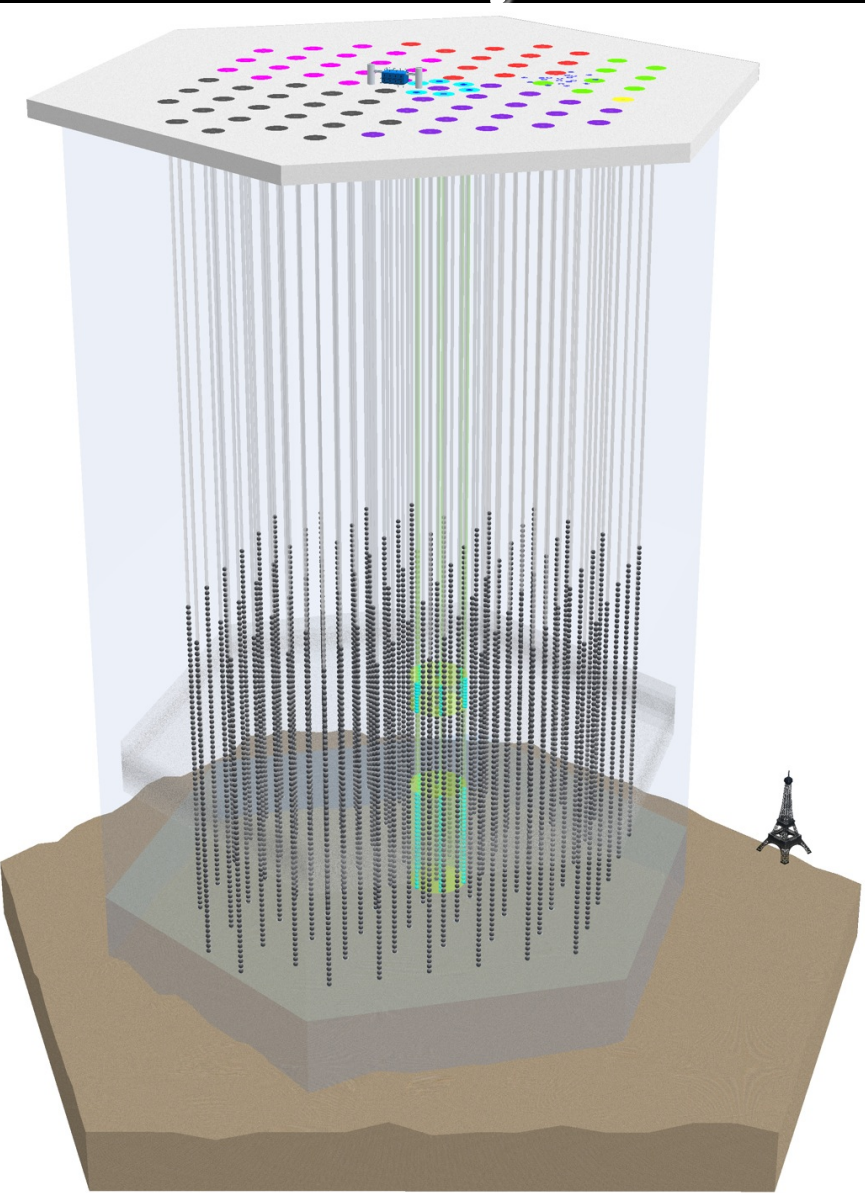
(click on “Enter the bioacoustics page”, then “Ligurian Sea”)

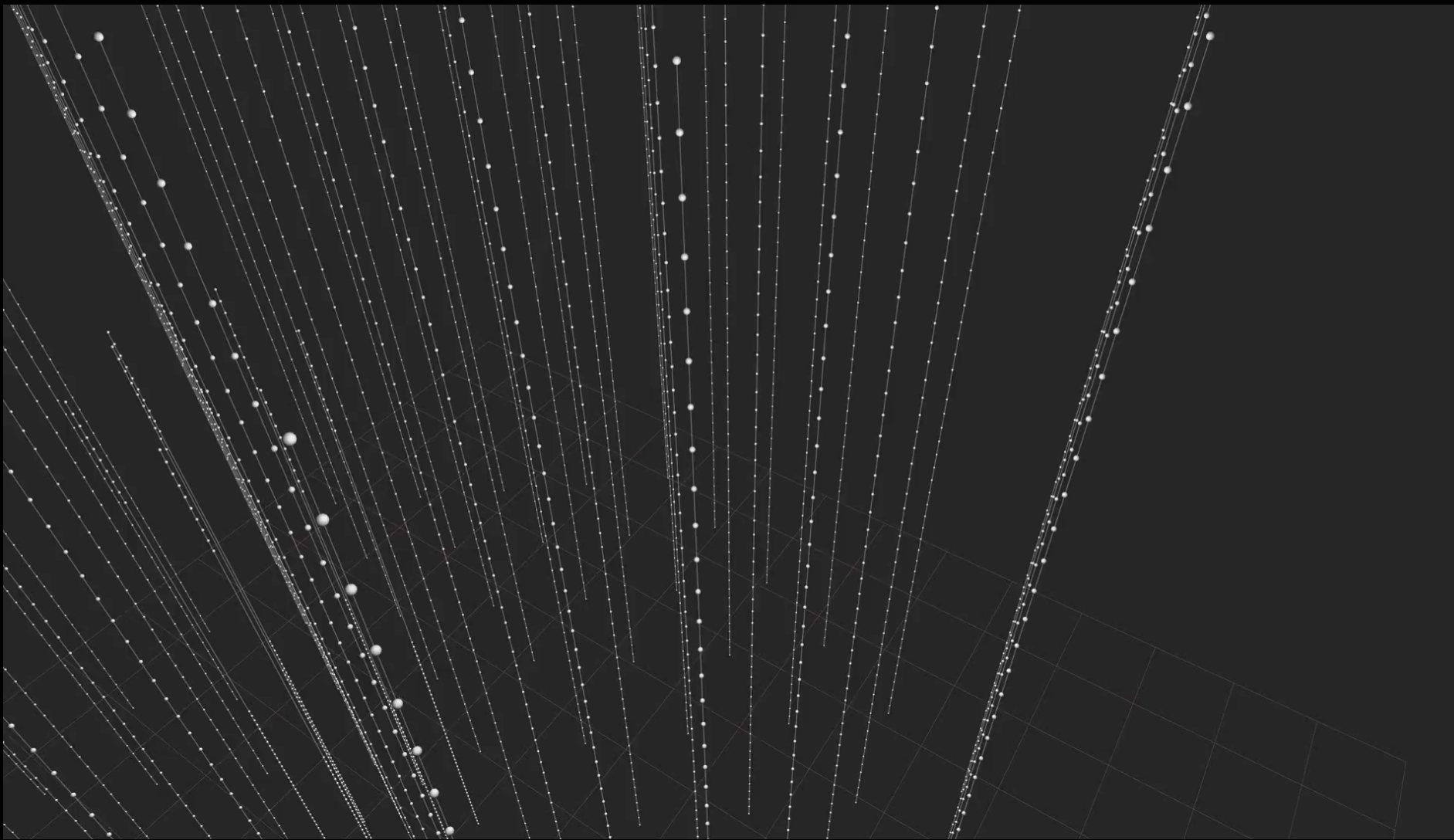


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Universitat Politècnica de Catalunya



The IceCube “Dust Layer”





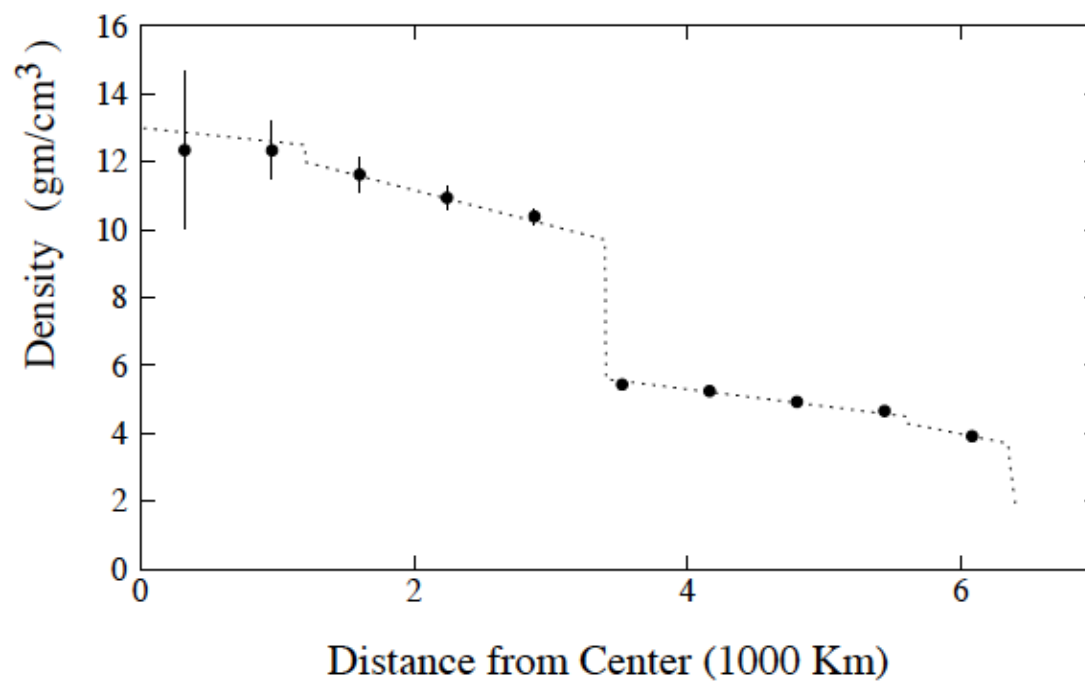
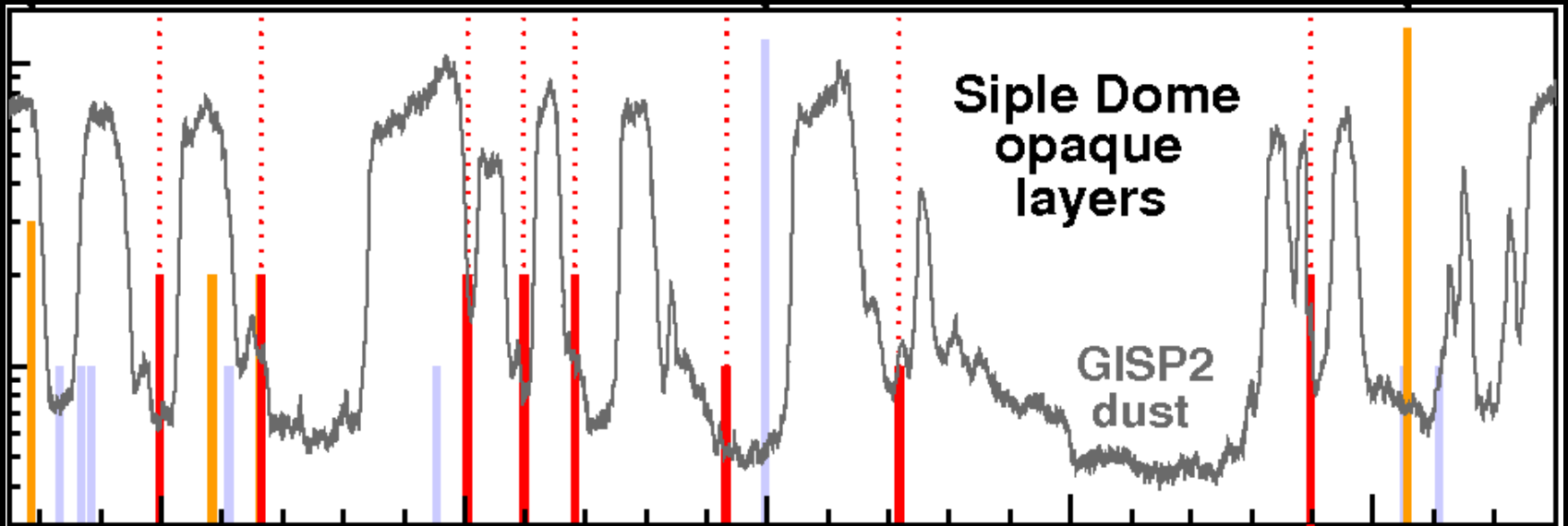
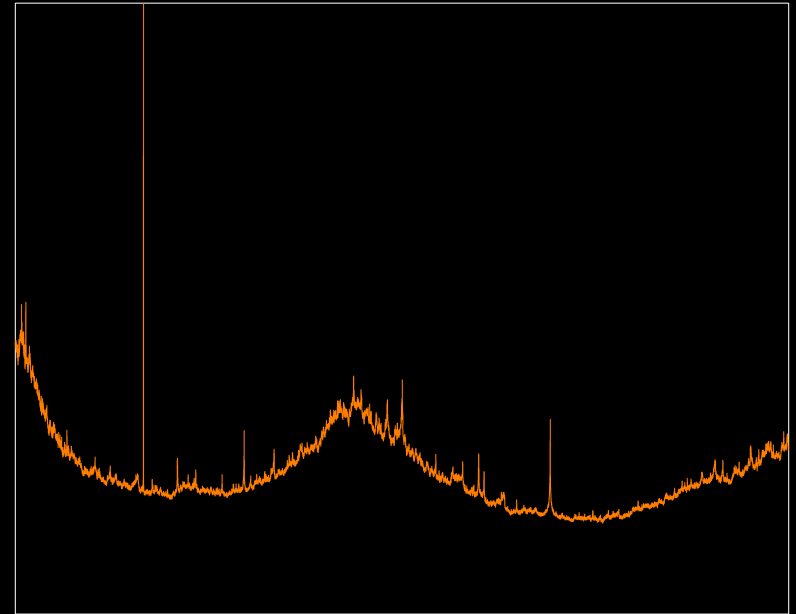
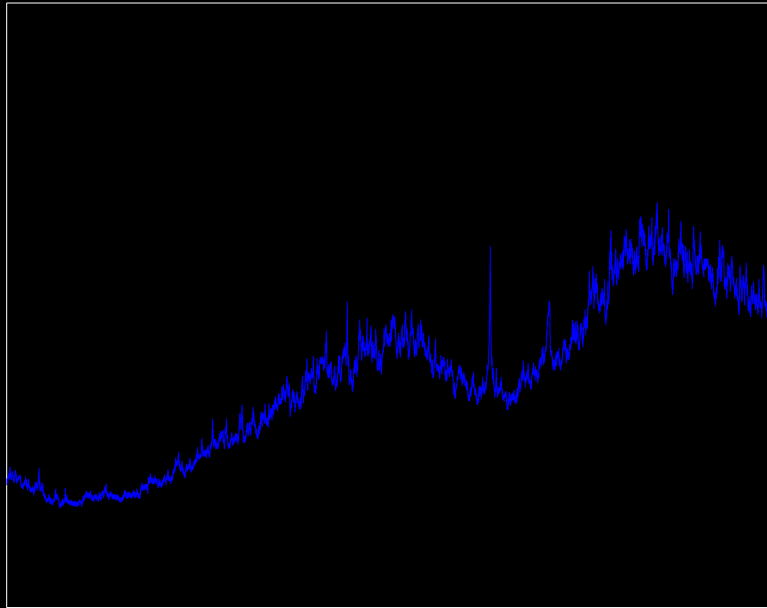
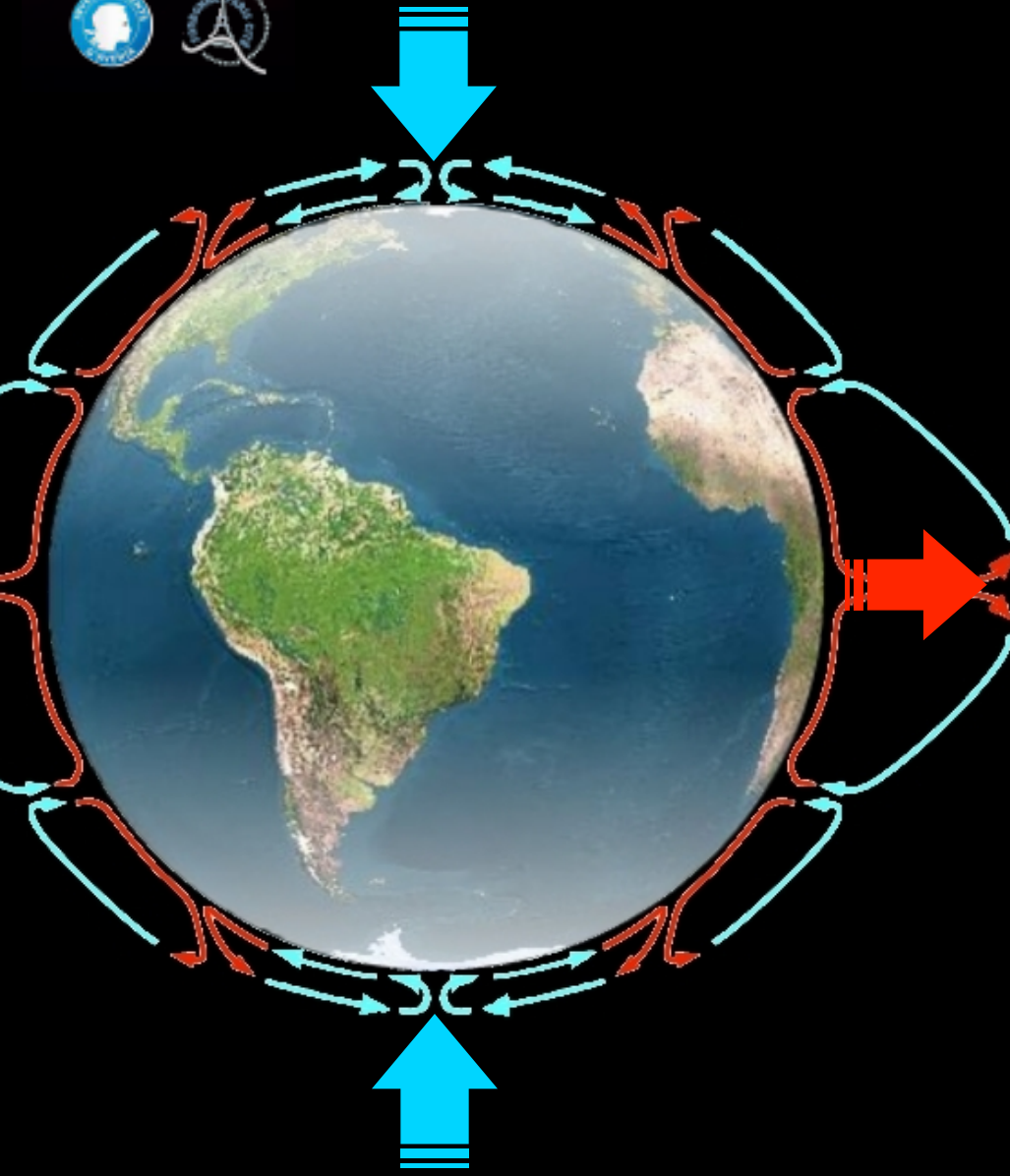


Figure 1: The extracted nucleon density using ten radial bins along with the statistical errors. The PREM model for the Earth's density, dashed curve, was used to generate data. The step between the core and lower mantle is very well resolved, while the difference between the inner and outer cores is not.

Fallout layers & millennial climate change





Dustier ice
during cold climate

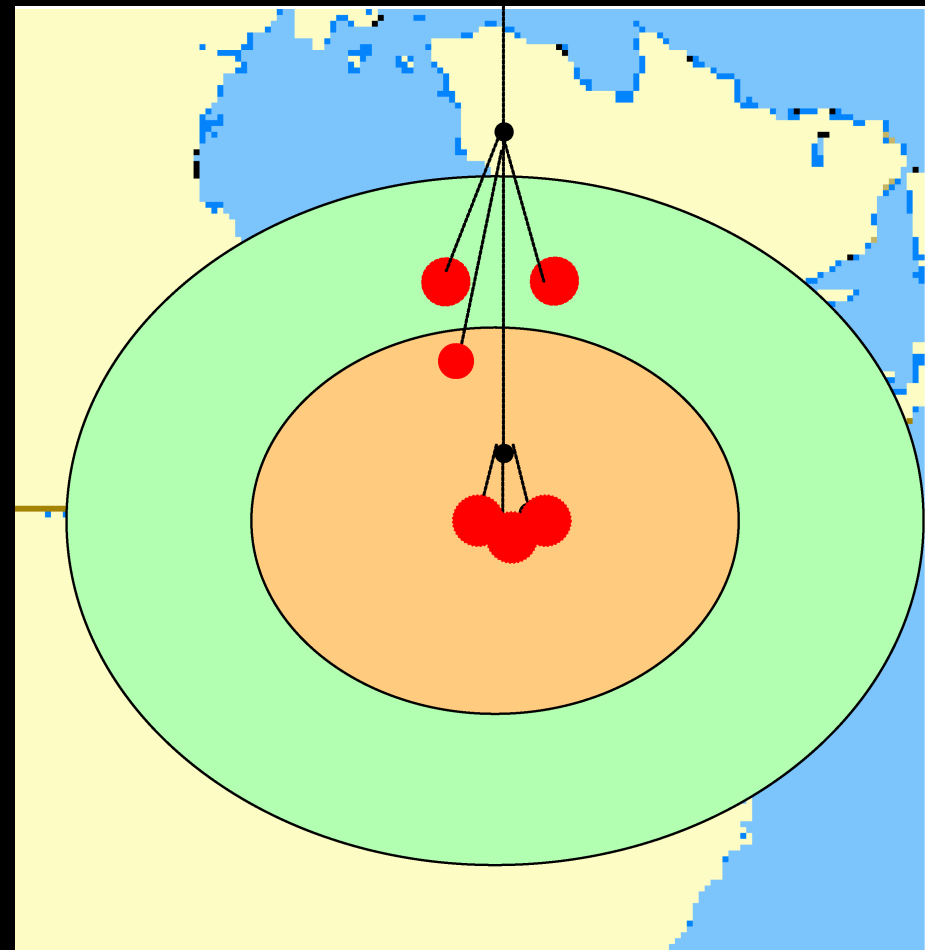
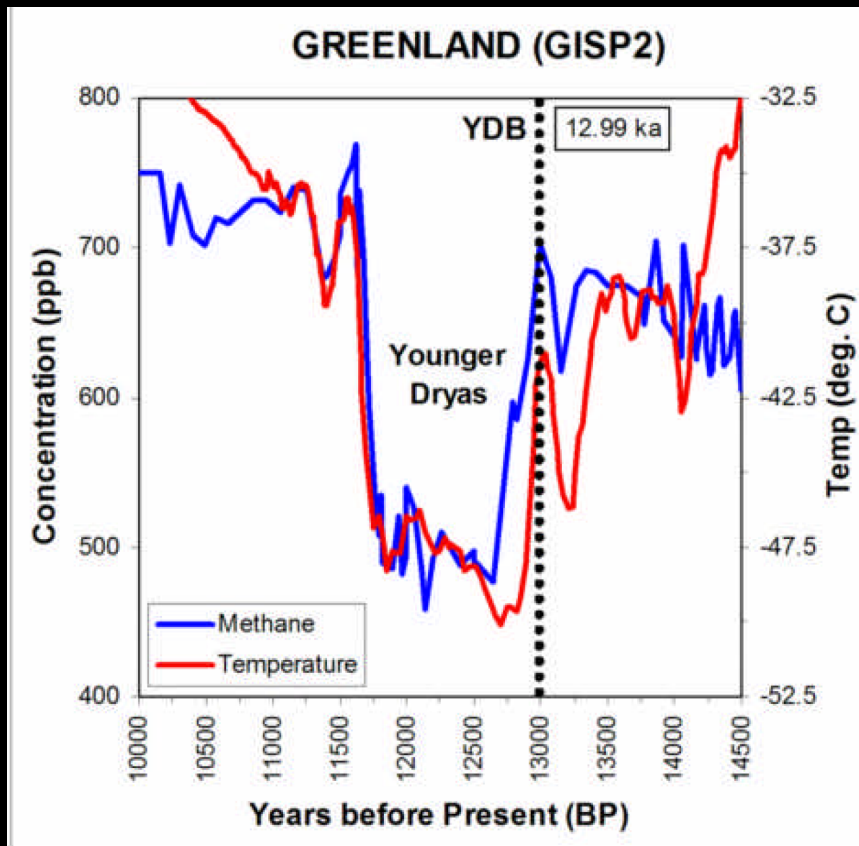
Greater dust
supply

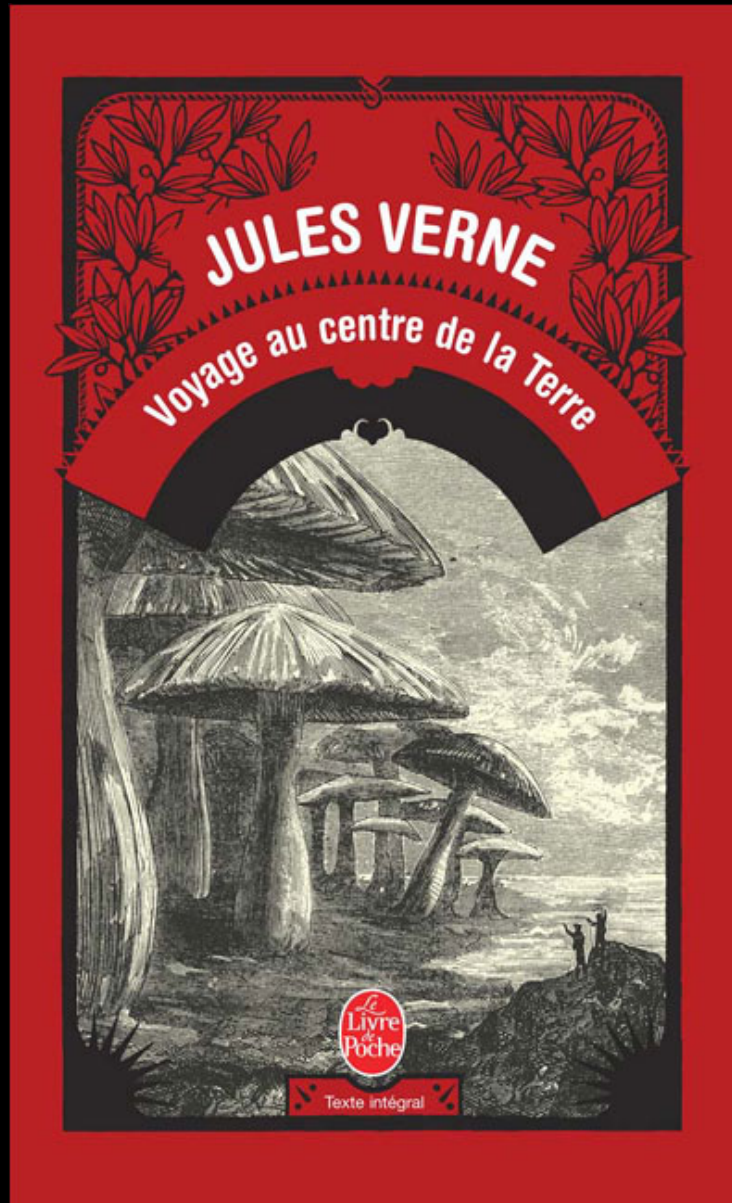
+

More vigorous
circulation, storms,
wind

Evidence for an extraterrestrial impact 12,900 years ago that contributed to the megafaunal extinctions and the Younger Dryas cooling

R. B. Firestone^{a,b}, A. West^c, J. P. Kennett^d, L. Becker^e, T. E. Bunch^f, Z. S. Revay^g, P. H. Schultz^h, T. Belgia^g, D. J. Kennettⁱ, J. M. Erlandson^j, O. J. Dickenson^j, A. C. Goodyear^k, R. S. Harris^h, G. A. Howard^l, J. B. Kloosterman^m, P. Lechlerⁿ, P. A. Mayewski^o, J. Montgomery^j, R. Poreda^p, T. Darrah^p, S. S. Que Hee^q, A. R. Smith^a, A. Stich^r, W. Topping^s, J. H. Wittke^f, and W. S. Wolbach^r





Underground synergies

Dark Matter Music





Laboratoire Souterrain de Modane

Depth: **4800 m.w.e.**

Surface: **400 m²**

Volume : **3500 m³**

Muon flux: **$4 \cdot 10^{-5} \mu.m^{-2}.s^{-1}$**

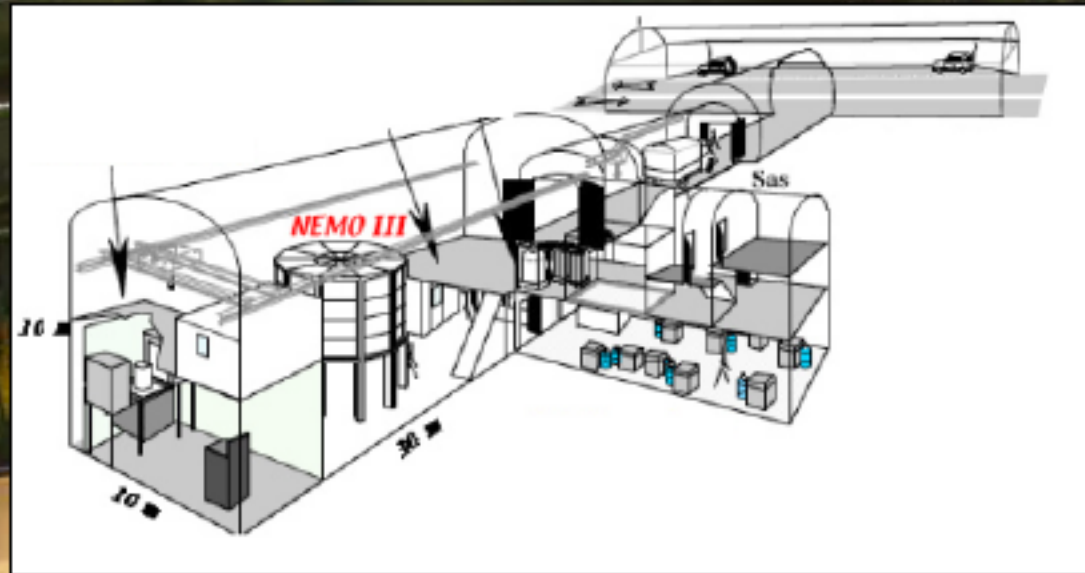
Neutrons:

Fast flux: **$4 \cdot 10^{-2} n.m^{-2}.s^{-1}$**

Thermal flux: **$1.6 \cdot 10^{-2} n.m^{-2}.s^{-1}$**

Radon: **15 Bq/m³**

Access : **horizontal**



Budget (full cost): **1 M€/yr**

Staff: **3 Physicists**

3 Engineers

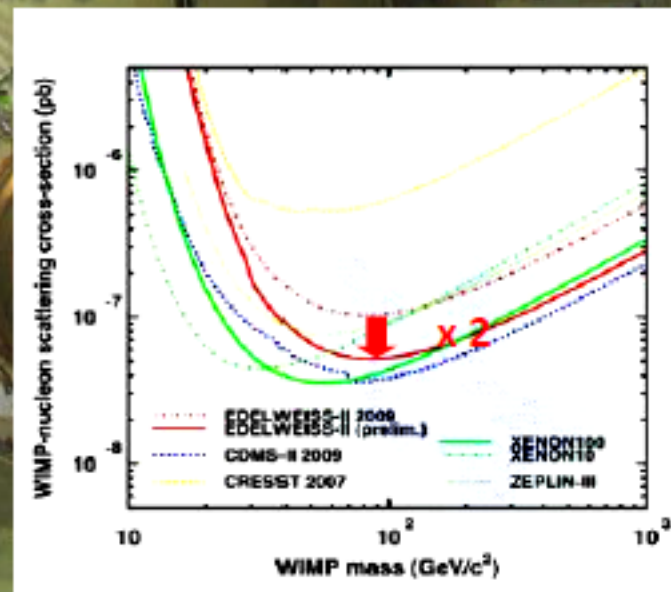
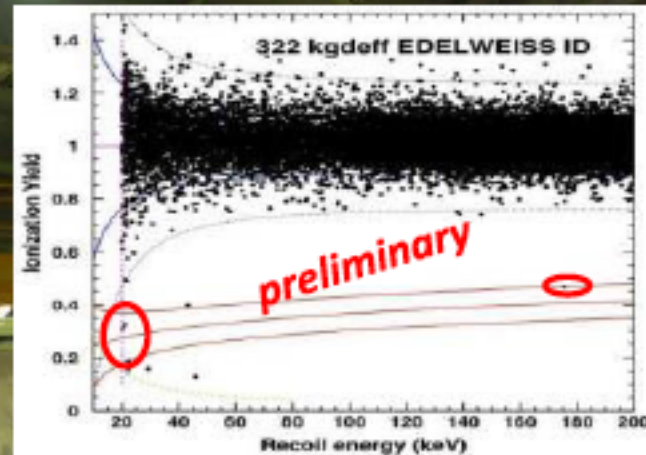
7 Technicians

International associated laboratory agreement with JINR Dubna (Russia) and CTU Prague (Czech Republic)



Dark matter: EDELWEISS

Bolometric technique: Heat + Ionization Ge crystals - 40 kg





Radon free air factory

Copy of the facility developed by SuperKamiokande

150 m³/h of air with an activity of 20 mBq/m³ (air in the lab 20 Bq/m³)



Build in Czech Republic



Gamma ray spectroscopy @ LSM

14 HPGe from 7 different laboratories of CNRS , CEA, JINR DUBNA and CTU Prague are available at LSM



- Material selection for astroparticle physics,
- Environnemental research (oceanography, climat, retro-observation,....)
- Environmental survey
- Applications (wine datation, salt origin,...)
- Developements of Ge detector (ILIAS)





Interdisciplinary @ LSM

Use of the ultra-low gamma-ray spectroscopy

Radio-isotopes are used as tracers in the environment or as chronometers for dating of glacial or sedimentary layers.

They are used also for archaeological objects which sometimes require non-destructive measurements

Some exemples:

- Environmental survey
- Characterization the age of the suspended solids and pollutants associated with them in rivers
- Marine and continental geochemistry
- Characterization of water masses, their origin and age in the ocean
- Retro-observation (effects on human activities on the environment)
- Radioactivity in the atmosphere



Interdisciplinary @ LSM

Use of the gamma-ray spectroscopy

Applications :

- Characterisation of water (lakes, rivers, underground water) EU directive
- For drug and food administration ex. wine dating, marine salt origin
- Judicial expertises
- Mean age of crustacean livestock for fishing regulation
-

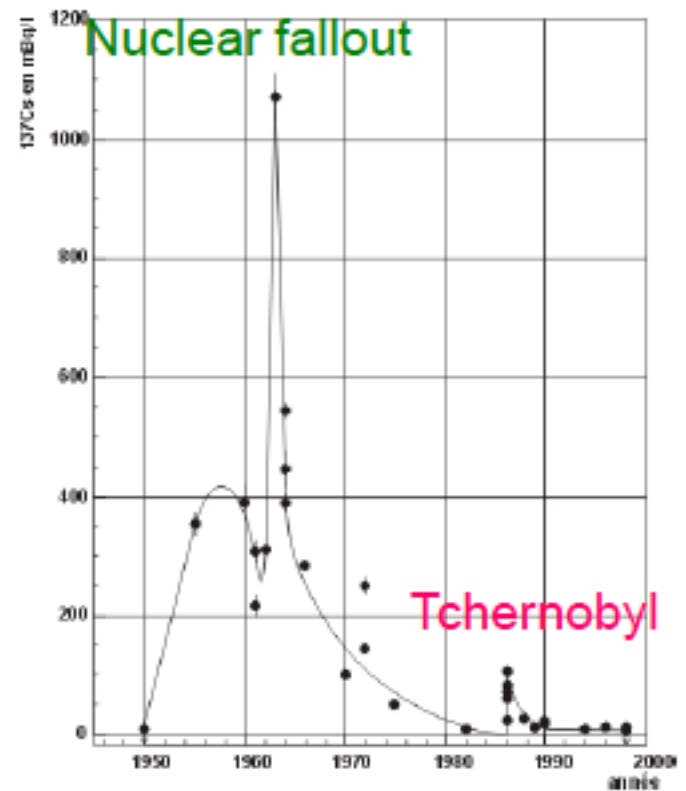
Developemnt of a national ultra-low radioactivity platform measurements with

EDYTEM (University of Savoie/CNRS), LGGE (University of Grenoble / CNRS), LSCE (CNRS/CEA), LPSC (University of Grenoble / CNRS) and LSM

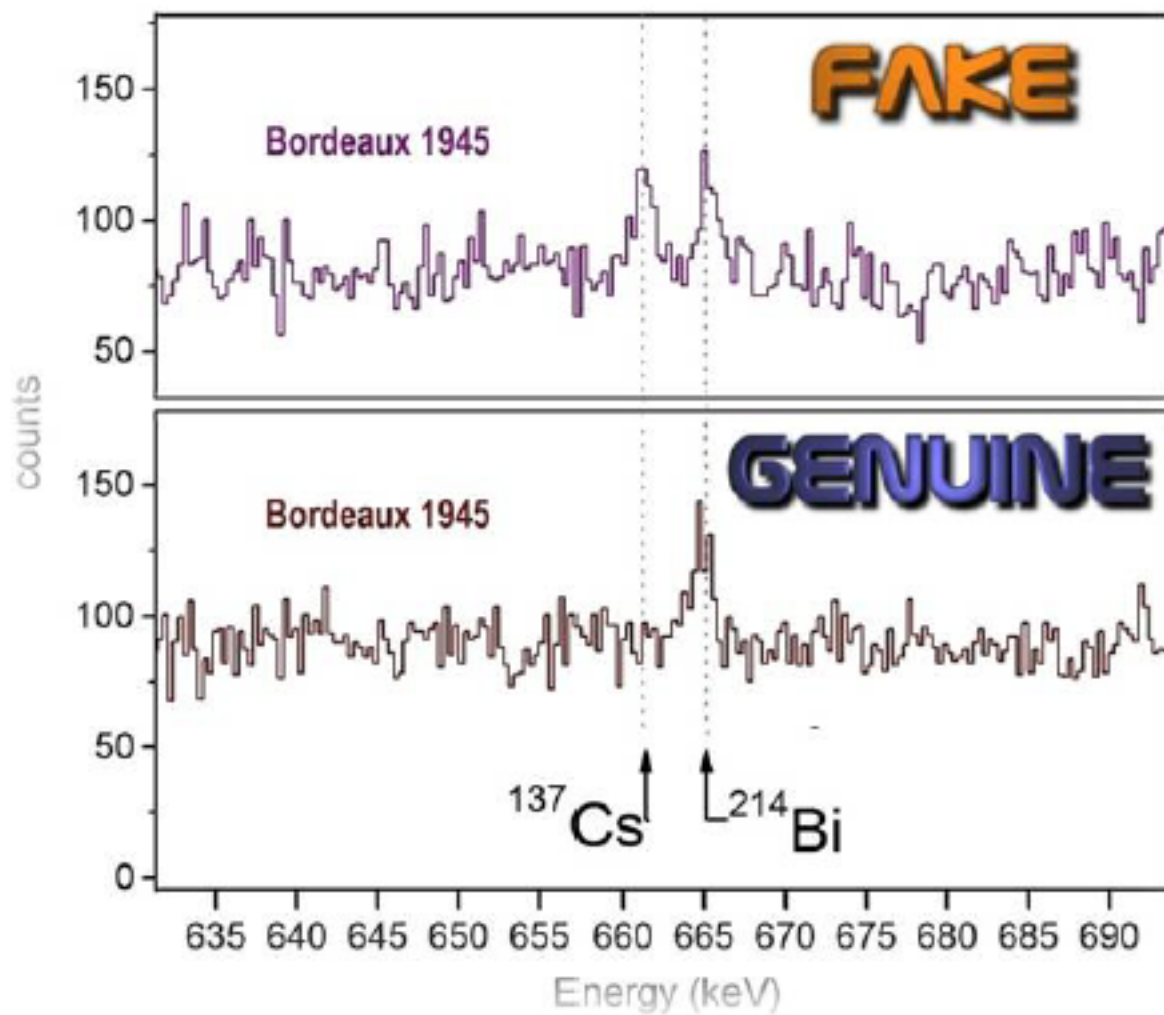
Wine datation by ^{137}Cs measurement

Developed by Ph. Hubert (Centre d'études Nucleaires de Bordeaux-Gradignan)

Châteaux
"Lafite"
&
"Margaux"
1900 ?



Comparison of 2 magnums
same château and same vintage?!?



DUSEL an ambitious program

<http://www.sanfordlab.org/publications/bge-scienceusel>]

Dark Life (Biology)

How deep does life go?

Do biology and geology interact to shape the world underground?

How does subsurface microbial life evolve in isolation?

Did life on earth originate beneath the surface?

Is there life on earth as we don't know it?

Restless Earth (Geosciences)

What are the interactions among subsurface processes?

Can we view complex underground processes in action?

Can we forewarn of earthquakes?

Ground Truth (Geoengineering)

What lies between boreholes?

How can technology lead to a safer underground?

How do we better harness deep underground resources?





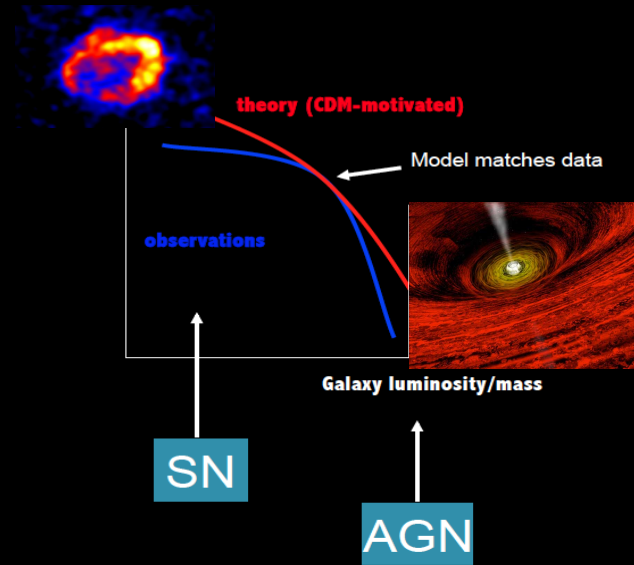
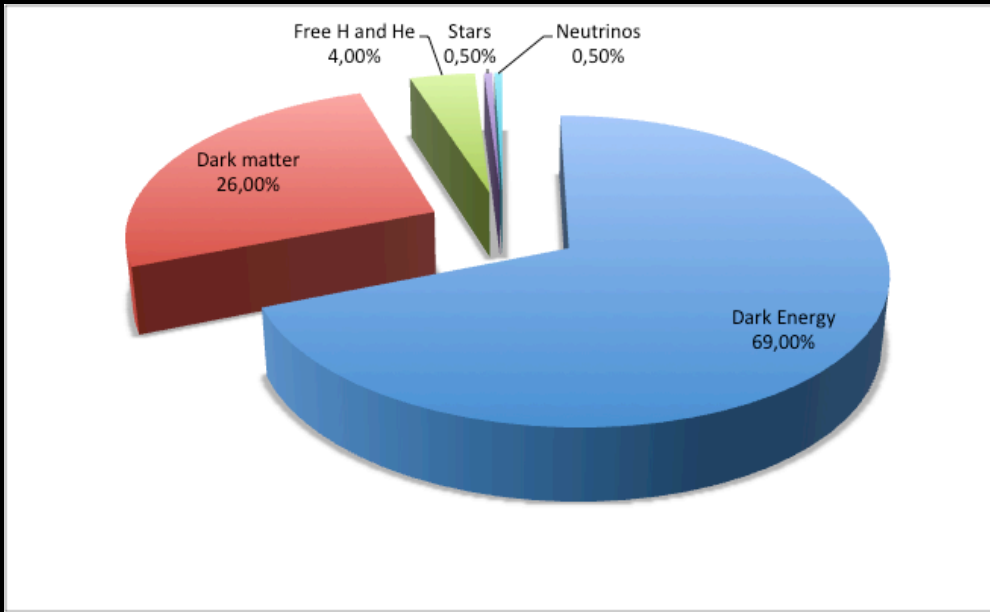
Conclusions

Synergy Geosciences-astroparticle: a new frontier:

- For astroparticle the Geosphere is both the target and the detecting medium: continuous time series data by deploying large (autonomous?) sensor networks in hostile environments (sea, desert, underground)
- Dating is important: radioactivity-free platforms
- High sensitivity instruments for probing very low intensity geological effects, metrology
- Large data manipulation and worldwide networking.
- Distribution of alerts
- A COMMON AMBITION: COVER THE EARTH WITH SENSORS

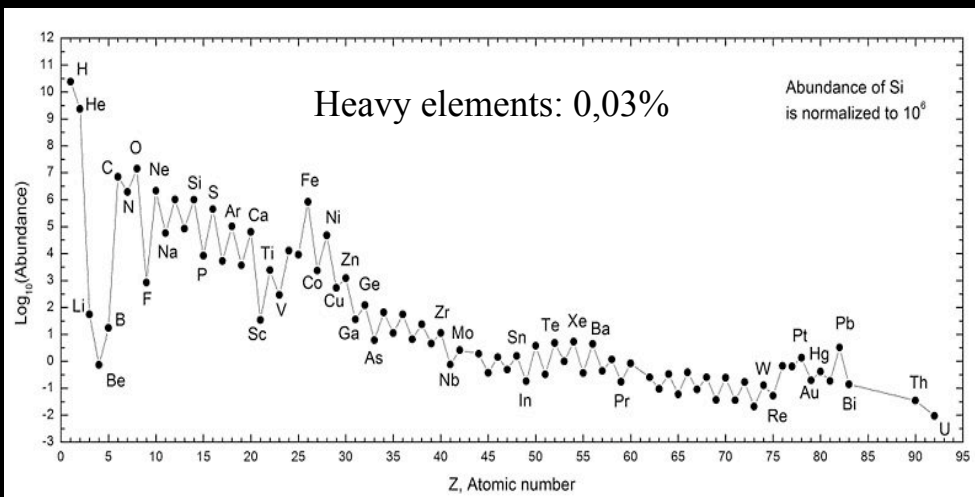


Cosmos in a nutshell composition and role of high energy phenomena



The existence of dark matter not sufficient to explain the formation of cosmic structures, one needs regulation (feedback) from violent phenomena

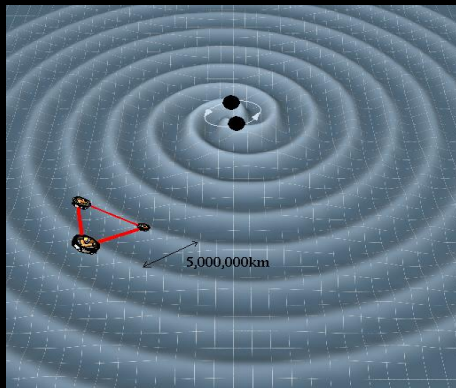
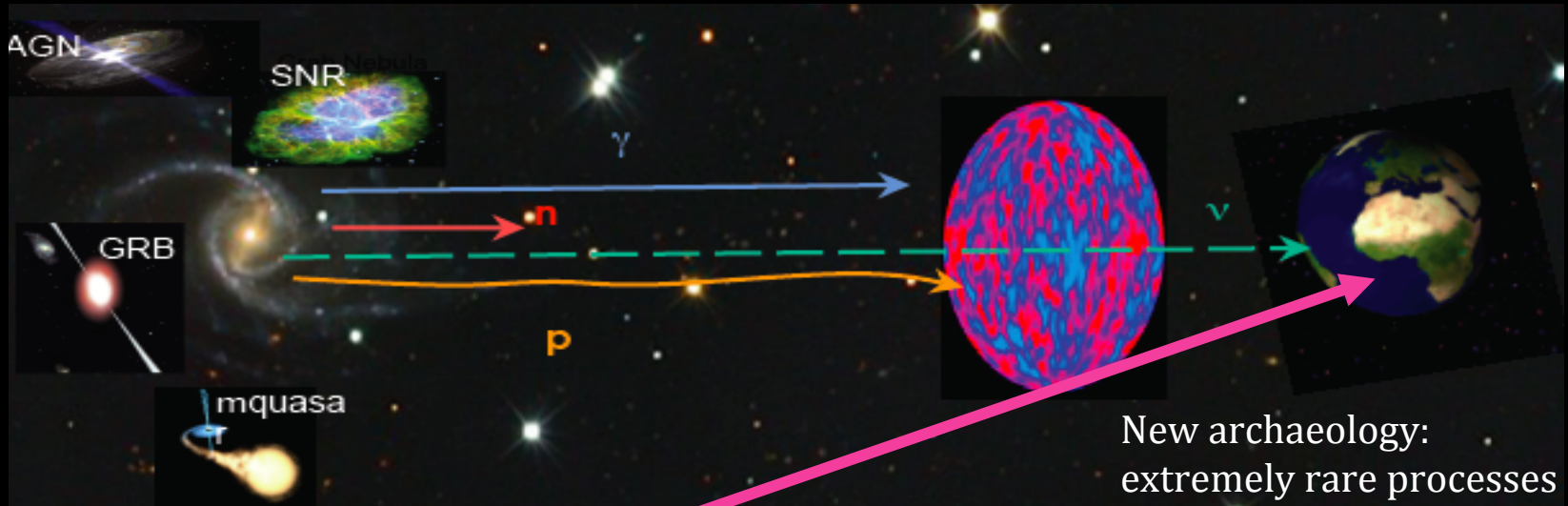
- The elements H, He (+ a little Li) mostly primordial nucleosynthesis
- The rest till Fe, Ni star nucleosynthesis
- Above Fe, Ni at the environment of violent phenomena and cosmic ray spallation





New astronomies

High energy photons, neutrinos, cosmic rays and gravitational waves



- Earth changes role: from privileged optical observatory of the cosmos to a generalised detector
- Out of the 2D projection on its surface one deduces the structure of the Universe in 3D.
- Radioactive decays help dating the structure of the earth and the Universe expanding the projection to 4D.
- Holography?



Astroparticle Physics Connecting scales

The Astroparticle theme after LHC/PLANCK/ ν results can be reduced to 2 fundamental questions:

- 1) Are there any intermediate scales between the EW and Inflation? If yes where are they? (Unification, SeeSaw, TeV, eV)
 - 1) Inflation, dark matter and energy
 - 2) Neutrino mass and proton decay
- 2) Are there new energy scales at work in the most violent phenomena of the Universe? How do particles and fields shape the formation and evolution of cosmic structures?
 - 1) High energy photons, neutrinos, CR
 - 2) Gravitational waves



Planck-Scale

Grand Unification

Leptogenesis

Dark Matter

Fermi-Scale

eV-Scale

The heavenly ladder